

Noise from stray light in interferometric gravitational waves detectors

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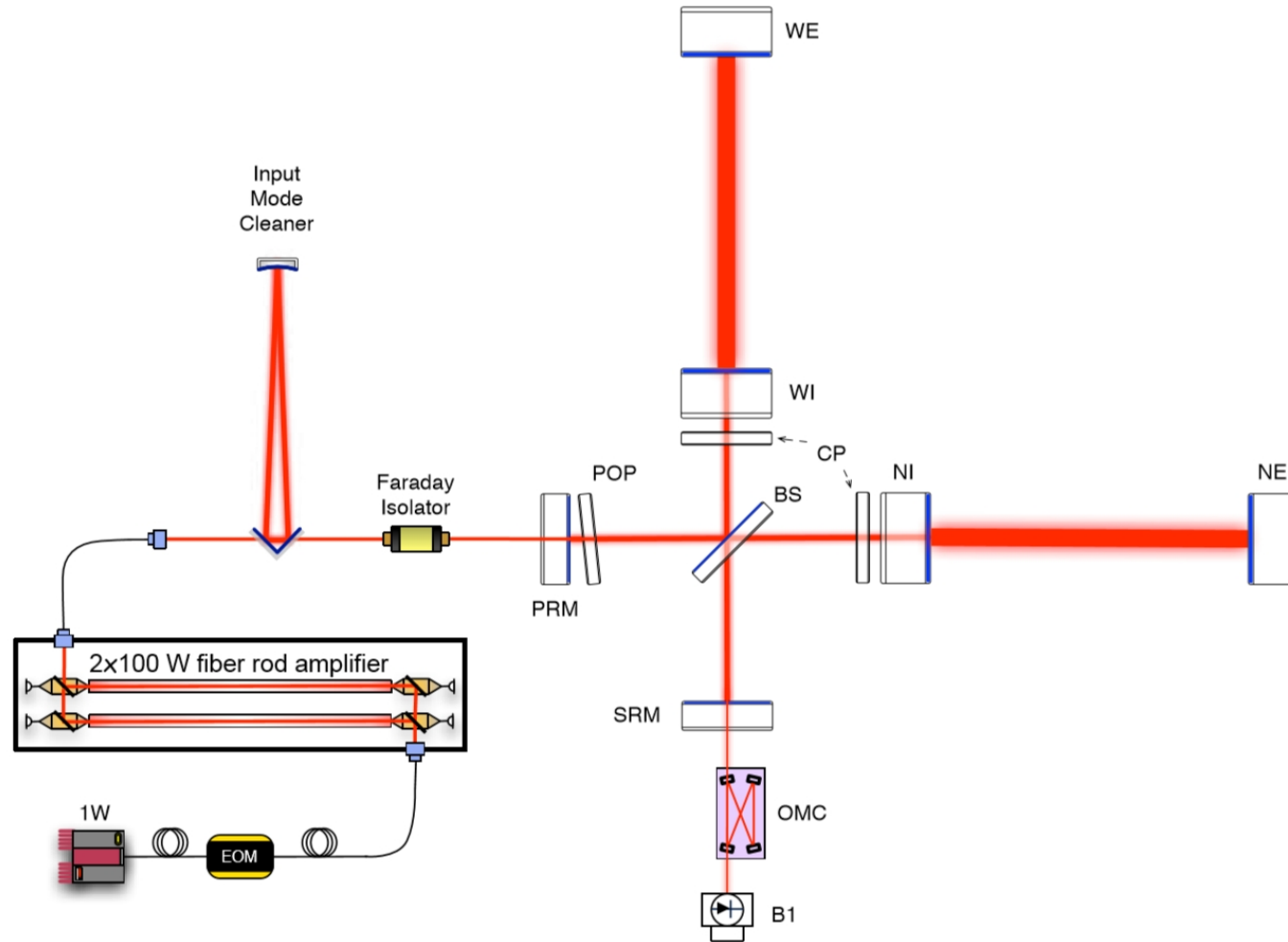


Outline

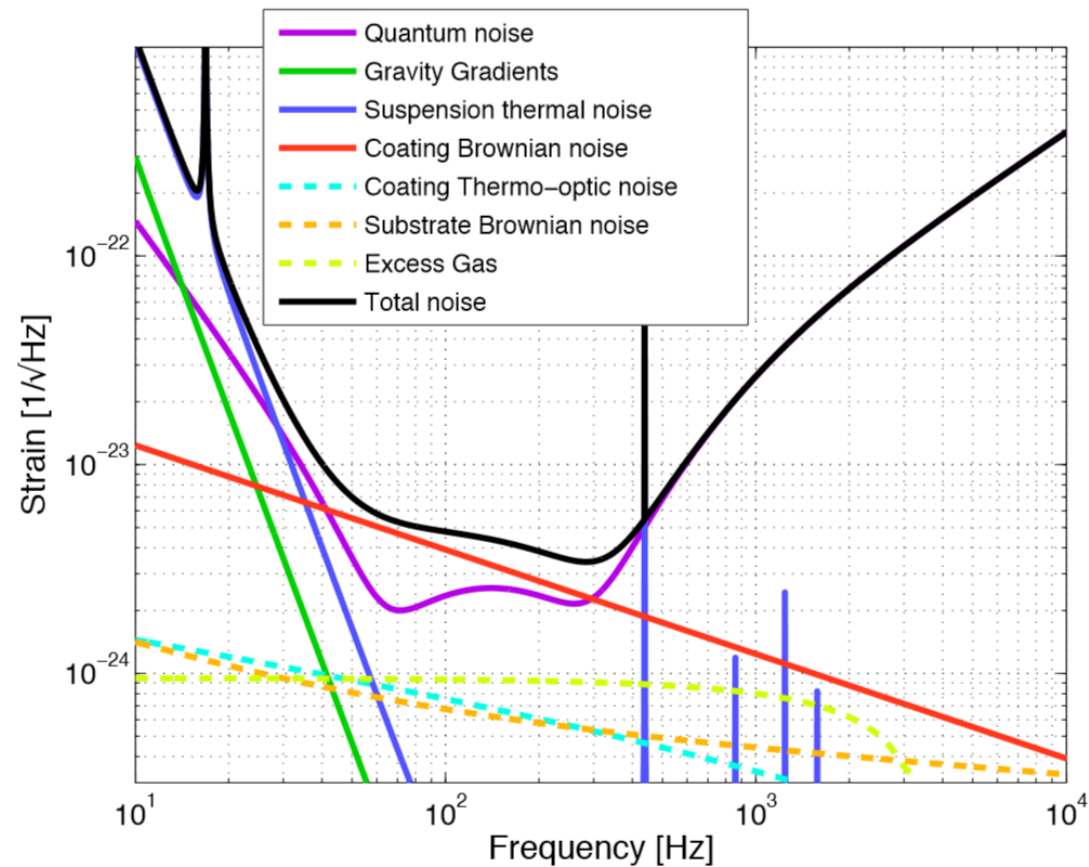
- Interferometric gravitational waves detectors
 - Advanced Virgo
 - Stray light issues
- Ray tracing
 - Ray tracing fundamentals
 - Shading
 - Interference with particle model of light
- Coupling between light and mechanics
 - Phase noise
 - Fabry-Perot cavities



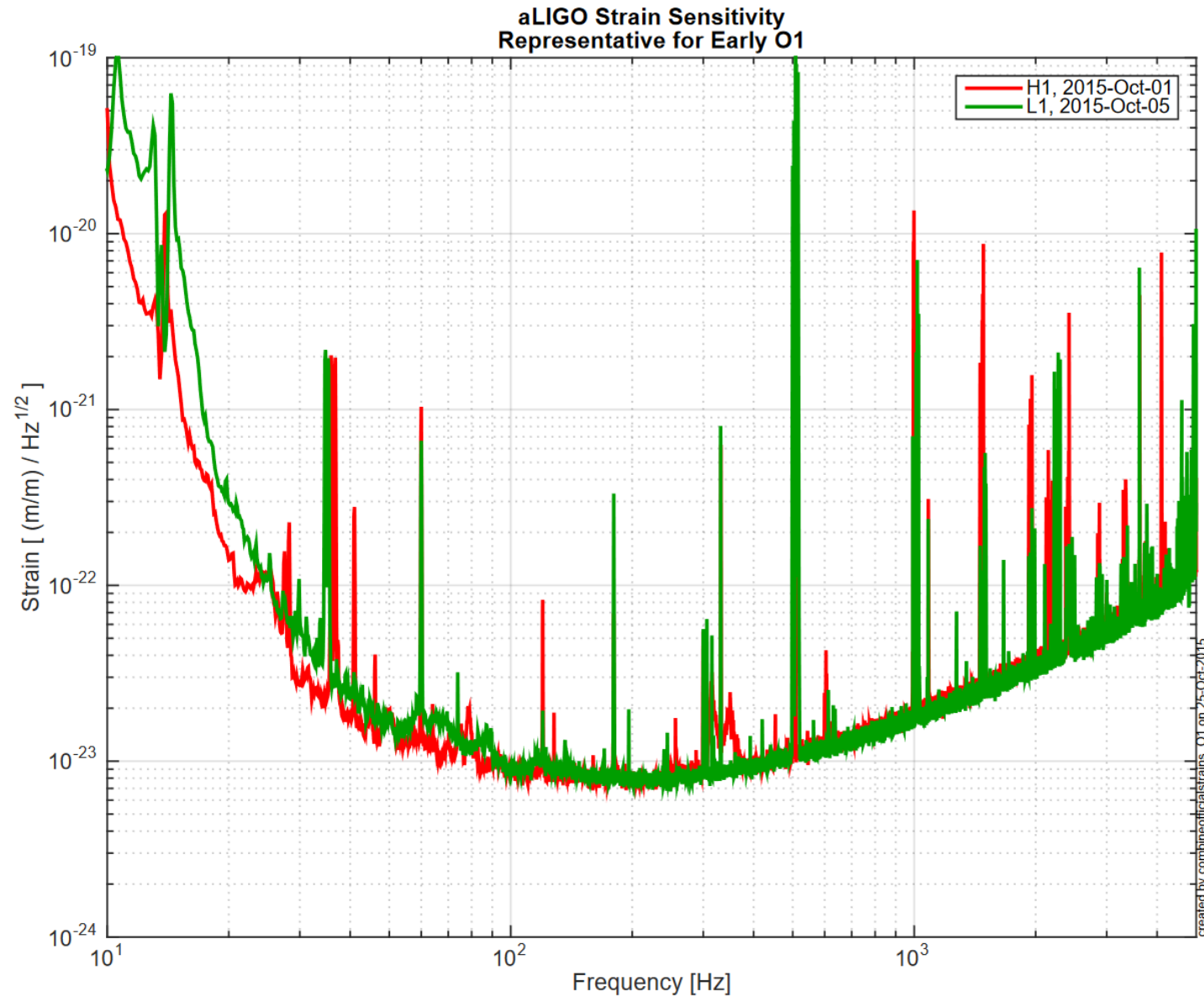
Advanced Virgo configuration



Advanced Virgo design Sensitivity



Ligo Sensitivity: SR1



Stray light

Stray light is all the light that follows a non expected path.

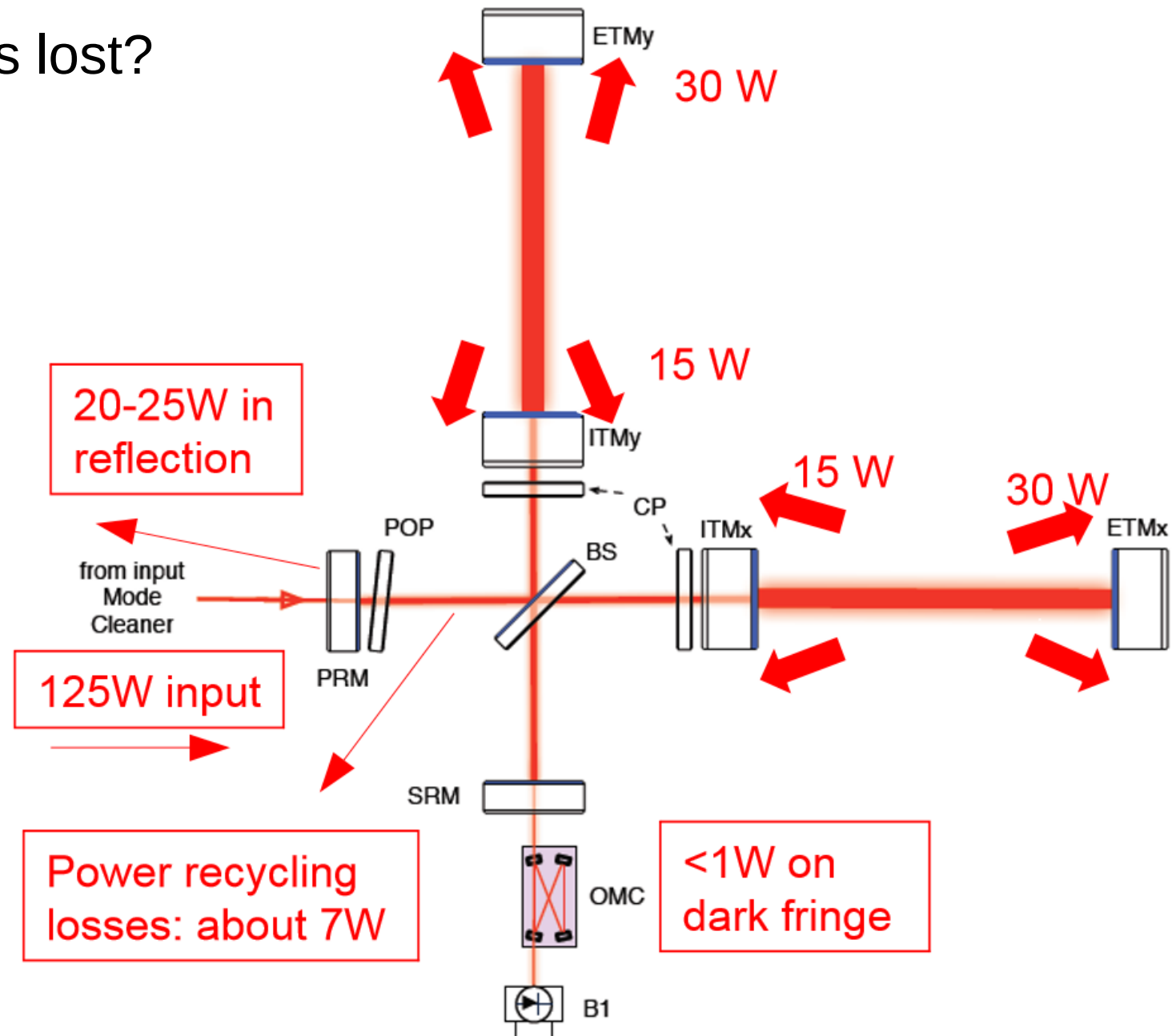
Mainly there are 4 sources:

- Secondary beams
- Rough surface scattering
- Point defect scattering
- Diffraction



Stray light

Where the light is lost?



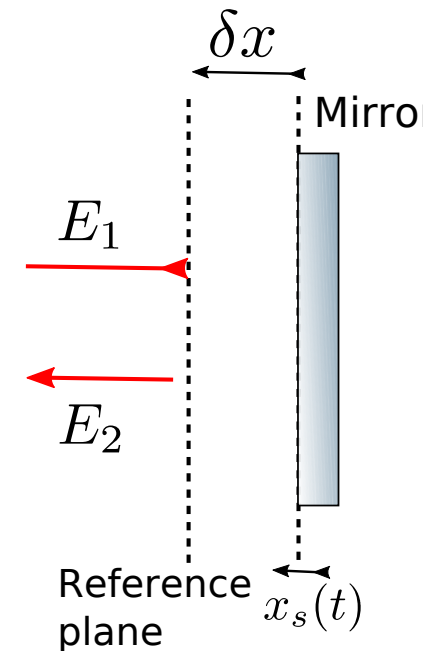
Stray light noise

- Phase change due to a moving element:

$$\phi(t) = \frac{4\pi}{\lambda}(x_0 + \delta x_{sc}(t)) = \phi_0 + \delta\phi_{sc}(t)$$

- Which can recouple as:

$$h = G \cdot \sin \phi_{sc}$$



Stray light noise II

- Two regimes:
 - Linear regime: small amplitude

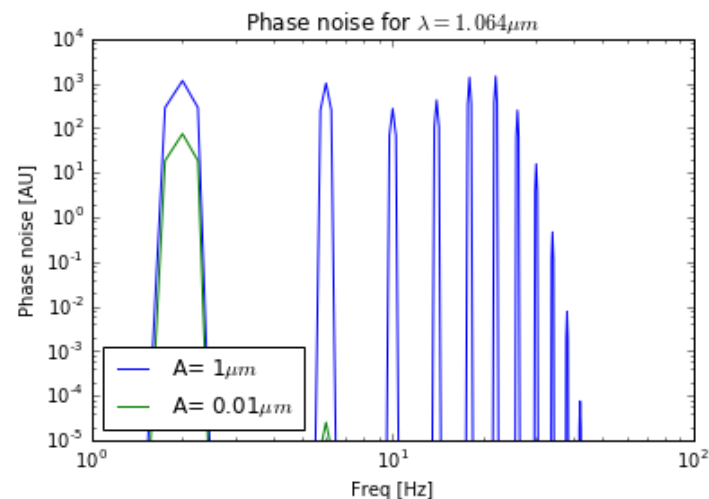
$$h(f) = G\phi_{sc}$$

- Non-linear regime

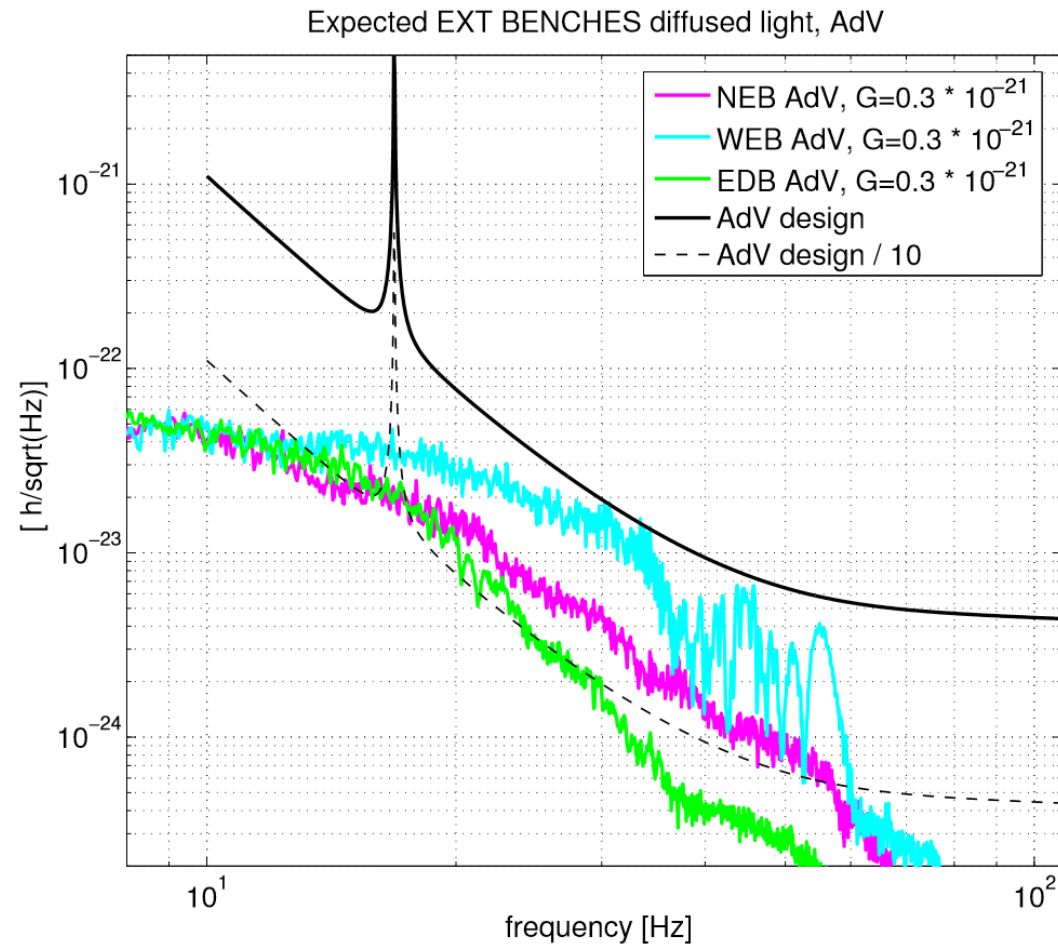
- Upconversion phenomena:

$$\sin(a \sin \psi) = 2 \sum_{n=1}^{\infty} J_{2n-1}(a) \sin[(2n-1)\psi]$$

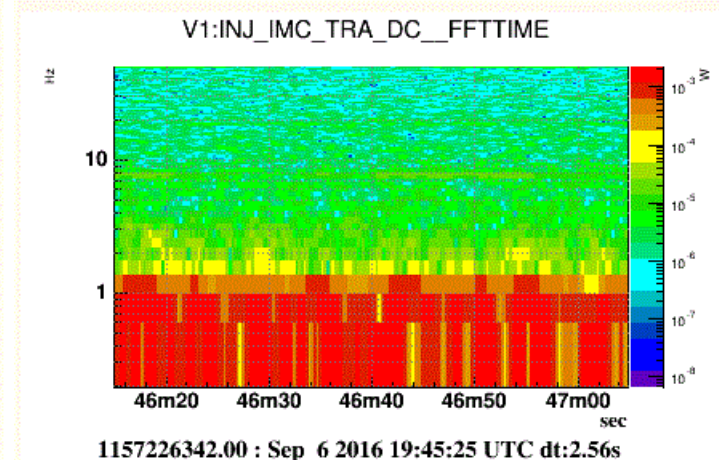
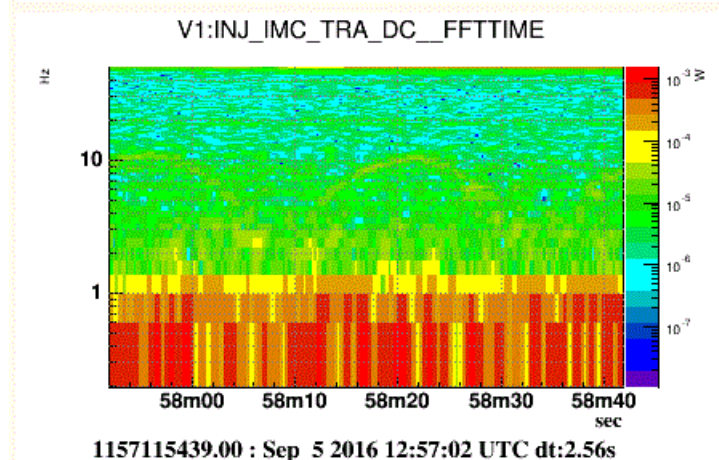
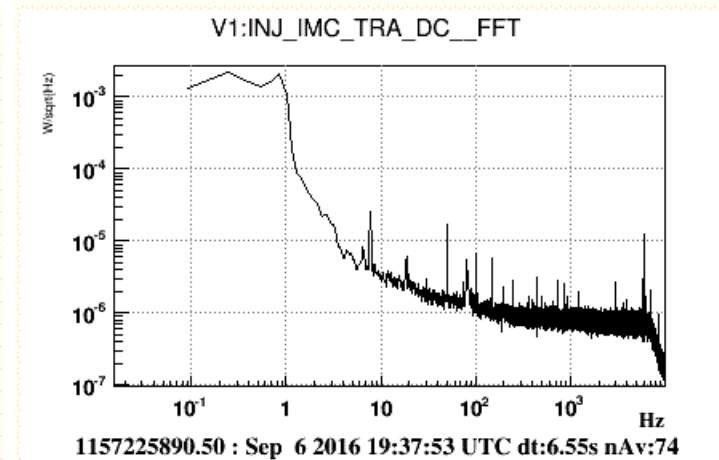
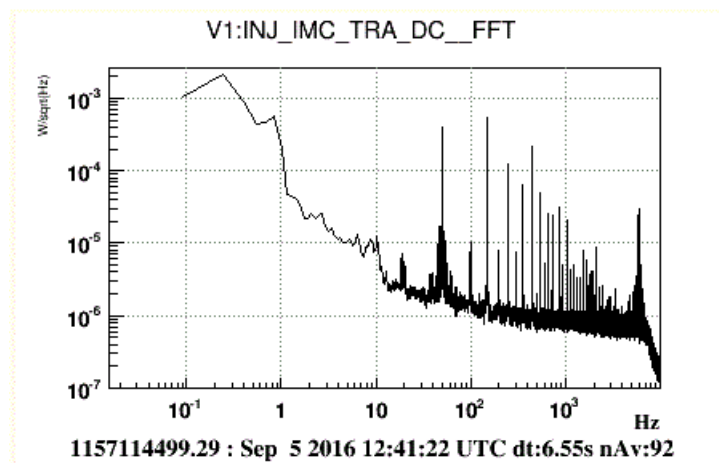
$$f_{max} = \frac{2A_x}{\lambda} 2\pi f_{sc}$$



Expected noise during sea activity



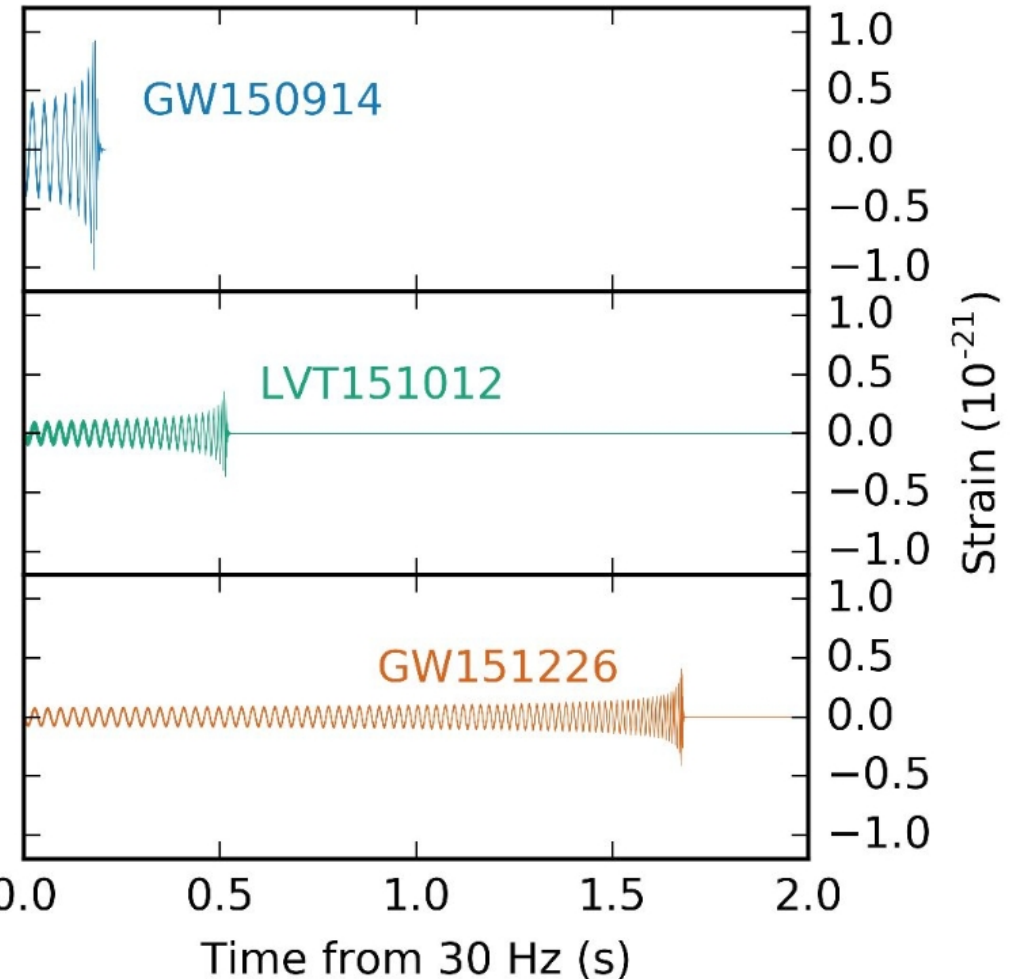
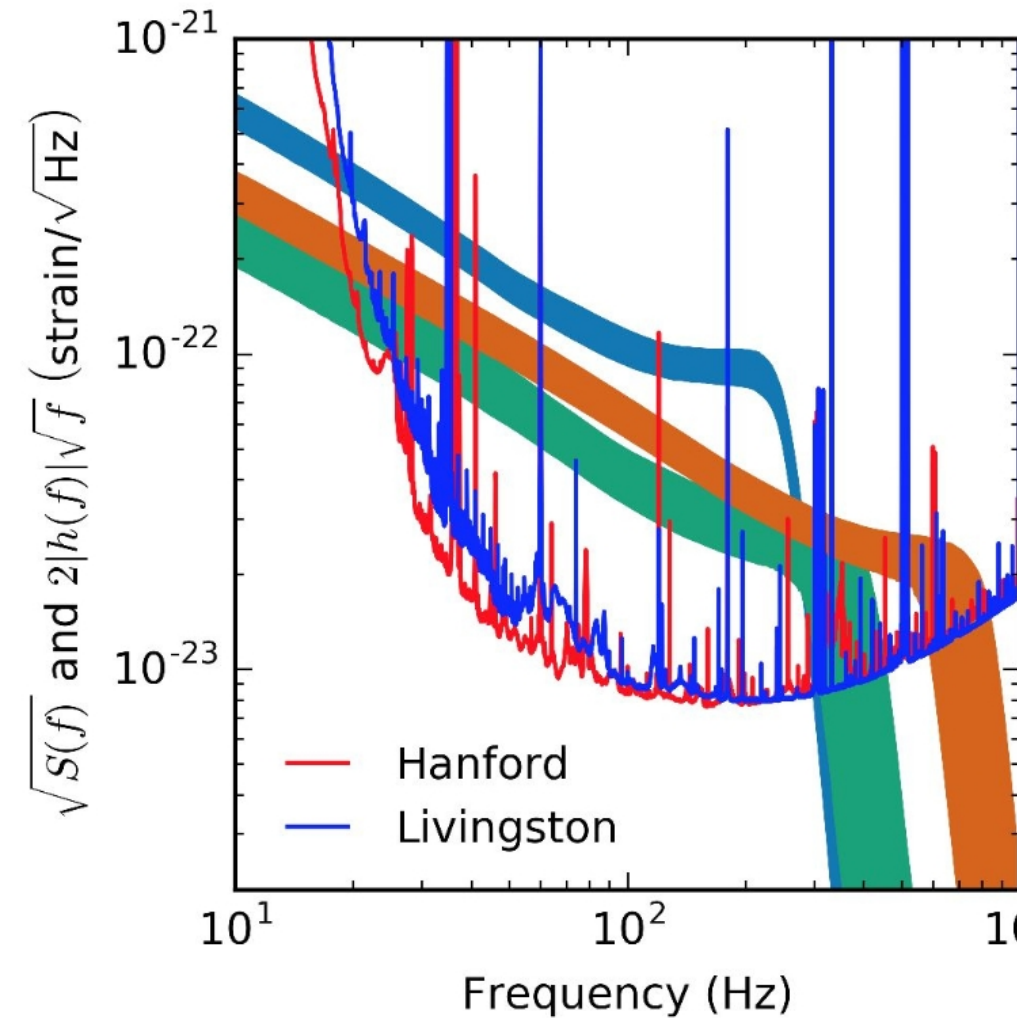
Recent case during AdV commisioning



Spurious light inside IMC at 50Hz before and after intervention



Importance of Low frequency sensitivity



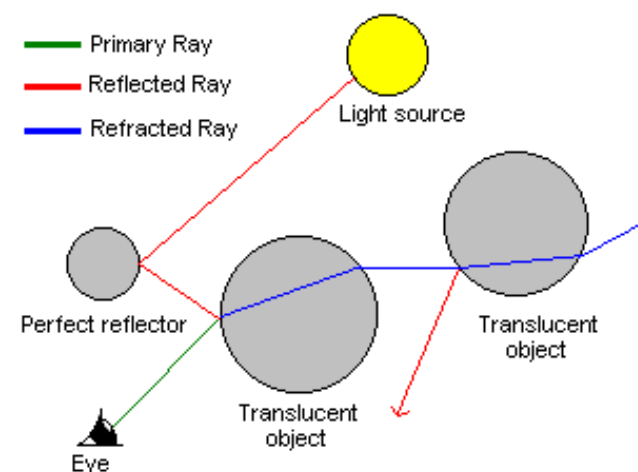
Ray tracing

- Basic technique to simulate ray as a particle

```

1 initialization;
2 Set up Objects;
3 for  $i \in \text{Range}(1, \text{MaxRays})$  do
4   Generate new Ray;
5   while exists Ray do
6     Find Interesection of Ray with Objects;
7     Select MinInteresection;
8     if MinInteresection = Null then
9       return 0
10    else if MinInteresection = Sensor then
11      Save the Ray on the Sensor;
12      return 0
13    else
14      Do Shading → Generate new Ray
15    end
16  end
17 end

```



Shading

Shading: Model of the interaction between light and matter, usually BRDF divided in:

- Reflection
- Transmission
- Scattering
- Absorption

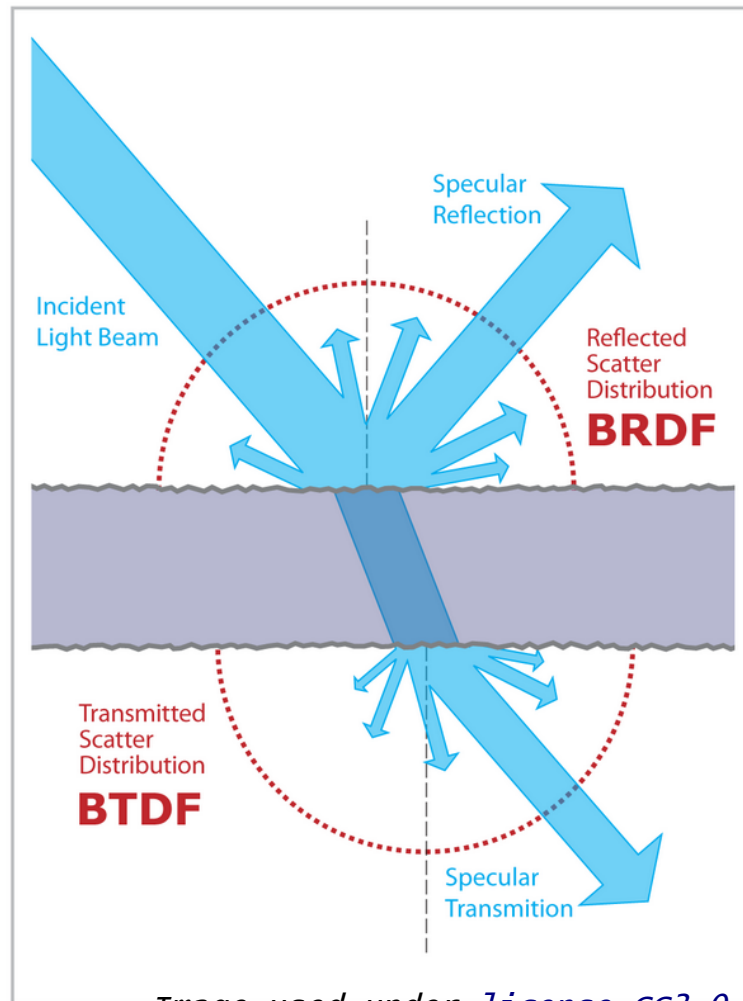


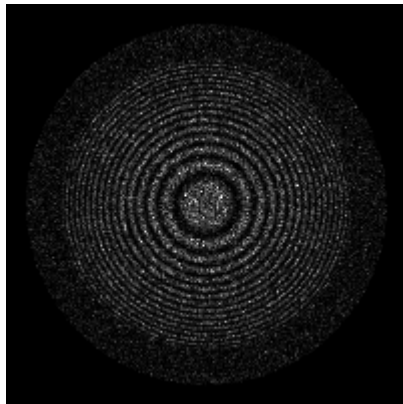
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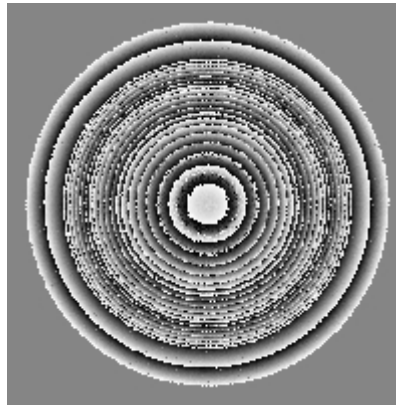
Interference

Wave property that can be simulated with ray tracing

- Transform Optical Path Length into Phase



Amplitude: $|E|^2$ [au]



$-\pi \leq \text{Phase} \leq \pi$

Interference pattern simulated with a Michelson configuration of a collimated beam with spherical mirrors of 2.2m and 9.2m



Coupling between light and mechanics

As has been previously explained, the movement of a mechanical component may induce noise on a specific frequency



Phase noise

- Modulation of a photon hitting a moving surface:

$$\Delta\phi(t) = \frac{2\pi}{\lambda}(\vec{s} - \vec{\omega}) \cdot \vec{X}(t)$$

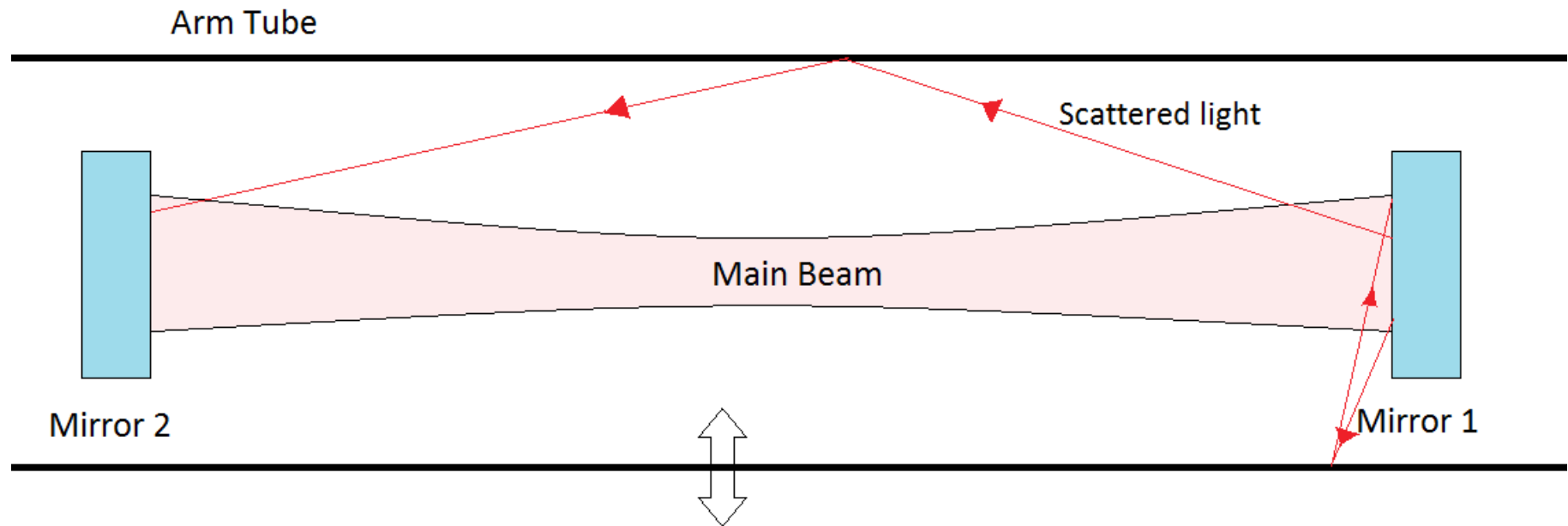
- Phase change on main beam and GW strain:

$$\Delta\phi_{rec}(t) = \text{Im}[\gamma(t)]/\sqrt{P}$$

$$\Delta\phi_{rec}(t) = \frac{4\pi L}{\lambda}h(t)$$



Noise on Fabry-Perot cavities



Total noise inside the
Fabry-Perot Cavity

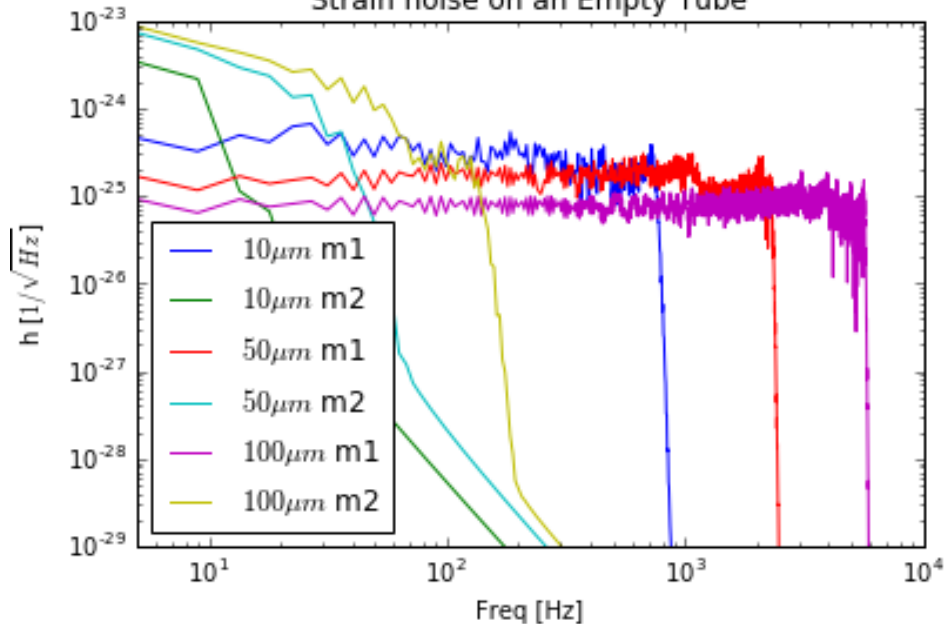
$$h_{tot}(f) = \frac{\lambda^2 \epsilon}{2^{2.5} \pi^2 L R_m} \sqrt{\frac{1}{n_s} \sum_{j=1}^n b(\theta_j) n_j(f)^2}$$



Noise on Fabry-Perot cavities

2 different cases:

Strain noise on an Empty Tube



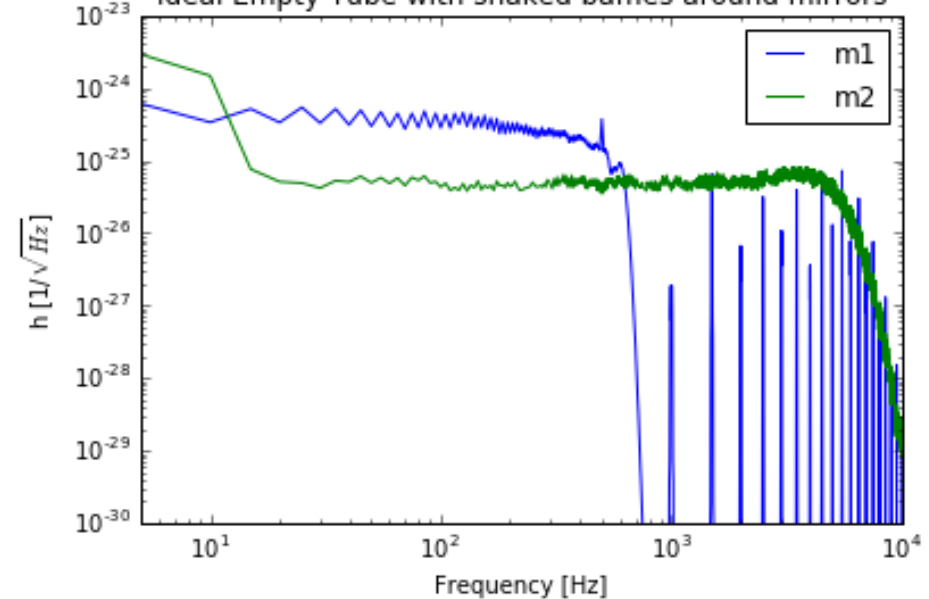
Tube shaking at 10Hz

Previous results obtained by Vinet et.al.

	Noise [$1/Hz$]
Analytical method	1.23×10^{-23}
Statistical	10^{-23}

Note: the statistical method was using seismic noise already in frequency domain

Ideal Empty Tube with shaken baffles around mirrors



Tube at 5Hz with $A=8\mu m$ $\rightarrow f_{\max}=0.5KHz$

Baffle at 500Hz with $A=1\mu m$ $\rightarrow f_{\max}=6KHz$



Conclusions and future work

- Conclusions
 - Ray tracing can be used to simulate some wave optics phenomena.
 - It is possible to estimate strain noise produced by stray light.
- Future work
 - Physics side:
 - Find new important cases to find out the contribution to the total noise.
 - Perform experiments to validate the results of the code.
 - SW side:
 - Optimize the code and include GPU computing.
 - Obtain input for realistic optomechanical components, probably from CAD files.



Thank you for your attention

