

Neutrino Oscillations and Astroparticle Physics (5)

John Carr

Centre de Physique des Particules de Marseille (IN2P3/CNRS)

Pisa, 10 May 2002

n High Energy Astronomy

Multi-Messenger Astronomy

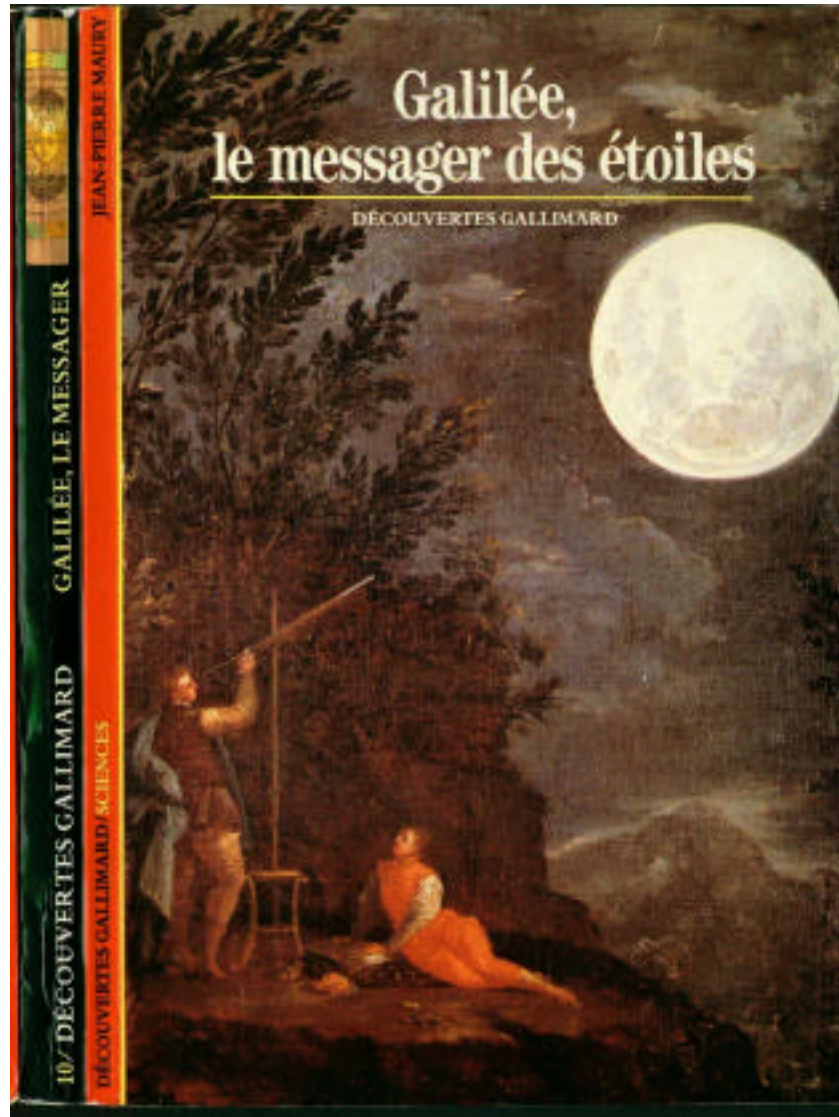
Cosmic Rays

Gamma Ray Astronomy

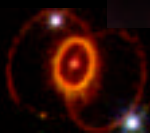
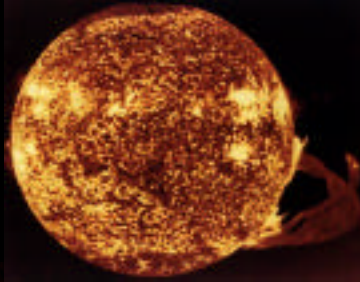
Neutrino Astronomy

(References)

Multi-Messenger Astronomy



Multi-Wavelength Photons
Charged Cosmic Rays
Neutrinos
Gravity Waves



Neutrinos
 (MeV: sun, SN
 GeV: atmosphere
 PeV: CR accelerators)



Cosmic ray
 particles $\rightarrow 10^{20}$ eV



Electromagnetic
 radiation $\rightarrow 100$ TeV



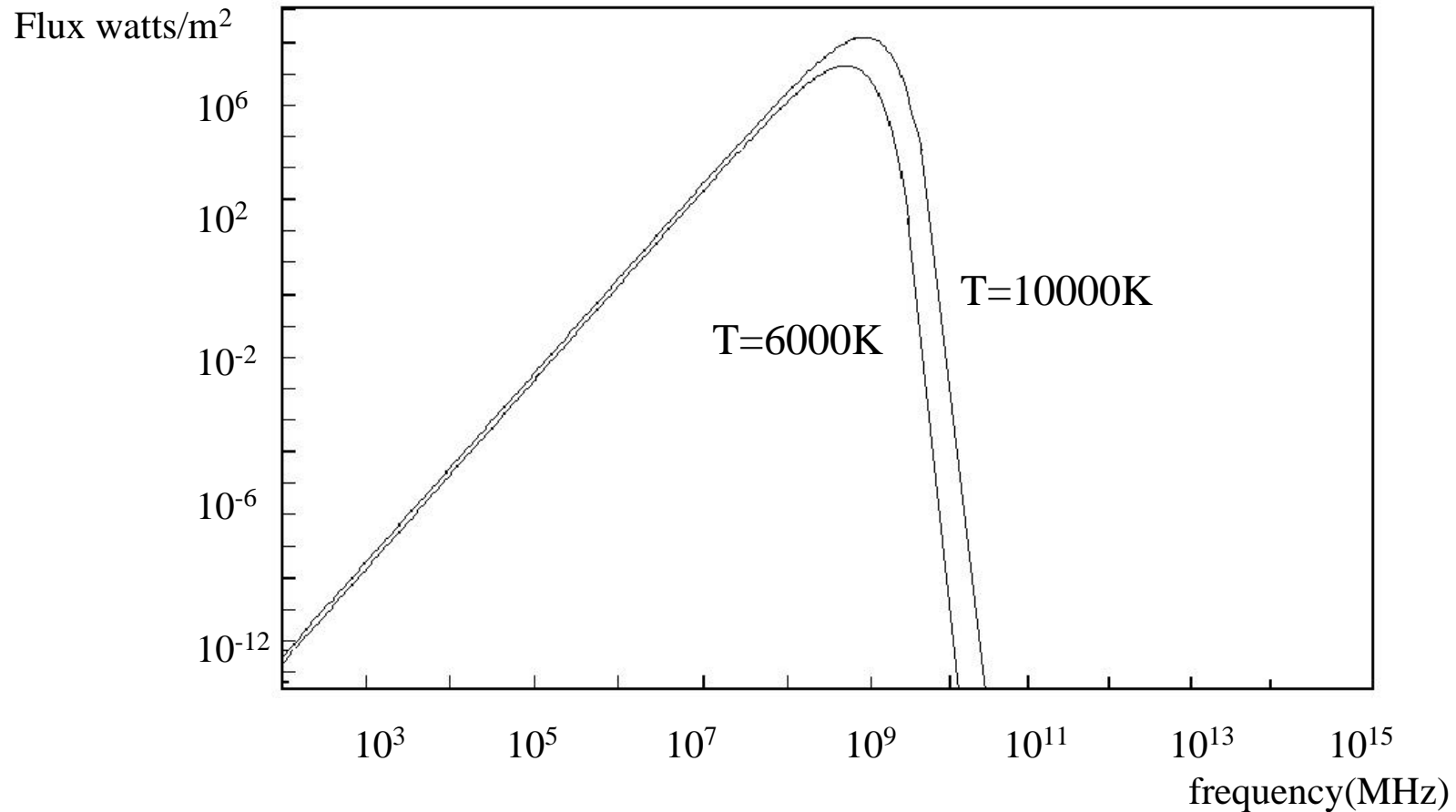
W Dark matter
s Monopoles
 Axions
 Grav. waves

v

u

(W. Hoffmann)

Thermal Radiation from Stars



Normal Stars surface temperature ~3000 to 30000K

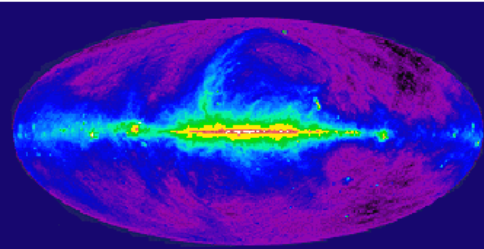
thermal radiation: radio ultra -violet

non-thermal radiation: X-rays, gamma rays

(higher in energy more extreme is the source)

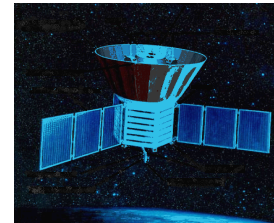
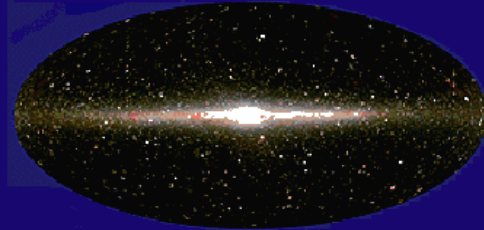
Multi-Wavelength Photons

Radio



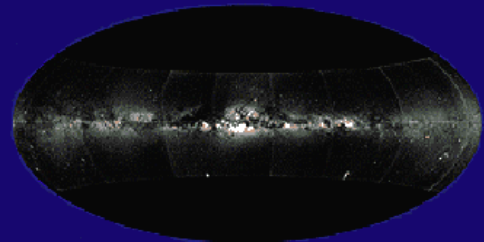
Radio télescope
de Bonn

Infrared



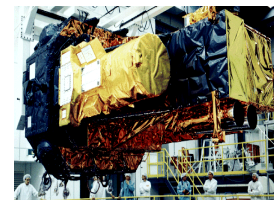
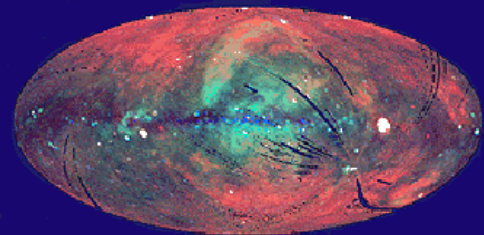
Satellite
COBE

Visible light



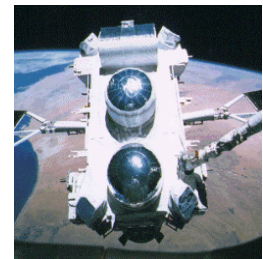
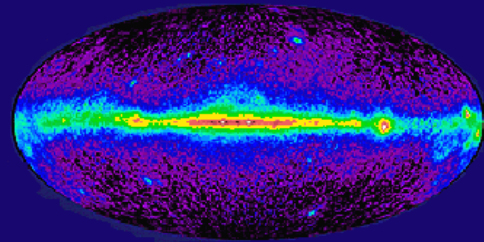
Télescope du
Mont Palomar

X-ray



Satellite
INTEGRAL

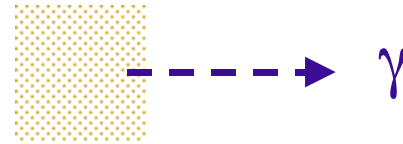
Gamma Ray



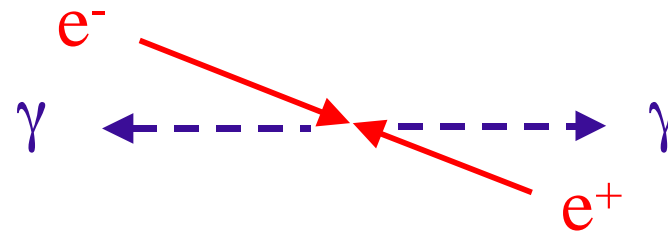
Satellite
CGRO

Production Mechanisms of Photons

Hot plasma (surface of stars)



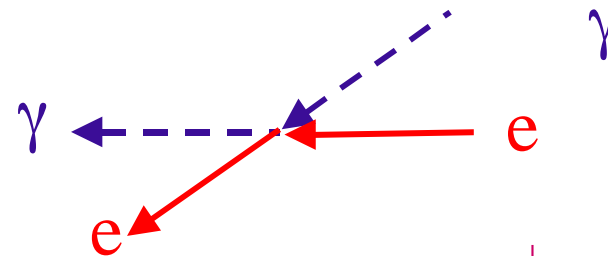
Annihilation of matter/antimatter



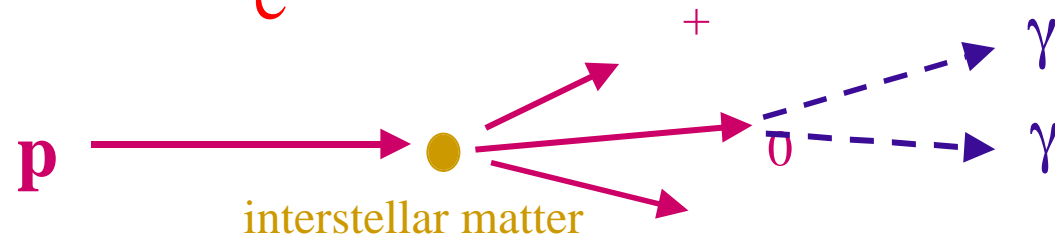
Bremsstrahlung /
Synchrotron Radiation



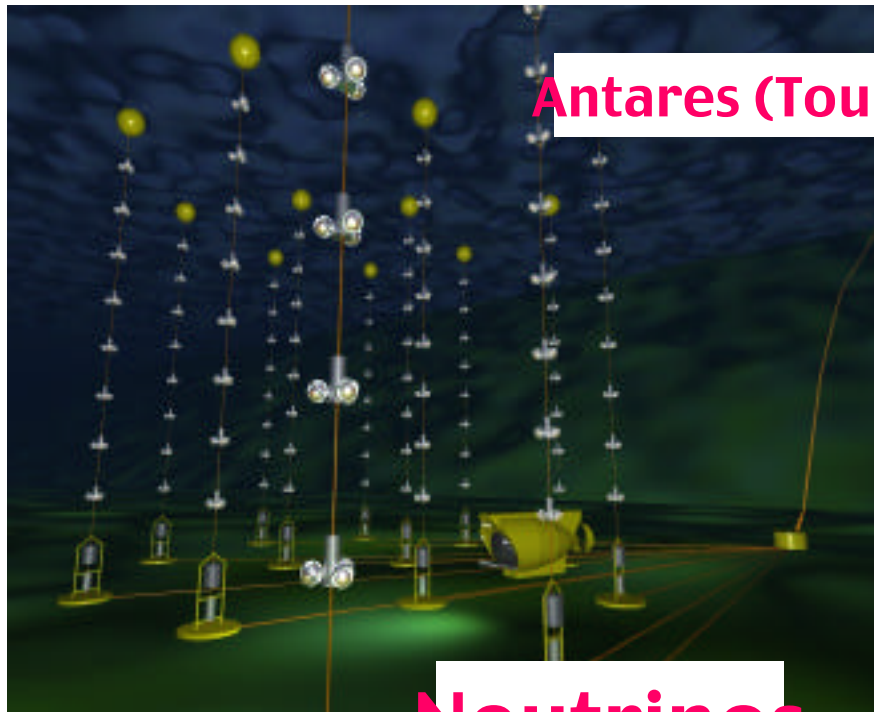
Inverse Compton Scattering



High energy showers



Non-Photonic Astronomy

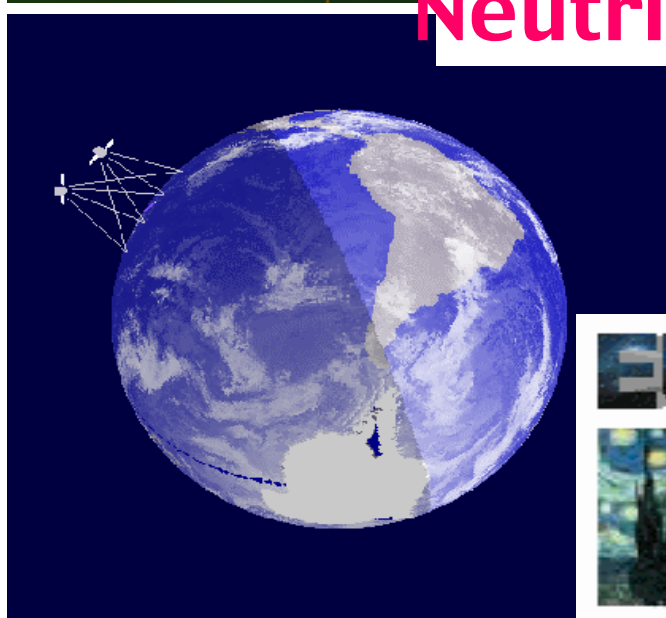


Antares (Toulon)



Gravitational Waves

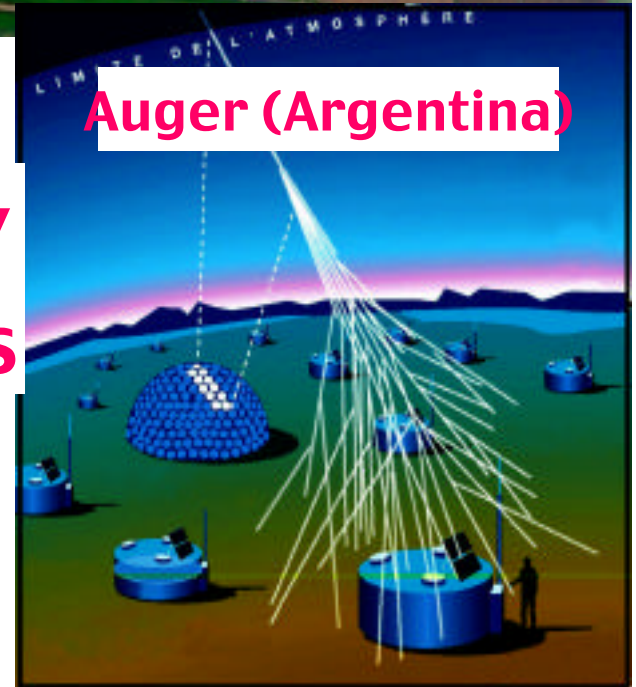
Virgo (Pisa)



Neutrinos

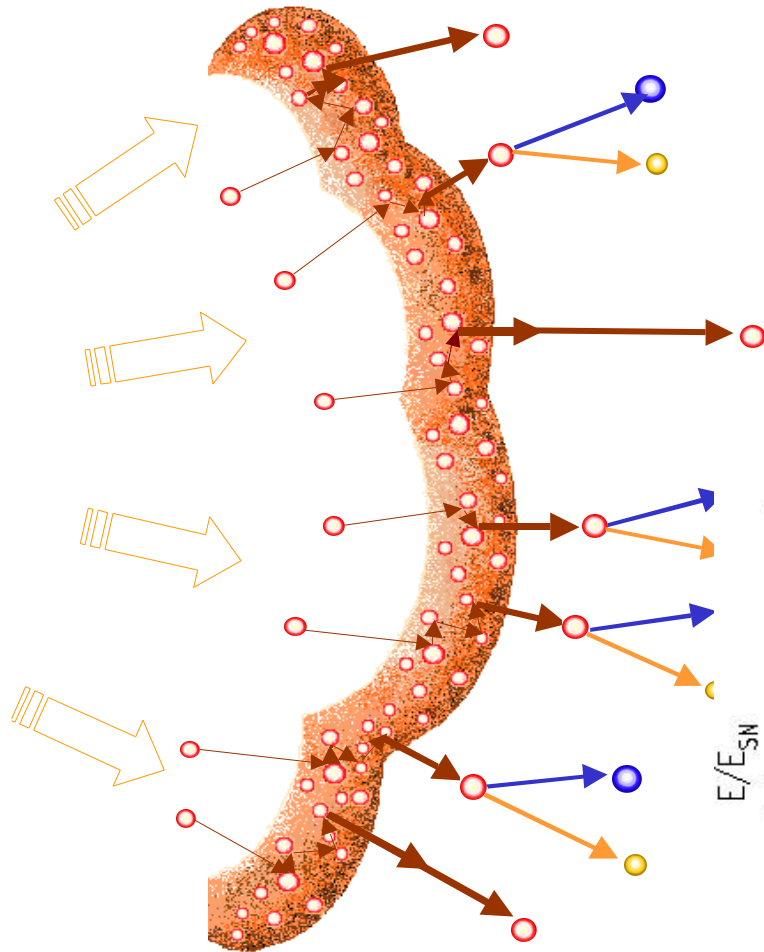
High Energy
Cosmic Rays

EUSO



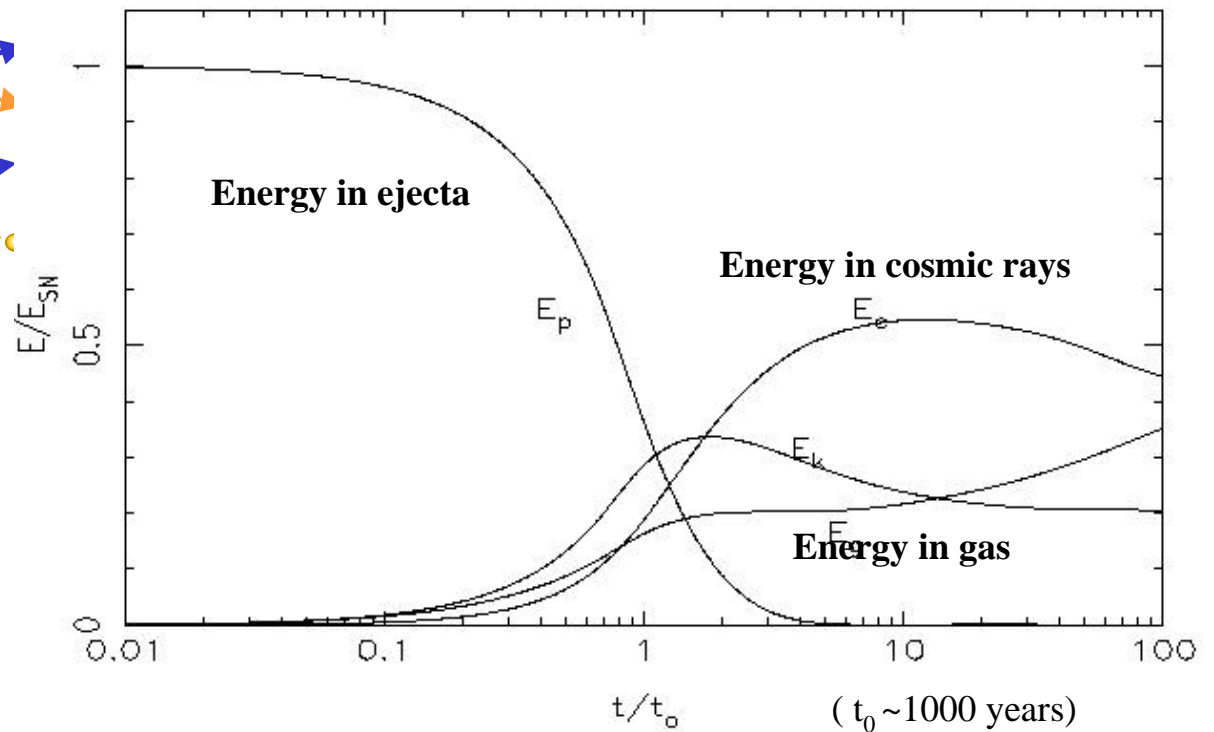
Auger (Argentina)

Acceleration Mechanism of Charged Particles



Simulations indicate can get
~ 50% of energy of supernova explosion
Cosmic Rays by ~1000 yrs

Example of Supernova Remnants



Production Mechanism of Neutrinos

p accelerated in shock waves:

non-relativistic supernova remnants

relativistic quasars/microquasars

p interact with interstellar matter and produce showers :

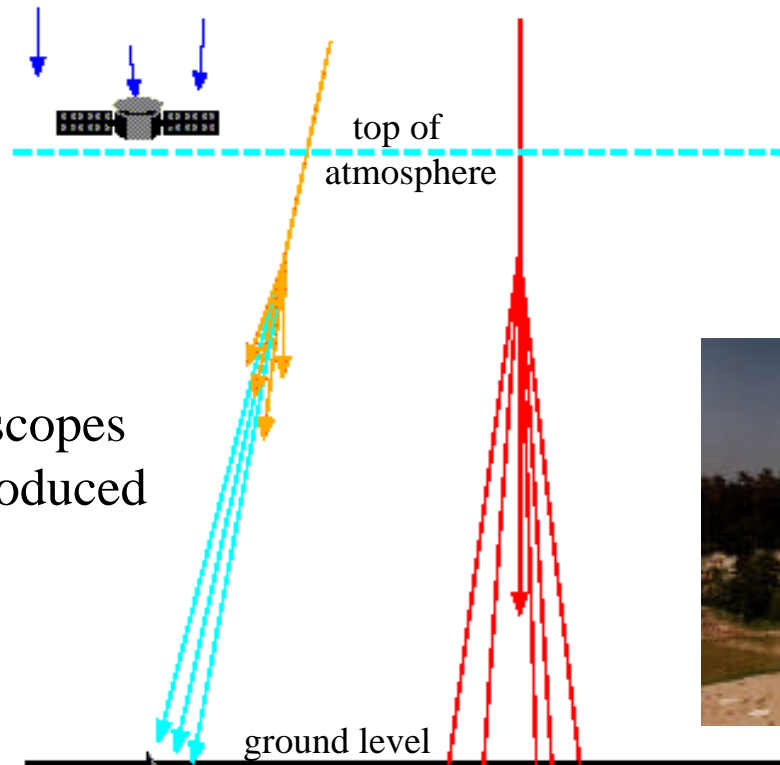
$$p/A + p/ \quad \quad \quad 0 \quad + \quad \pm \quad + \quad \dots$$

$$\gamma \gamma \quad \nu_{\mu} \mu$$

$$\nu_{\mu} \nu_e e$$

Types of Cosmic Ray Detectors

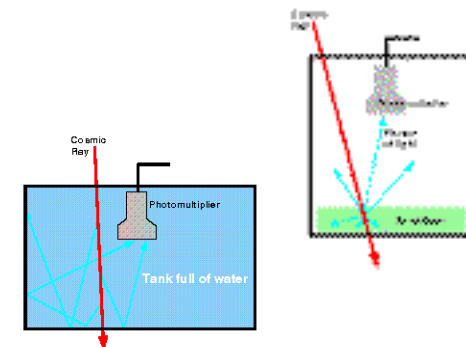
Satellites



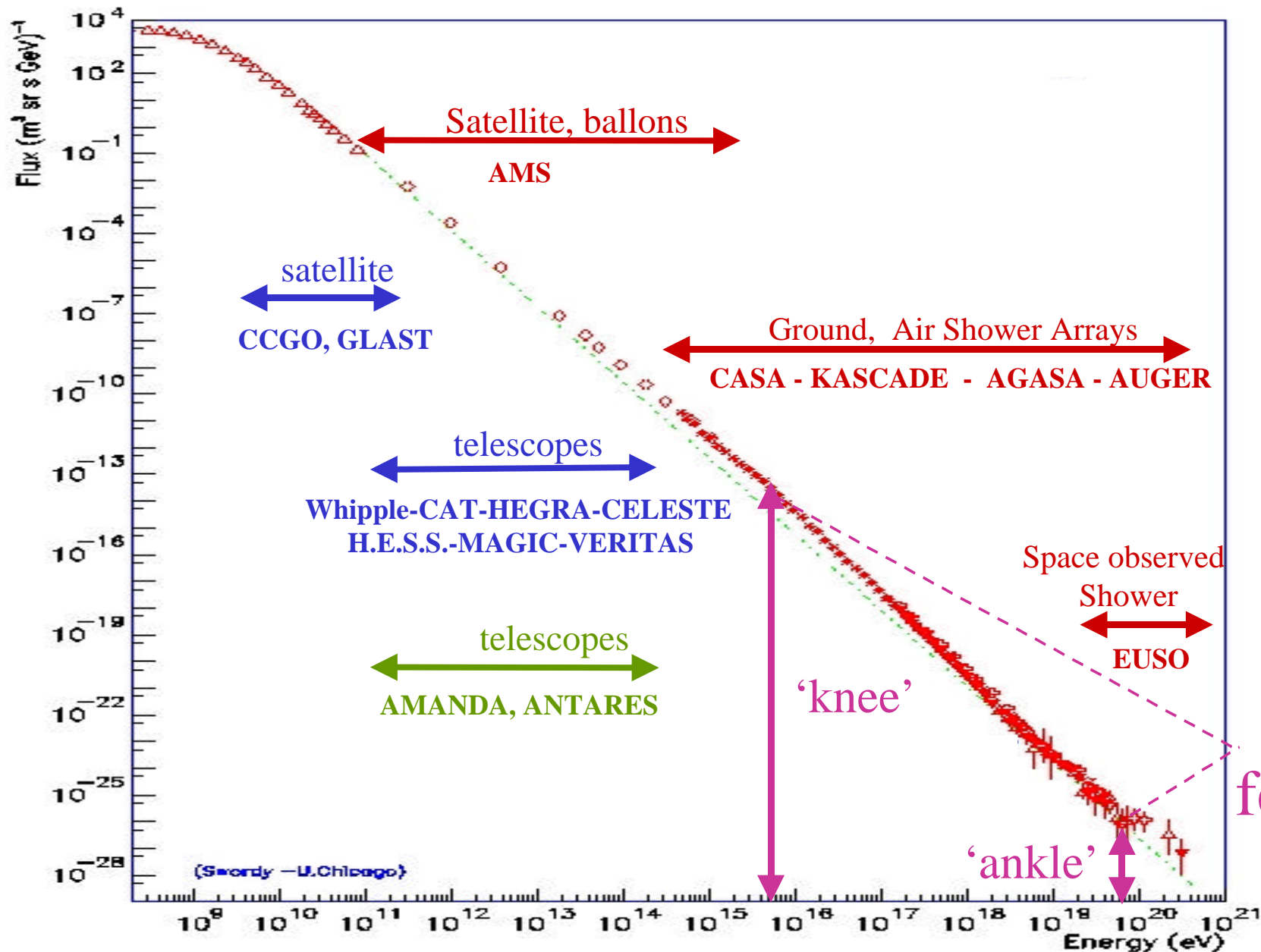
Ground based telescopes looking at light produced in atmosphere



Arrays of particle detectors

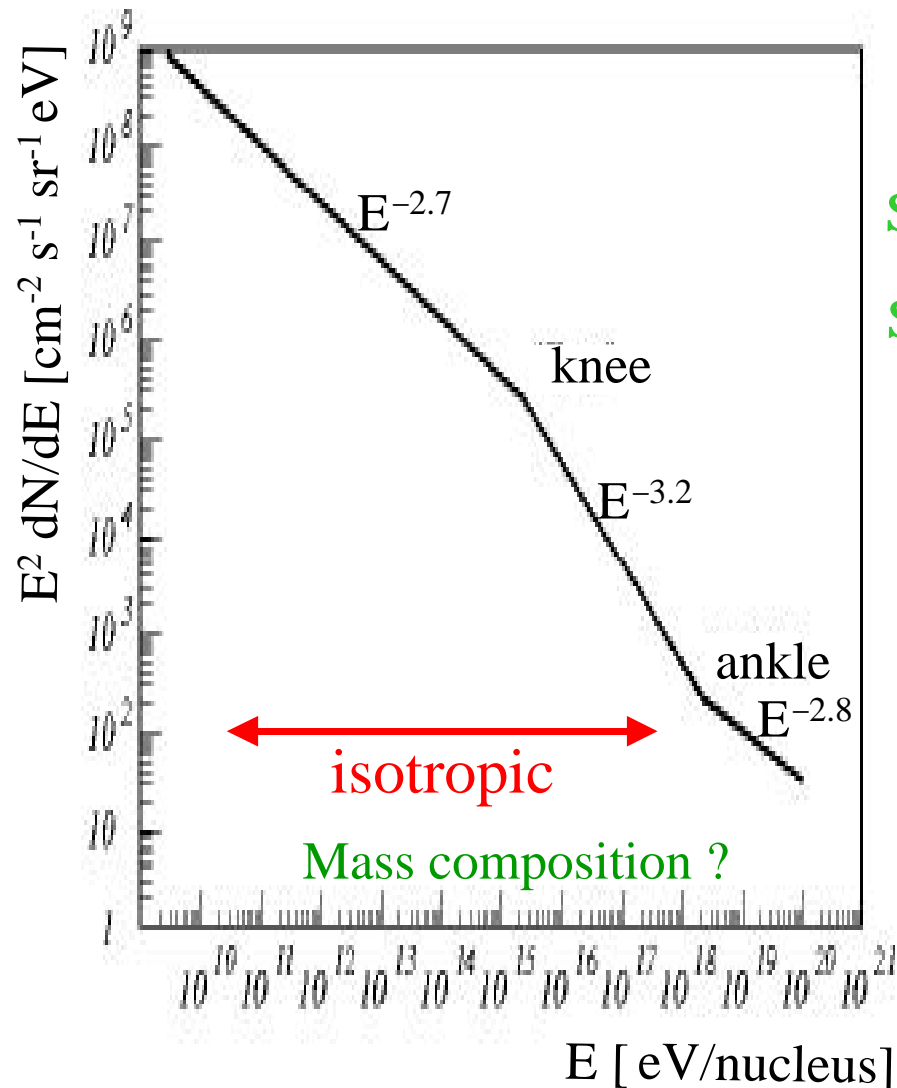


Charged Cosmic Ray Energy Spectrum



Why these features ?

Features of Cosmic Ray Spectrum



Ingredients of models:

$$\frac{dN}{dE} \propto E^{\alpha}$$

source
+
propagation

Source acceleration: $\alpha = -2.0$ to $-2.2, \dots$

Source cut-off $E < 10^{18} Z \left[\frac{R}{\text{kpc}} \right] \left[\frac{B}{\mu\text{G}} \right] \text{eV}$

Diffusion models $\alpha = -0.3$ to -0.6

GZK cut-off on CMB $E \approx 7 \cdot 10^{19} \text{ eV}$

‘Conventional Wisdom’:

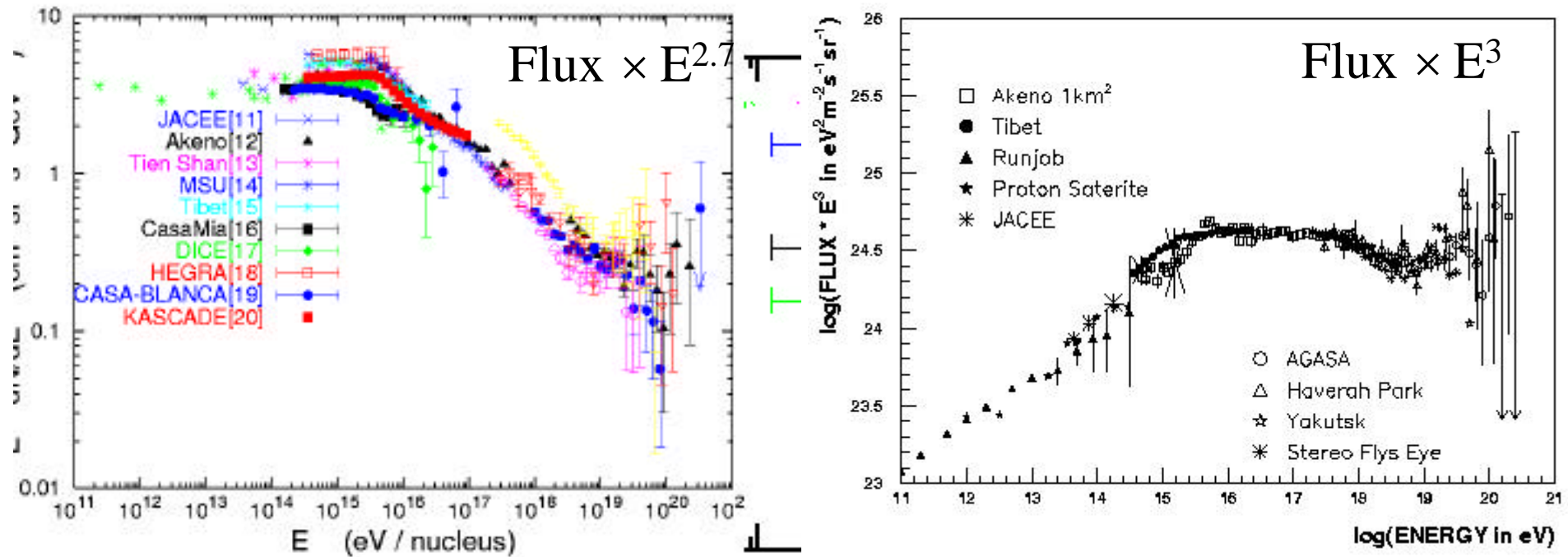
Galactic SNR $E < 3 \cdot 10^{18} \text{ eV}$

Galactic losses $E > 4 \cdot 10^{14} \text{ eV}$

Extragalactic $E > 3 \cdot 10^{18} \text{ eV}$

exotic $E > 7 \cdot 10^{19} \text{ eV}$

Cosmic Rays Spectrum: Knee and Ankle



Explanations of knee ($E \sim 3 \cdot 10^{15}$ eV)

Astronomy type explanations

- Galactic de-confinement
- Single dominant source
- Single SNR acceleration multiple SNR acceleration

Particle Physics type explanations

- Absorption on massive neutrinos in galaxy

Various interactions invoked to give threshold at $E = 3 \cdot 10^{15}$ eV.

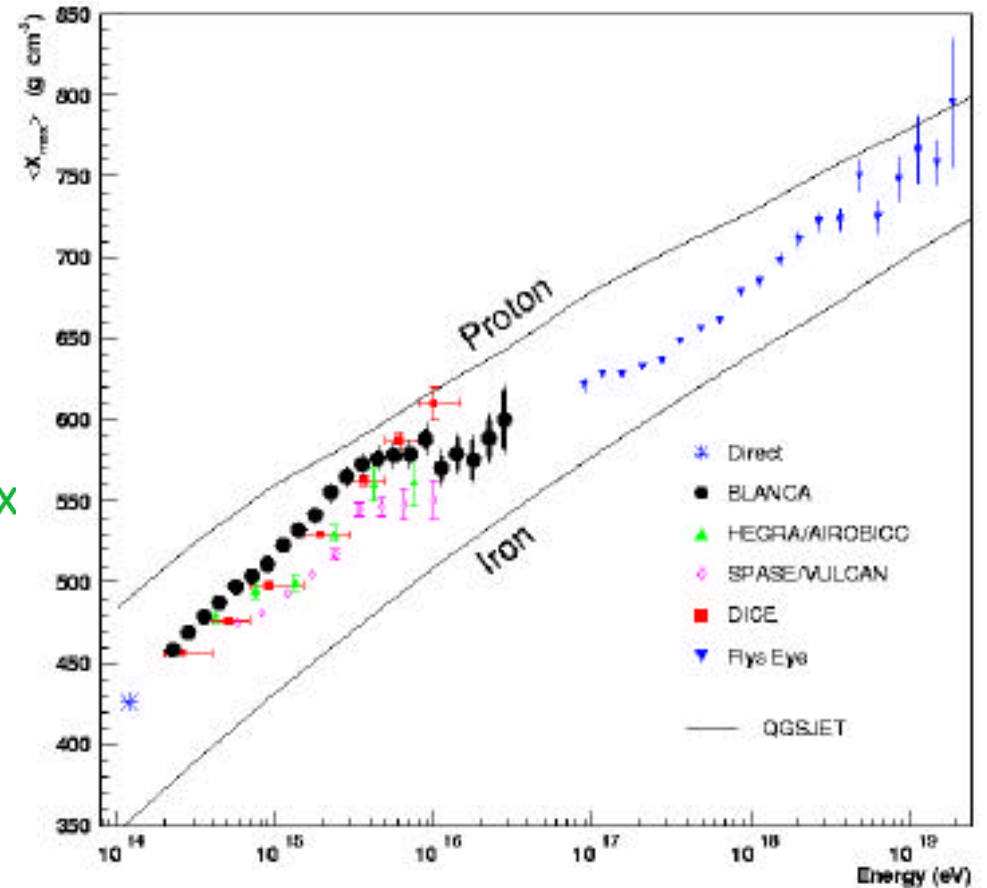
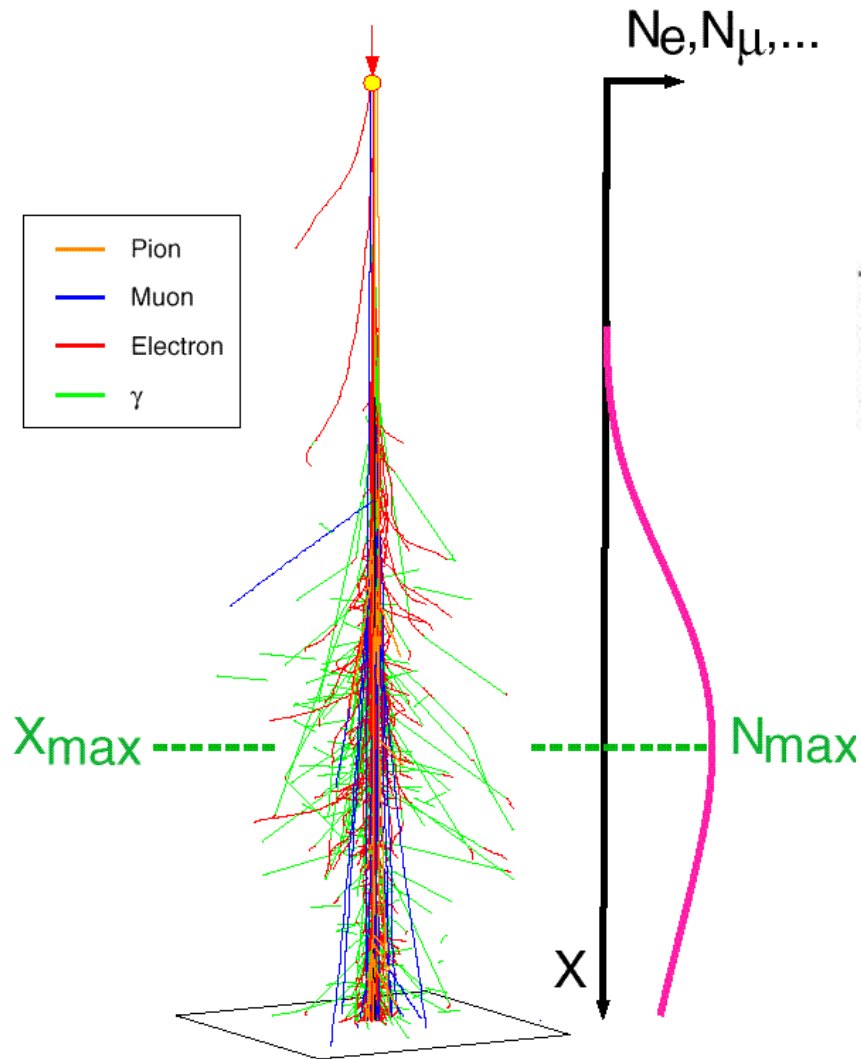
eg. $p + \bar{\nu}_e \rightarrow n + e$, with $M(\bar{\nu}_e) = 0.1$ eV

$p + \bar{\nu}_\mu \rightarrow \mu + \bar{\nu}_\mu$, with $M(\bar{\nu}_\mu) = 100$ eV

(not very convincing due to various problems: mass, rates)

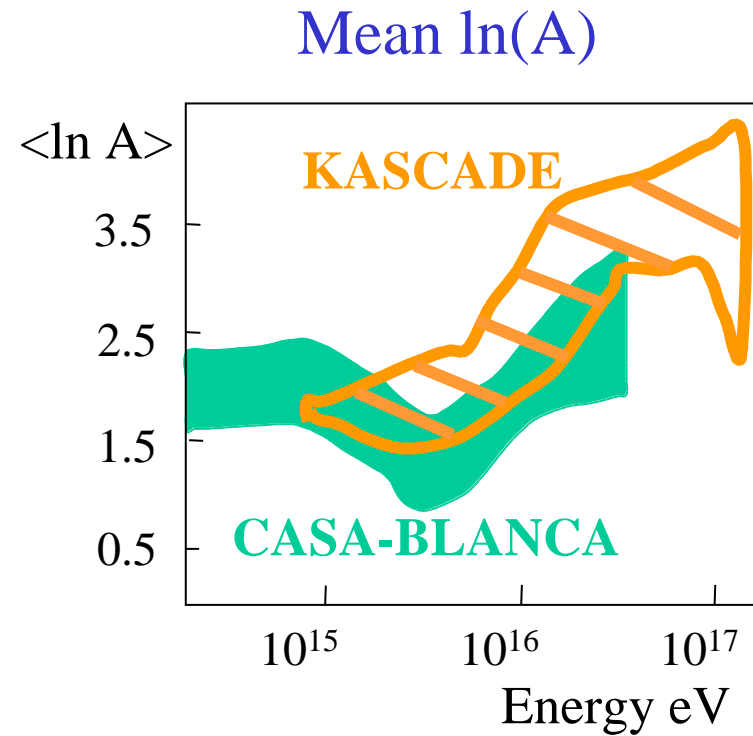
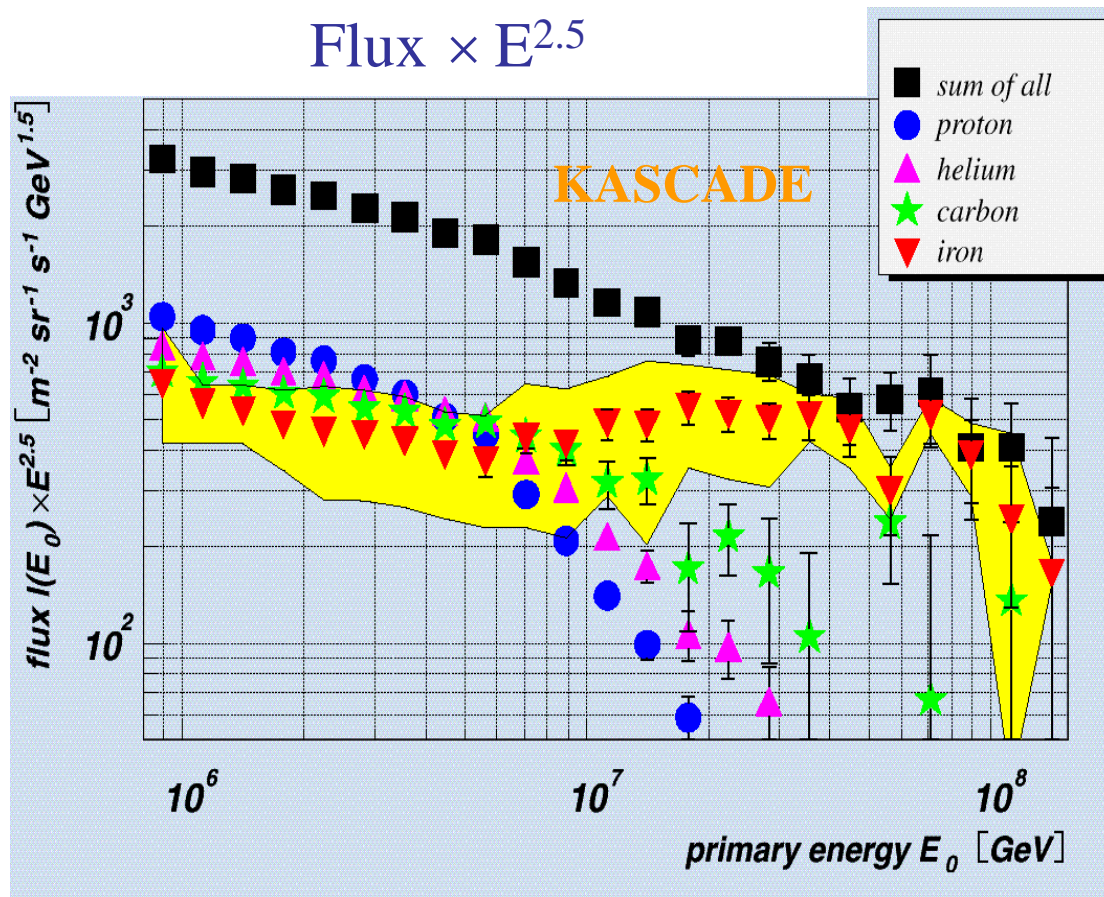
- New interaction effects in atmosphere

Mass composition from shower depth



Mass composition at knee

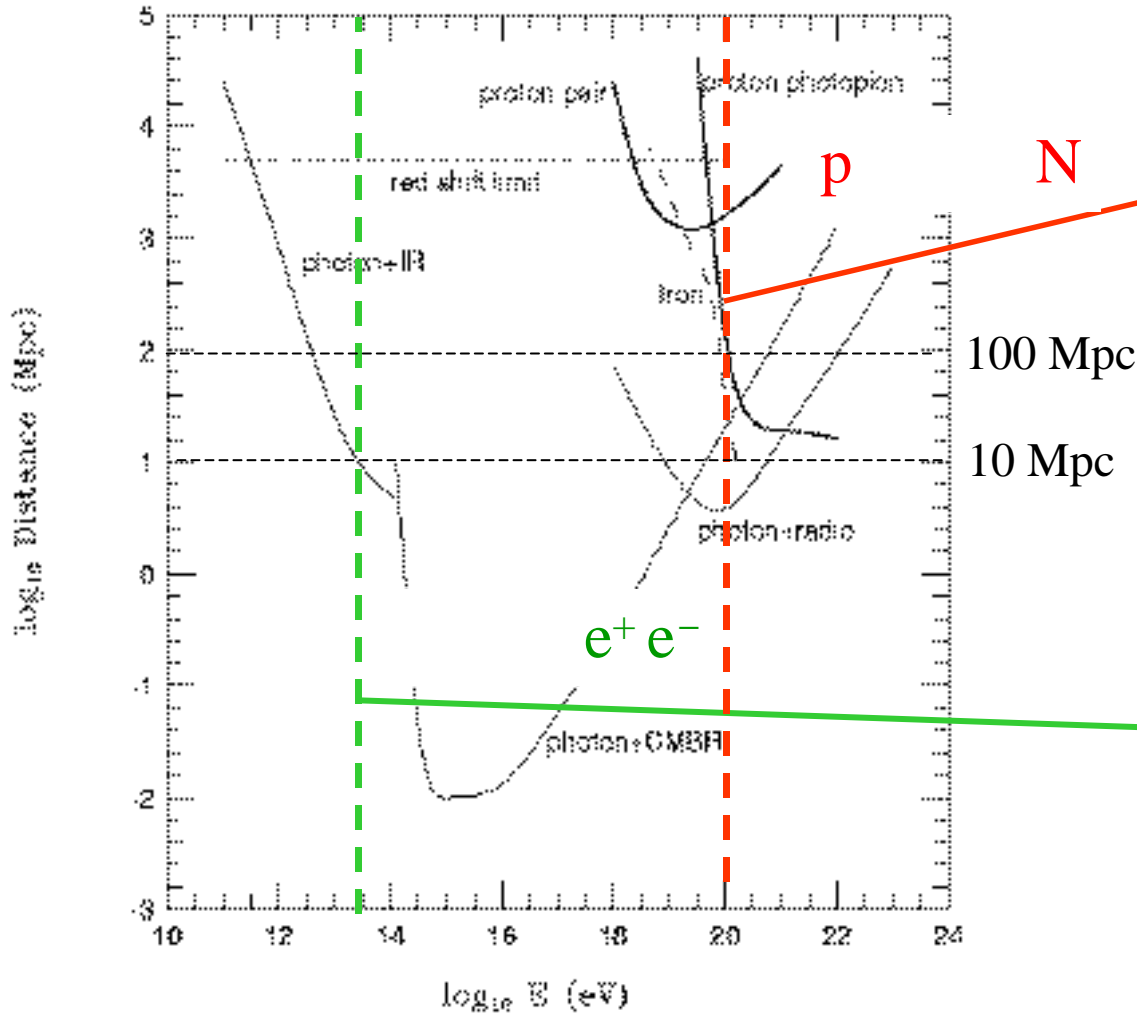
Average shower depth and ratio N_μ / N_e sensitive to primary mass
 (NB. Mass composition extracted is very sensitive to Monte Carlo simulation)



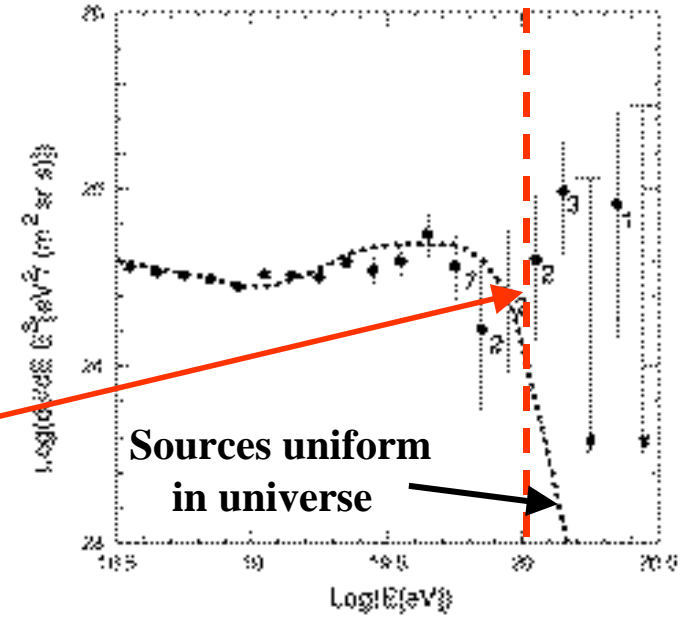
KASCADE series of knees at different energies: p,He,...,C,...,Fe.
 E(Knee) Z knee due to source confinement cut-off ?

'GZK cutoff'

Interaction with background
(infrared and 2.7K CMBR)

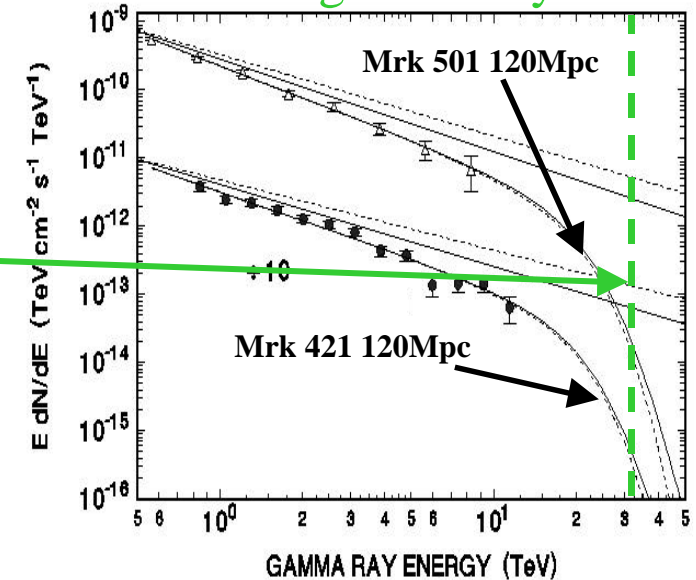


HE cosmic rays



Sources uniform
in universe

HE gamma rays



Explanations of Ankle/ $E > 10^{20}$ eV events

Astronomy type explanations

- ‘Bottom-Up’ : acceleration
 - pulsars in galaxy,
 - radio lobes of AGN (proximity a problem due to GZK, also should see source)

Particle Physics type explanations

- ‘Top-Down’ : decay of massive particles
 - GUT X particles with mass $> 10^{20}$ eV and long lifetimes
 - Topological defects
 - Neutrinos as messenger particle
- New Physics