High Intensity Laser-Solid Interaction: Collisionless Absorption and Instabilities

Andrea Macchi

Istituto Nazionale per la Fisica della Materia (INFM)
Dipartimento di Fisica "Enrico Fermi", Università di Pisa
www.df.unipi.it/~macchi



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- "Fast" electron transport and instabilities
 - Resistivity effects

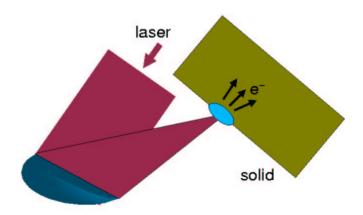
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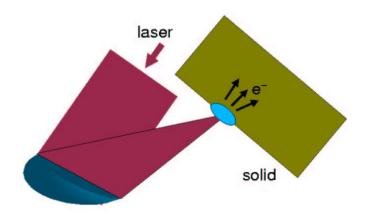
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Laser–Solid Interaction is a route for laser energy conversion into thermal or suprathermal electrons and ions and into coherent and incoherent XUV radiation.

We may identify four stages of the interaction:

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- 2. collisional absorption and plasma heating

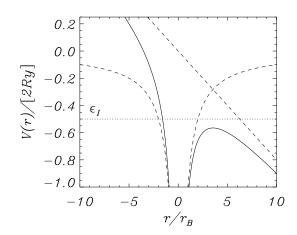
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- 4. electron energy transport and conversion (radiation, ions, fields . . .)

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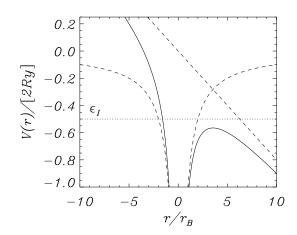
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Oscillating free electrons contribute to collisional ionization (quiver energy $\mathcal{E}_{osc} \simeq 6$ keV at $I_L \lambda_L^2 = 3.5 \times 10^{16}$ W cm⁻² μ m²).

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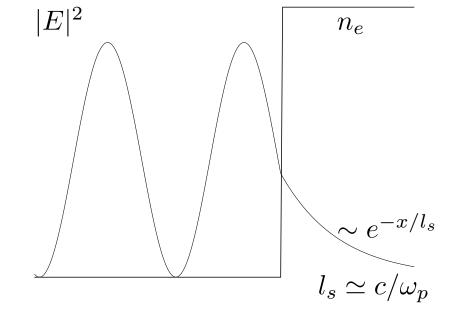
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$$n = 1 - \omega_p^2/\omega^2 < 0 \qquad (\omega_p = \sqrt{4\pi n_i e^2/m_e})$$

$$B_z(x,t) = B_z(0^+)e^{iky\sin\theta - x/l_p - i\omega t} + \text{ c. c.},$$

$$E_y(x,t) = -\frac{i\omega l_p}{c}B_z(0^+)e^{iky\sin\theta - x/l_p - i\omega t} + \text{ c. c.},$$

$$E_x(x,t) = \left(\frac{\omega l_p}{c}\right)^2\sin\theta B_z(0^+)e^{iky\sin\theta - x/l_p - i\omega t} + \text{ c. c.},$$

$$l_p = \frac{c}{\omega_p}\left(\cos\theta(1 - \omega^2\cos^2\theta/\omega_p^2)\right)^{-1/2},$$

$$\frac{B_z(0^+)}{B_{z,i}} = \frac{2\mathsf{n}^2\cos\theta}{\sqrt{\mathsf{n}^2 - \sin^2\theta} + \mathsf{n}^2\cos\theta}$$

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In addition to the electric force at frequency ω , the force at the plasma surface has a magnetic component at frequencies 0 and 2ω :

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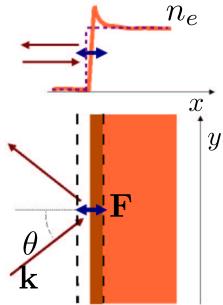
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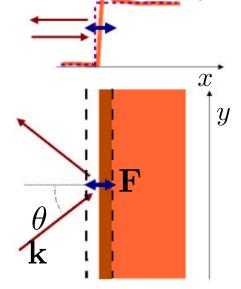
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It turns out that the dynamics at the plasma surface is dominated by *the force component* normal to the surface. The latter strongly depends on polarization and incidence angle.



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Effective dielectric function becomes (Drude's model)

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$$n_e \partial_t T_e = -\nabla \cdot (\kappa \nabla T_e), \qquad \kappa = \frac{n_e T_e}{m_e \nu_{eI}}, \quad \nu_{eI} = \frac{Z n_e e^4 \ln \Lambda}{m_e^{1/2} T_e^{3/2}}$$

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Weak absorption dependence on target material [Price et al, PRL **75**, 252 (1995)].

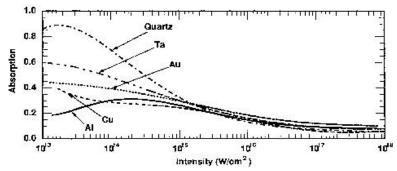


FIG. 1. Absorption fraction vs peak laser intensity for aluminium, copper, gold, tantalum, and quartz targets. In Figs. 1, 3, 4, and 5 laser intensity is the temporal and spatial peak value of the laser intensity.

Where does absorption come from?

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- mode conversion (i.e. linear or nonlinear excitation of waves)
- kinetic effects (the distribution function is modified leading to a different phase between ${\bf J}$ and ${\bf E}$.

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From $4\pi J_x + \partial_t E_x = 0$ one obtains $J_x E_x = -\partial_t E_x^2/8\pi$.

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This constraint may be however violated: non-steady state effects, aperiodic motion, 2D effects, . . .

The solution of the Vlasov-Maxwell system should in principle contain all effects leading to absorption.

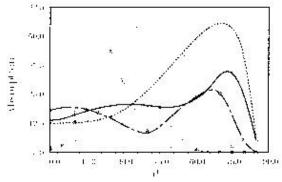


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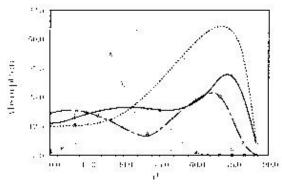


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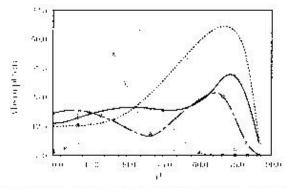


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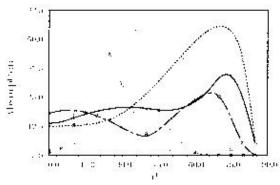


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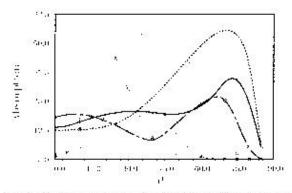


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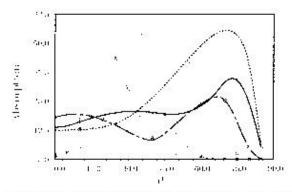


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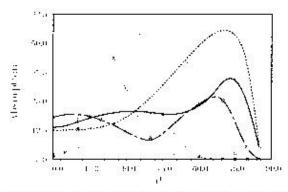
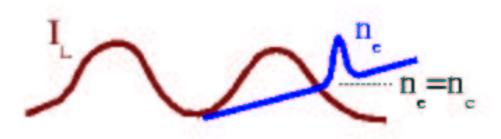


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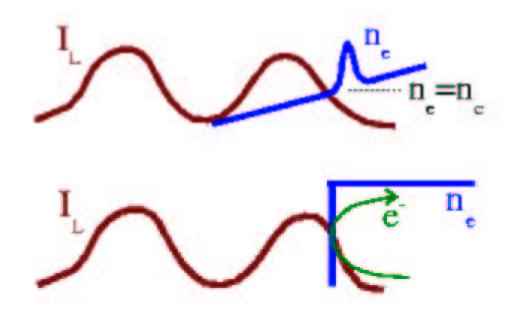
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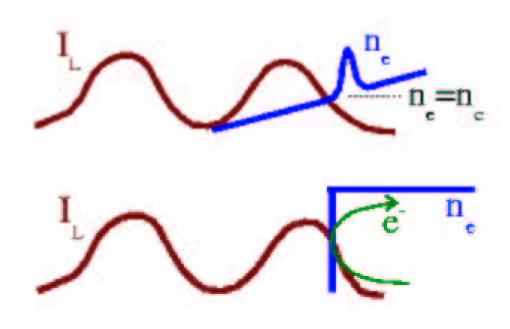
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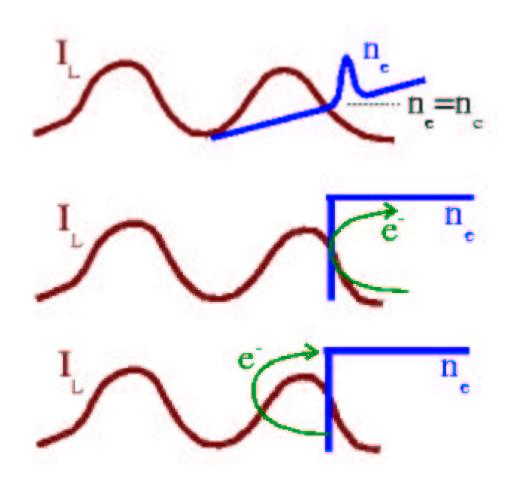
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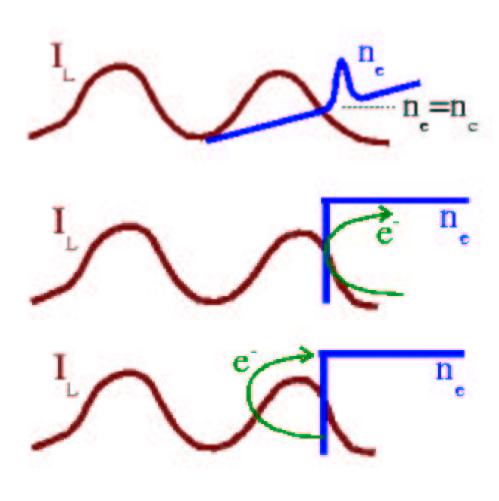
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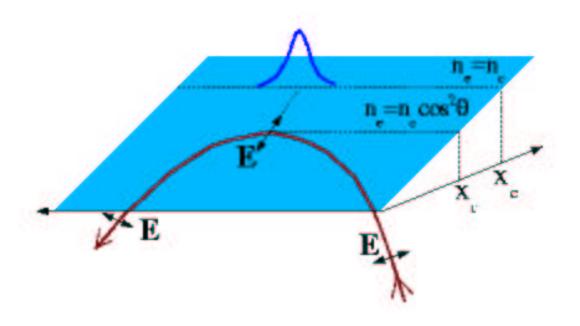
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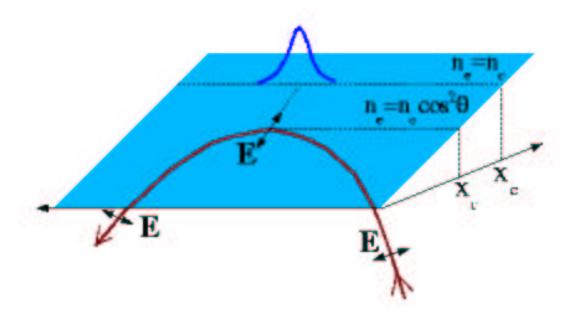
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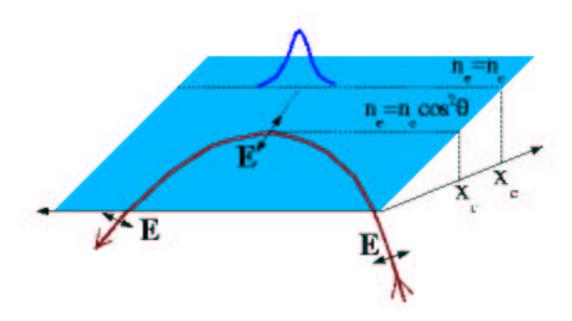
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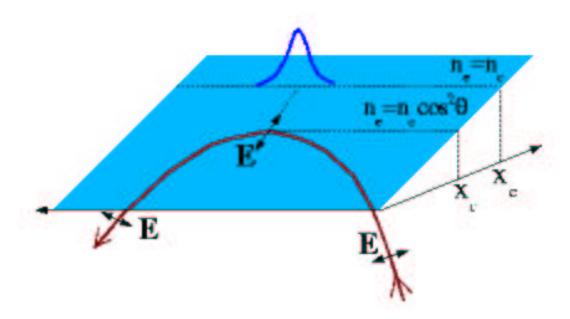
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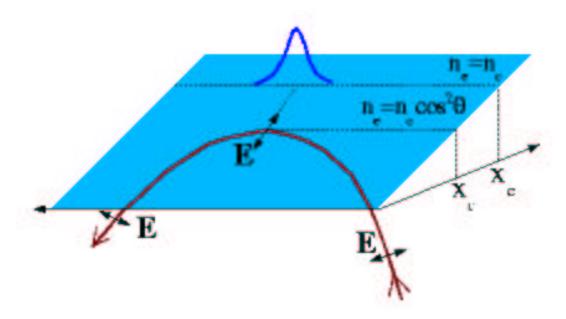
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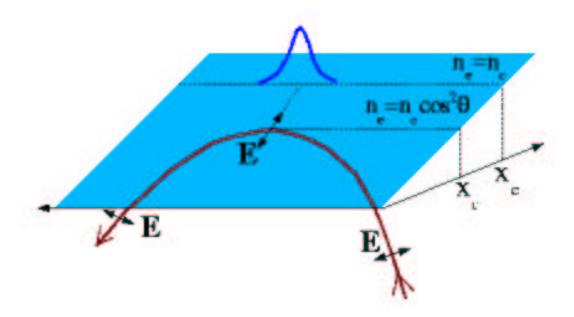
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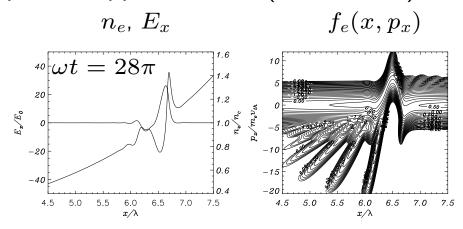
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In a warm plasma, the plasma oscillation propagates in the $n_e < n_c$ region and can accelerate electrons.

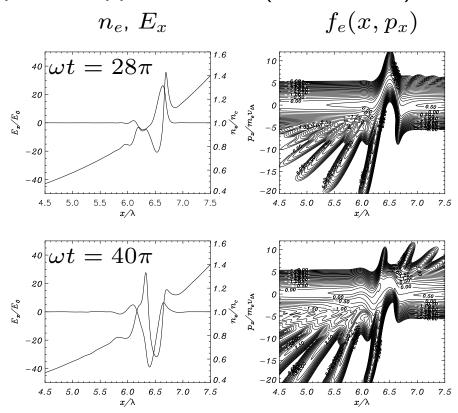
Numerical solution of the Vlasov-Poisson system within the capacitor approximation (uniform \mathbf{E}_d)

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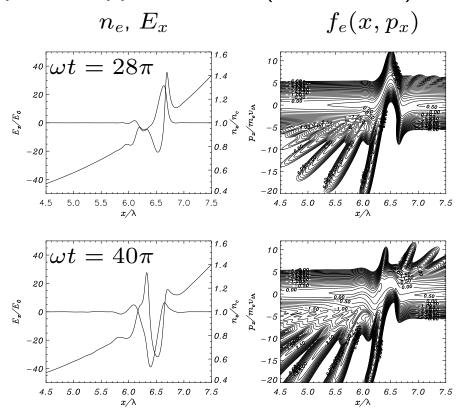
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A. Macchi and H. Ruhl, GSI Report, Darmstadt, 2000

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[E. S. Weibel, Phys. Fluids 10, 741 (1967)].

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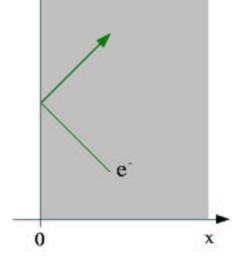
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Specular reflection at x = 0 is assumed:

$$f(x = 0, v_x, v_y) = f(x = 0, -v_x, v_y)$$



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Absorption degree A_{abs} and extinction length l_s in the limit $l_s \ll v_{th}/\omega$:

$$P_{abs} = \int_{-\infty}^{+\infty} dx \langle J_y E_y \rangle \equiv A_{abs}(c|E_L|^2/4\pi)$$

$$A_{abs} = \frac{8}{3\sqrt{3}} \left(\frac{v_{th} \omega_L^2}{c \omega_p^2} \right)^{1/3}, \quad l_s = \left(\frac{c^2 v_{th}}{\omega_p^2 \omega_L} \right)^{1/3}.$$

Anomalous skin effect absorption

Absorption degree A_{abs} and extinction length l_s in the limit $l_s \ll v_{th}/\omega$:

$$P_{abs} = \int_{-\infty}^{+\infty} dx \langle J_y E_y \rangle \equiv A_{abs}(c|E_L|^2/4\pi)$$

$$A_{abs} = \frac{8}{3\sqrt{3}} \left(\frac{v_{th} \omega_L^2}{c \omega_p^2} \right)^{1/3}, \quad l_s = \left(\frac{c^2 v_{th}}{\omega_p^2 \omega_L} \right)^{1/3}.$$

ASE + simple diffusion model for heat losses [i.e. $T_e = T_e(t)$] explains well absorption data by Price et al. in solid target at $I \le 10^{18}$ W cm⁻². [Rozmus et al, Phys. Plasmas **3**, 360 (1996)]

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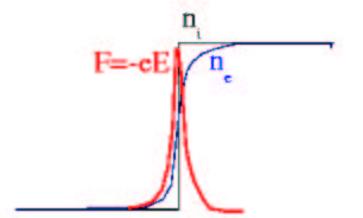
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In an external field E_d , electrons are reflected from the sheath if $v_{osc} = eE_d/m_e\omega < v_{th}$.

Since the sheath is very thin $(\approx \lambda_D)$, $E_s \sim \delta(x)$ may be assumed (reflecting boundary).

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At normal incidence, the $\mathbf{v} \times \mathbf{B}$ force may drive VH.

Simple electrostatic model of "Vacuum heating"

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4\pi j_x &= \partial_t E_e, \\
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$$x = x_0 + \xi(x_0, t) \,, \ \partial_t \xi = v_x$$

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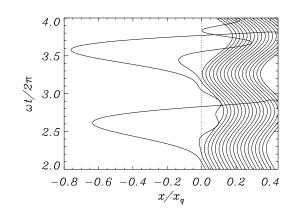
Electrons crossing the surface towards vacuum $(x_0 + \xi < 0)$ feel a discontinous force with an effective secular acceleration $\omega_p^2 x_0$.

Numerical solution of the equation of motion in Lagrangian coordinates:

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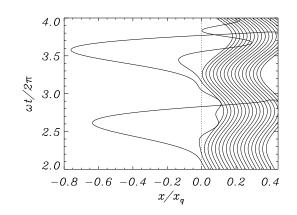
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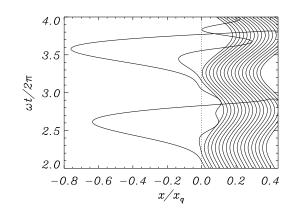
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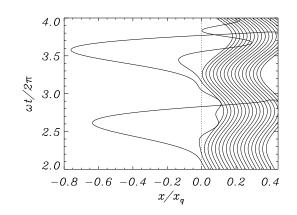
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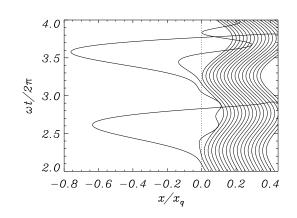
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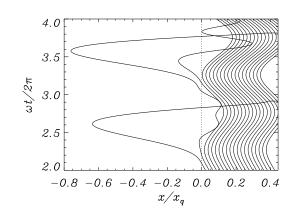
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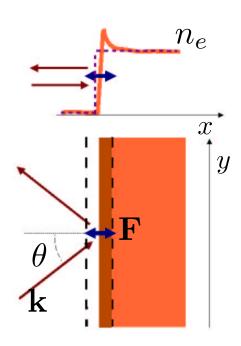
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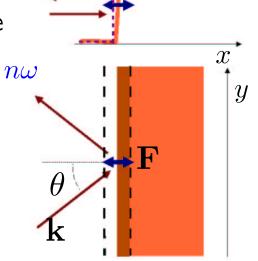
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 \rightarrow "moving mirror" effect: generation of high harmonics 3ω , 4ω , ..., $n\omega$...



 n_e

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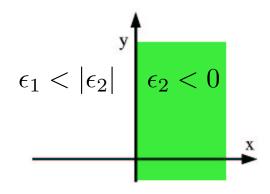
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- Magnetic collimation of "fast" electrons

• . . .

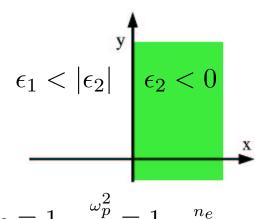


$$E_y = E_0 \left[\theta(-x)e^{q_-x} + \theta(x)e^{-q_+x} \right] e^{iky-i\omega t}$$

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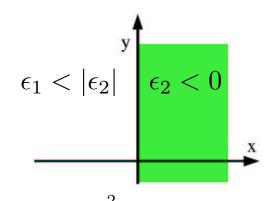
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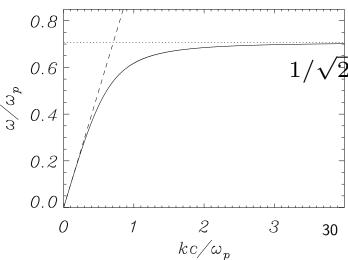
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$$k^2 = \frac{\omega^2}{c^2} \frac{\epsilon_1 |\epsilon_2|}{|\epsilon_2| - \epsilon_1} = \frac{\omega^2}{c^2} \frac{\omega_p^2 - \omega^2}{\omega_p^2 - 2\omega^2}$$



$$\epsilon_2 = 1 - \frac{\omega_p^2}{\omega^2} = 1 - \frac{n_e}{n_c}$$

$$< -1 \text{ if } n_0 > 2n_c$$



Linear mode conversion of the laser pulse into a SW at a plane vacuum–plasma interface requires $\omega_L = \omega_s$, $k_L \sin \theta = k_s$ where $k_L = \omega_L/c$ ($L \rightarrow$ laser, $s \rightarrow$ SW).

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Peak absorption occurs at optimal incidence angle $\sin \theta = \frac{k_s(\omega_L) + k_g}{\omega_L/c}$

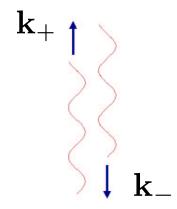
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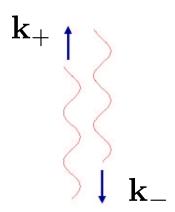
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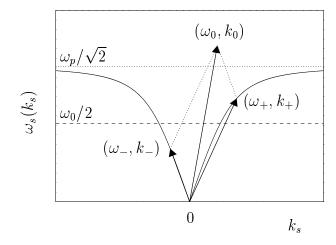


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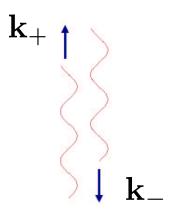


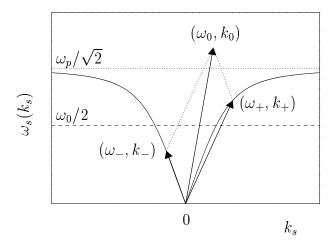
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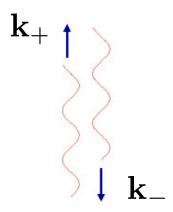
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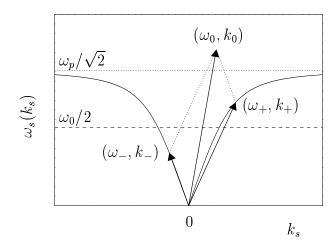
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However, also the $\mathbf{v} \times \mathbf{B}$ force at $2\omega_L$ may drive TSWD at normal incidence: $k_+ = -k_-$, $\omega_\pm = \omega_L$. [Macchi et al, PRL **87**, 205004 (2001); Phys. Plasmas **9**, 1704 (2002).]





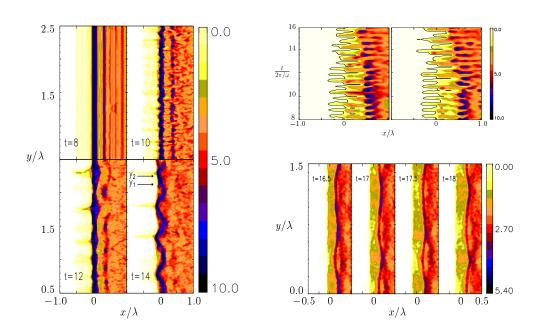
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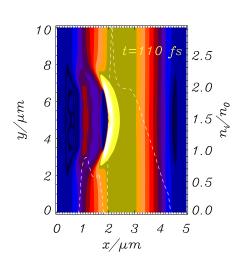
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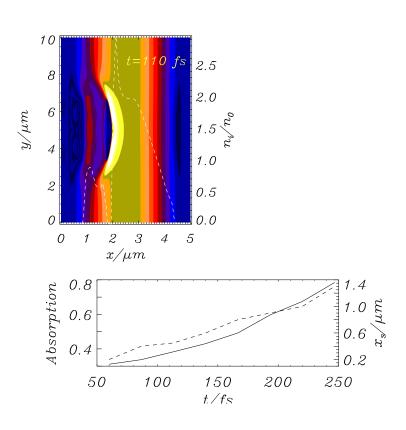
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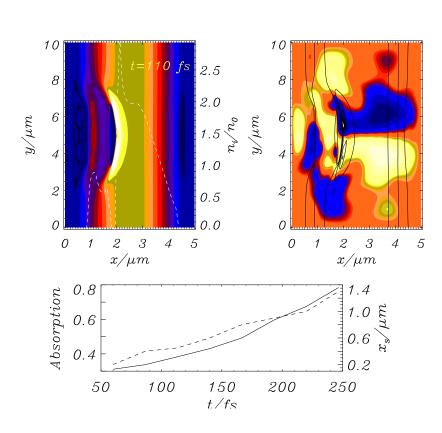
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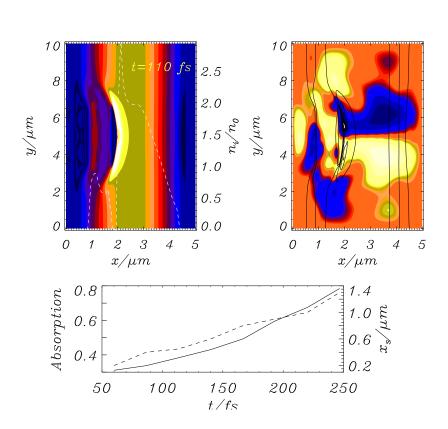
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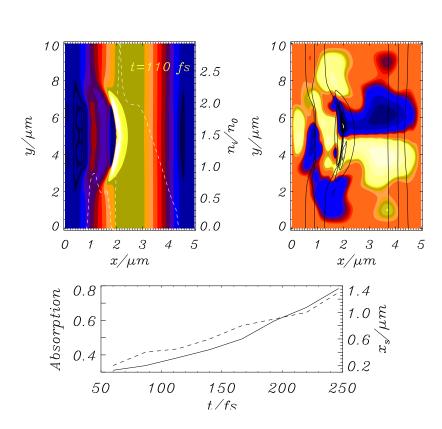
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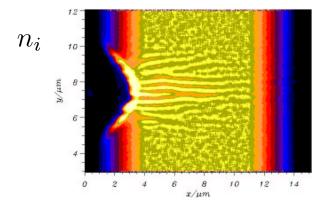


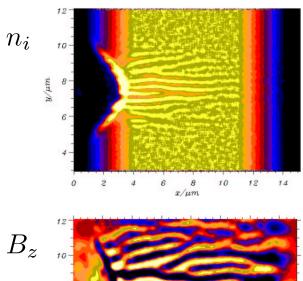
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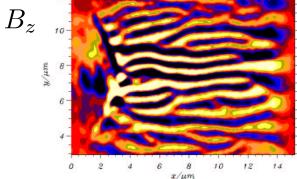
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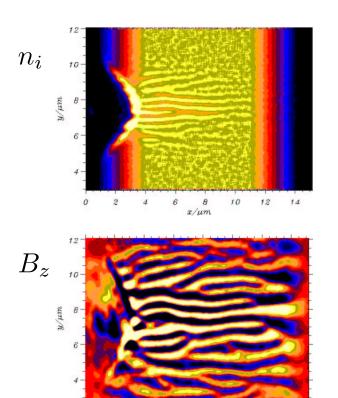
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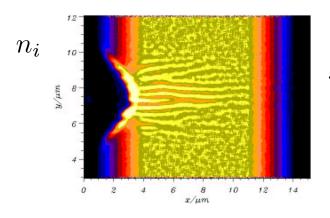




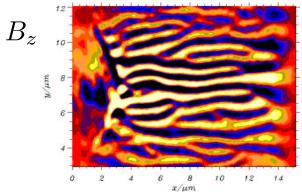


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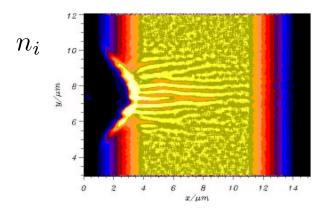
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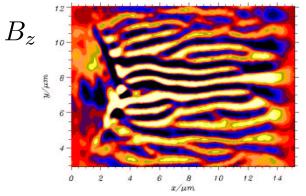
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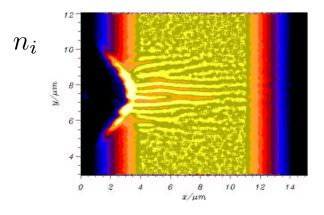
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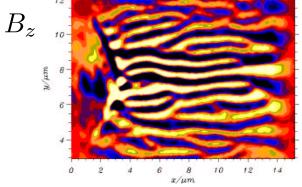
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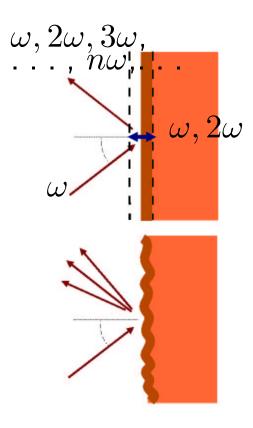
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spatial correlation with surface corrugations

PIC simulations by H. Ruhl, in Mulser et al., Las. Phys (1999).

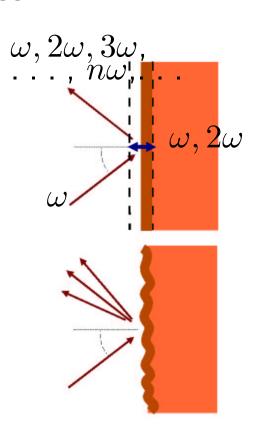
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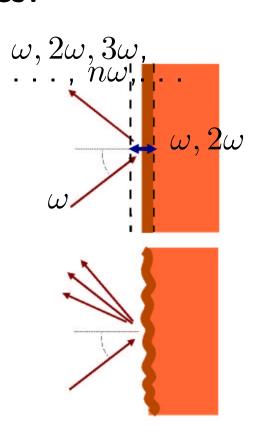
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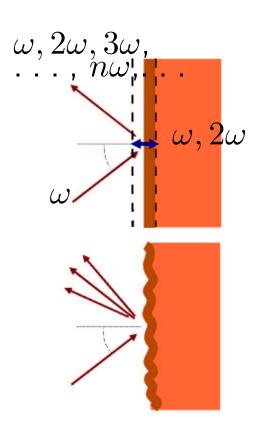
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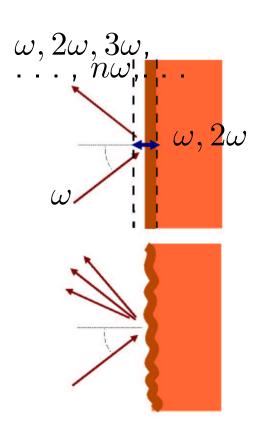


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The effect is detrimental to high harmonic generation from "moving mirrors".



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⇒ current neutralization by "background" electrons is needed to avoid "self-stopping" by associated electric and magnetic fields.

Current neutralization

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Neutralization of the "fast" electron current \mathbf{j}_f by a current \mathbf{j}_s of "slow" background electrons within a time:

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The equilibrium condition of opposite, neutralizing currents $\mathbf{j}_f = -\mathbf{j}_s$ is however affected by instabilities and additional effects.

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The field ${\bf E}$ has a slowing effect for fast electrons \Rightarrow collisions affect fast electron transport.

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$$\begin{split} n_f &= n_0 \exp(e\Phi/T_f) \\ &+ j_f = -j_s = -\sigma E \\ &+ \text{Poisson \& continuity eqs.} \\ \text{yield diffusion equation :} \\ \partial_t n_f &= \partial_x \left[\left(\frac{\sigma_s T_f}{n_f} \right) \partial_x n_f \right] \end{split}$$

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):
$$n_f(x,t) = \begin{cases} n_0 \left(\frac{t}{\tau_L}\right) \left(\frac{x_0}{x + x_0}\right)^2 & (t < \tau_L), \\ \frac{2n_0 x_0}{\pi} \frac{L(t)}{x^2 + L^2(t)} & (t > \tau_L). \end{cases}$$

$$n_0 = (2I_{abs}^2 \tau_L)/(9eT_f^3 \sigma_s),$$

$$x_0 = 3T_0^3 \sigma_s/I_{abs},$$

$$L(t) = x_0 \left[(t - \tau_L)(5\pi \sigma_s T_0)/(3en_0 x_0^2) + 1 \right]^{3/5}.$$

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$$\partial_t \mathbf{B} = \mathbf{\nabla} \times (\eta \mathbf{j}_f)$$
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Much additional physics is (or *should* be) inserted: target heating, slow electron diffusion, ionization, . . .

[Haines, PRL 47, 917 (1981).]

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Growth rate $\gamma \approx (2m_e/M_I)\nu_{eI}$, wavelength $\lambda \approx (M_I/m_e)^{1/2}\ell_{mfp}$.

(Ion mass appears because Ohmic dissipation is balanced by equipartition to ions.)

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The "Weibel" instability has been invoked to explain filamentation of currents observed in PIC simulations (moderate densities, relativistic electrons, collisions not important).

Simple model of transverse "Weibel" instability - I

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Linearized Vlasov+Maxwell equations

$$(-i\omega + ikv_y)f_1 = \frac{e}{m} \left[\left(E_x + \frac{v_y}{c} B_z \right) \partial_{v_x} - \frac{v_x}{c} B_z \partial_{v_y} \right] f_0,$$
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Dispersion relation $\omega = \omega(k)$ $(v_0^2 = v_1 v_2)$

$$\omega^{2} - k^{2}c^{2} = \omega_{p}^{2} \left(1 + \frac{v_{0}^{2}}{v_{te}^{2}} \int dv_{y} \frac{1}{\sqrt{\pi}v_{te}} e^{-v_{y}^{2}/v_{te}^{2}} \frac{kv_{y}}{\omega - kv_{y}} \right)$$

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Imaginary root $\omega = i\gamma, \gamma > 0$

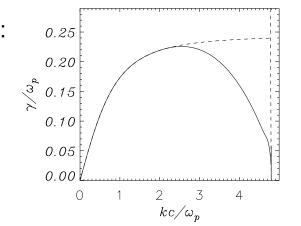
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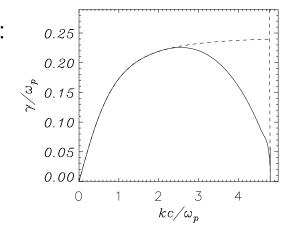
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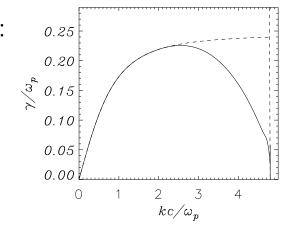


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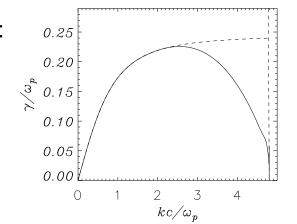
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Saturation when $B^2/8\pi \approx n_e m v_0^2/2$ (\approx energy equipartition) [Califano et al. PRE 1998]

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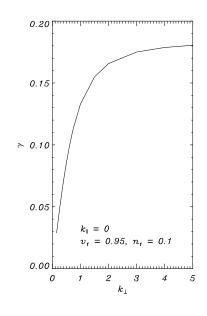
$$f(x, y, v_x, \mathbf{v}_\perp, t) = f_0(v_x, v_\perp^2) + f_1(v_x, \mathbf{v}_\perp) e^{i(k_\parallel x + ik_\perp \mathbf{r}_\perp - \omega t)}$$

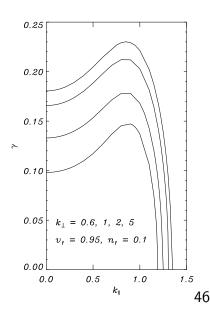
The most unstable wavevector $\mathbf{k}=(k_\parallel,\mathbf{k}_\perp)$ has $k_\perp\neq 0$ \rightarrow fields structures have finite length along the beams direction.

The distribution function (1) is also unstable with respect to electrostatic, longitudinal perturbations (two–stream instability). In the relativistic regime $(v_1 \to c)$ the longitudinal modes are coupled to the transverse "Weibel" modes for asymmetrical initial equilibria $(v_1 \neq v_2)$ [Califano et al. PRE 1998].

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The most unstable wavevector $\mathbf{k} = (k_{\parallel}, \mathbf{k}_{\perp})$ has $k_{\perp} \neq 0$ \rightarrow fields structures have finite length along the beams direction.





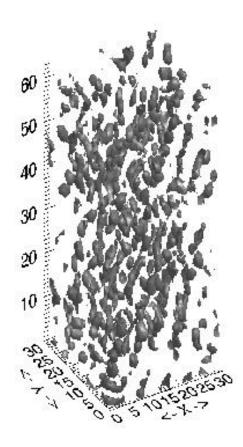
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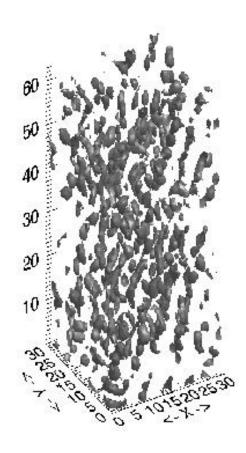
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3D "bubble–like" magnetic structures are formed with typical length scales $\sim d_e = c/\omega_p$. No extended filaments in beam direction are observed.

[Simulations by F. Califano;

Macchi et al, Nucl. Fus. 43, 362 (2003)]



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- → We performed test particle simulations of electron motion in the pump+SW fields involved in TSWD.

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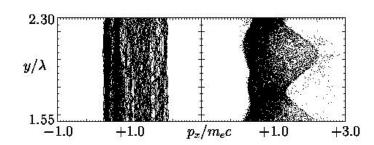
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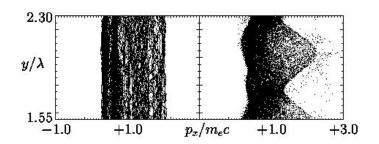
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Top: (y, p_x) phase space projections from PIC simulations at two subsequent times



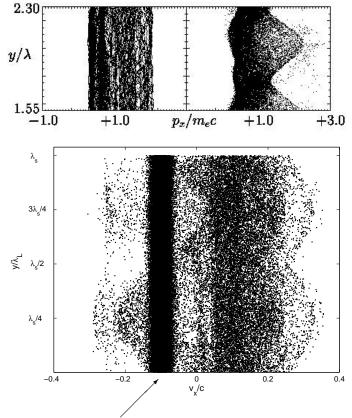
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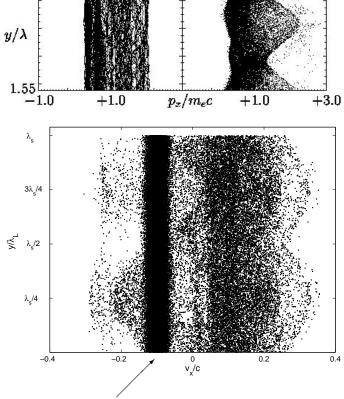
"Black stripe": initial conditions

2.30

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PIC and test-particle simulations both show enhanced electron heating near SW maxima



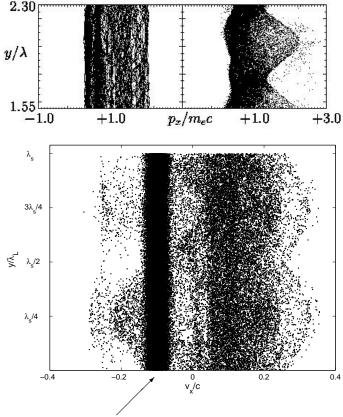
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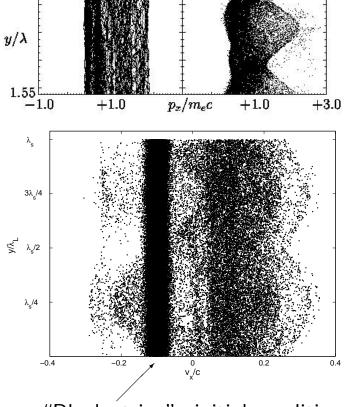
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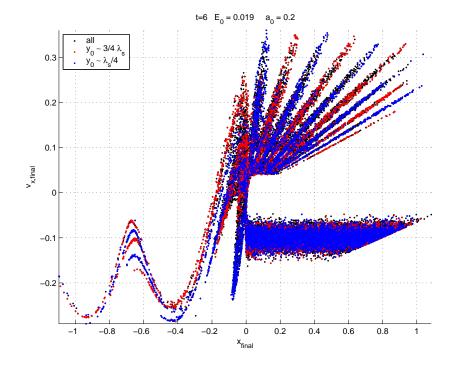
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Enhanced acceleration in time domain

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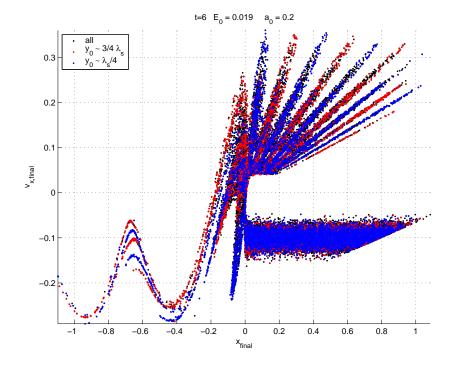
 (x, p_x) phase space

Black: all electrons in simulation



 (x, p_x) phase space

Black: all electrons in simulation Blue: electrons starting around $y=\lambda_s/4$



 (x, p_x) phase space

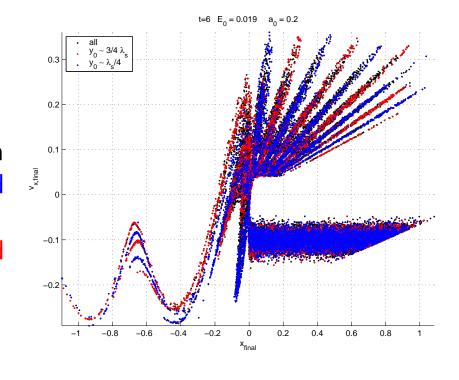
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Blue: electrons starting around

$$y = \lambda_s/4$$

Red: electrons starting around

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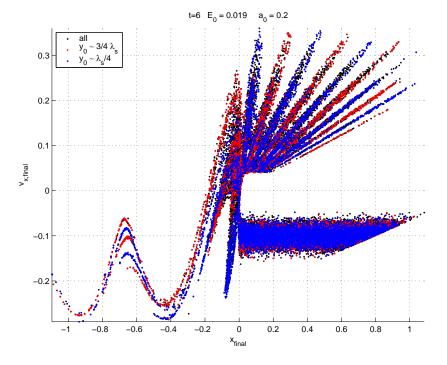
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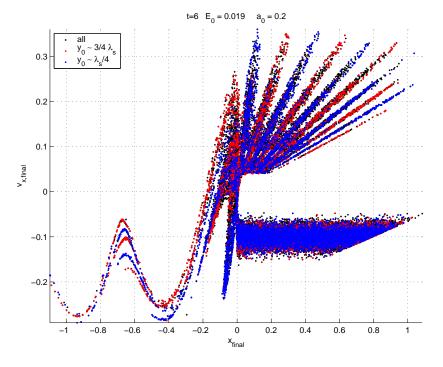
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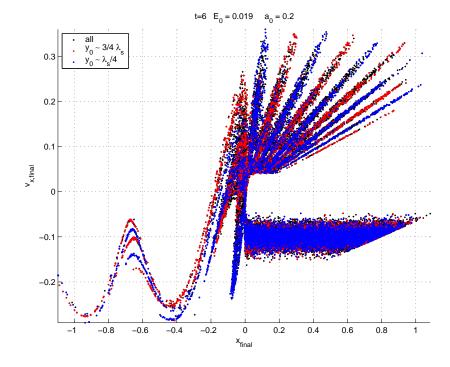


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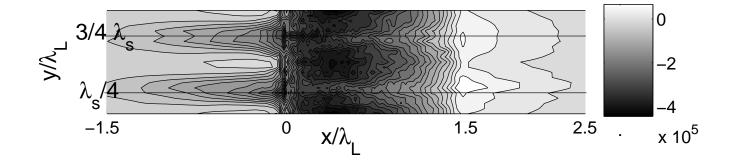
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- Near SW maxima some electrons are emitted into vacuum (x < 0) $(p_x \text{ modulated by } \mathbf{v} \times \mathbf{B} \sim \cos 2k_L x \text{ in vacuum})$

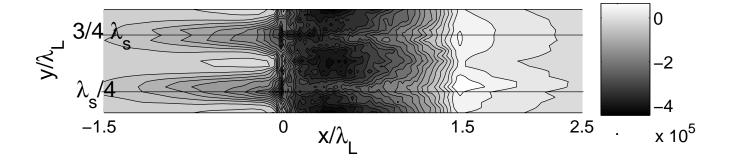
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Spatial imprint for current filamentation?

This talk was prepared on a Linux laptop,

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This was a 100% Microsoft-free presentation!

