Coherent Laser-Plasma Acceleration Examples and Recent Results

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Summary of This Talk

A short review of self-organization examples in the context of laser-plasma acceleration of ions, electrons and photons:

- "Light Sail"acceleration dynamics
- Light-pressure driven Rayleigh-Taylor instability
- Electron acceleration and XUV harmonic generation by surface plasma waves

Laser-Plasma Ion Acceleration: Artist's View ...



A. Macchi et al. Rev. Mod. Phys. 85 (2013) 571

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Laser-Plasma Ion Acceleration: Physicist's View ...

The "black box" hinders the acceleration mechanisms (not clear at time of discovery) The acceleration physics is of collective (cooperative, coherent) nature, based on self-consistent, non-linear plasma dynamics (complex and difficult to control)



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"Is plasma involved? It can't work" (Edward Teller on an early proposal of controlled fusion)



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The Vision of "Coherent" Acceleration (1957)

"The principles of coherent acceleration of charged particles"

V. I. Veksler, At. Energ. 2 (1957) 525



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- accelerating field on each particle proportional to the number of accelerated particles
- automatic synchrony between the particles and the accelerating field
- field localization in the region where the particles are
- acceleration of quasi-neutral bunches with large numbers of particles
- principles realized in laser-plasma acceleration of ions!

Example: Coherent "Radiation Drag" Acceleration

Equations of motion for a particle (radius $a \ll \lambda$) undergoing Thomson Scattering of a plane EM wave ($P_{sc} = \sigma_T I$)



For coherent scattering by a cluster with $N \gg 1$ particles

 $M \rightarrow NM$ $P_{sc} \rightarrow N^2 P_{sc}$ $\sigma_T \rightarrow N^2 \sigma_T$

 \longrightarrow *N*-fold increase in acceleration (Veksler, 1957)

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Light Sail Acceleration

EoM for a plane perfect mirror boosted by radiation pressure (2I/c in rest frame)is the same as for radiation drag of a particle Landau & Lifshitz, *The Classical Theory of Fields* ch.78

 $\frac{\mathrm{d}X}{\mathrm{d}t} = \beta c$ $\frac{\mathrm{d}(\gamma\beta)}{\mathrm{d}t} = \frac{2}{\rho\ell c^2} I\left(t - \frac{X}{c}\right) \frac{1 - \beta}{1 + \beta}$ E×B (104) а Analytical simulation eflecter 20 solution 4 analvtic $\propto t^{1/3}$ E_i (GeV) Sev 3 "observed" 00 2 laser in simulations -20 of thin foil 40 60 x/λ 20 40 60 $\omega t/2\pi$ acceleration at ~ 10^{23} Wcm⁻² 0 Esirkepov et al, PRL 92 (2004) 175003

 $V = \beta c$

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Beyond the Mirror: Charge Separation Effects

Real targets are not perfect rigid mirrors: local light pressure separates charges until electrostatic tension balances $P_{\rm L} = 2I/c$ Space-charge field E_x accelerates and bunches ions in the skin layer ($x_d < 0 < x_s$) until hydrodynamic "breaking" at $x = x_s$, $t = t_b$ produces a ion bunch with velocity v_b



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Ion Bunch Modeling and Dynamics



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From Ion Bunch to Light Sail

With proper choice of thickness a single ion bunch can be produced and re-accelerated as laser front advances Macchi et al, PRL 103 (2009) 85003; New J. Phys. 12 (2010) 045013

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Light Sail motion emerges as the multiple average over re-acceleration stages Grech et al, New J. Phys. 13 (2011) 123003]



Multispecies Effects

- For coherent light sail motion all species move with same V
- \rightarrow ideally same energy/nucleon for each Z/A
 - self-organized acceleration is complicated by multiple ion species (e.g. no simple self-similar motion & bunching)
 - LS proton acceleration requires tight control of target thickness and species spatial distribution (experimentally challenging)

[see e.g. Macchi et al PRL 103 (2009) 85003;

Qiao et al PRL 105 (2010) 155002]

 Single species target contain hydrogen surface impurities: removal necessary (but not straightforward) to prevent multispecies dynamics

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Impurity Removal for Carbon Ion Acceleration

Exploitation of picosecond prepulse for all-optical impurity removal

 \rightarrow for 15 nm thickness the energy/nucleon is higher for C⁶⁺ than for protons (H⁺)





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Possible Biomedical Application?



REVIEW published: 17 January 2020 doi: 10.3389/fonc.2019.01563

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Ultra-High Dose Rate (FLASH) Radiotherapy: Silver Bullet or Fool's Gold?

Joseph D. Wilson^{1†}, Ester M. Hammond^{1†}, Geoff S. Higgins^{1†} and Kristoffer Petersson^{1,2*†}

¹ Department of Oncology, The Oxford Institute for Radiation Oncology, University of Oxford, Oxford, United Kingdom, ² Radiation Physics, Department of Haematology, Oncology and Radiation Physics, Skåne University Hospital, Lund, Sweden

For Carbon ions only laser-driven accelerators may deliver the necessary flux in an ultrashort time

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Public Coverage (FLASH-Stimulated?)

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PARTICLE THERAPY | RESEARCH UPDATE

Intense radiation pressure enables selective acceleration of carbon ion beams

Physics Today 75, 1, 19 (2022); https://doi.org/10.1063/PT.3.4916

Irish boffins' laser to help beat cancer

Physics

particle therapy

A laser selectively kicks carbon out of a foil

Experiments and simulations show how the shape of a laser's profile determines which target atoms make up the resulting ion beam.

A New Trick to Make Short-Pulse Ion Beams

A new laser technique could lead to ultrashort-pulse, high-energy ion beams for medical use.



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Relativistic High Gain 3D Regime

For "extremely" tightly focused & intense pulses $\rightarrow \gamma \gtrsim 1$ (GeV energies)

- Transverse expansion of the target reduces surface density $\rho\ell$

- Decrease of laser frequency in "sail" frame delays the transparency onset

 \rightarrow enhanced self-regulated acceleration (at the expense of the number of ions)

Analytical model shows faster & higher gain in 3D than in 1D:

$$\gamma(t) = \left(\frac{t}{\tau_k}\right)^k \qquad k = \frac{D}{D+2} = \begin{cases} 1/3 & (1D) \\ 2/4 & (2D) \\ 3/5 & (3D) \end{cases}$$

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S.V.Bulanov et al. PRL 104 (2010) 135003

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Fast 3D Scaling in Simulations

Laser: 24 fs, 4.8 μ m spot, $I = 0.85 \times 10^{23}$ W cm⁻² \implies 1.5 kJ Target: $d = 1 \ \mu$ m foil, $n_e = 10^{23}$ cm⁻³



 $\mathscr{E}_{max} \simeq 2.6 \text{ GeV} > 4 \text{ X 1D}$ prediction (still limited by transparency) Sgattoni et al, Appl. Phys. Lett. **105** (2014) 084105

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Light Sail Rayleigh-Taylor Instability



3D light sail simulation: formation of net-like structures with size ~ λ (laser wavelength) and ~ hexagonal shape



Interpretation: Rayleigh-Taylor instability driven by light pressure A. Sgattoni et al. Phys. Rev. E **91** (2015) 013106 see also: E. Ott, PRL **29** (1972) 1429; F. Pegoraro & S.V.Bulanov, PRL **99** (2007) 065002; B. Eliasson, New. J. Phys. **17** (2015) 033026

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Rayleigh-Taylor Instability in Space and Lab



Heavy fluid over a light fluid is unstable († gravity ↓ acceleration)





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Exagon formation in RTI: an example of "spontaneous symmetry breaking" S.I.Abarzhi, PRE **59** (1999) 1729

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Plasmonic Imprint on Light-Driven RTI

The EM field at a rippled surface of spatial period *d* is modulated The *P*-component is resonantly enhanced when $d \sim \lambda$ due to the excitation of surface plasmons



Electron Acceleration in Plasma Waves

Much like surfers electrons can gain energy by longitudinal plasma waves when "injected" close to the phase velocity $v_p = \omega/k$ (provided $v_p < c$)

Plasma-based acceleration of electrons exploits wake waves at $v_p \lesssim c$ driven either by short laser pulses or charged particle bunches [see e.g. A.Macchi,

"A Superintense Laser-Plasma Interaction Theory Primer" (Springer, 2018), chap.4;

Am. J. Phys. 88 (2020) 723]



T. Katsouleas, Nature 444 (2006) 688



Electron Acceleration in Surface Plasma Waves

SPW can be excited at a sharp laser-plasma interface (by using "grating" targets, shooting on edge or at grazing incidence: see e.g. J. Sarma et al. New J. Phys. **24** (2022) 073023)



electrons are accelerated along the surface by "surfing" the SPW

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Observation of "Surfing" Electrons





Experiment at SLIC, CEA Saclay (France) Collimated ($\simeq 20^{\circ}$ cone) multi-MeV electrons near the surface tangent ($\phi \simeq 2^{\circ}$) Large total charge up to $\simeq 650$ pC Efficient self-injection in the SPW! L.Fedeli et al. PRL **116** (2016) 015001 G.Cantono et al. PoP **25** (2018) 031906



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Boosting of HH by Electron Nanobunching

Electrons (\rightarrow) trapped and accelerated by the SP self-organize into short bunches

Coherent scattering of the laser field by the electron bunches produce bright quasi-collinear HH

(similar to collective instability operation in a Free Electron Laser) 2D simulations by L. Fedeli



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2D simulations by L. Fedeli

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References to Our Work (1/2)

- A.Macchi, S.Veghini, F.Pegoraro, Light Sail Acceleration Reexamined, Phys. Rev. Lett. 103 (2009) 085003
- A.Macchi, S.Veghini, T.V.Liseykina, F.Pegoraro, Radiation Pressure Acceleration of Ultrathin Foils, New J. Phys. 12 (2010) 045013
- A.Sgattoni, S.Sinigardi, L. Fedeli, F. Pegoraro, A.Macchi, Laser-driven Rayleigh-Taylor instability: Plasmonic effects and three-dimensional structures, Phys. Rev. E 91 (2015) 013106
- A. McIlvenny, D. Doria, L. Romagnani, H. Ahmed, N. Booth, E. J. Ditter, O. C. Ettlinger, G. S. Hicks, P. Martin, G. G. Scott, S. D. R. Williamson, A. Macchi, P. McKenna, Z. Najmudin, D. Neely, S. Kar. M. Borghesi, *Selective Ion Acceleration by Intense Radiation Pressure*, Physical Review Letters **127** (2021) 194801

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References to Our Work (2/2)

L. Fedeli, A. Sgattoni, G. Cantono, D. Garzella, F. Réau, I. Prencipe, M. Passoni, M. Raynaud, M. Květoň, J. Proska, A. Macchi, T. Ceccotti, *Electron acceleration by relativistic surface plasmons in laser-grating interaction*, *Phys. Roy. Lett.* **116** (2016) 015001

Phys. Rev. Lett. 116 (2016) 015001

 G. Cantono, L. Fedeli, A. Sgattoni, A. Denoeud, L. Chopineau, F. Reau, T. Ceccotti, A. Macchi, *Extreme ultraviolet beam enhancement by relativistic surface plasmons*, Phys. Rev. Lett. **120** (2018) 264803

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 J. Sarma, A. McIlvenny, N. Das, M. Borghesi, A. Macchi, Surface Plasmon-Driven Electron and Proton Acceleration without Grating Coupling,

New J. Phys. 24 (2022) 073023

THANKS FOR WATCHING!

Download this talk: osiris.df.unipi.it/~macchi/TALKS/

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Heating Suppression for Light Pressure Dominance



The "Optical Mill" (Crookes radiometer) rotates in the *opposite* way to that suggested by the imbalance of light pressure between white (reflecting) and black (absorbing) plates

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(white: $R \simeq 1$, black: $R \simeq 0$)

Thermal pressure dominates due to stronger heating of the black plate in the imperfect vacuum of bulb In laser-driven light sail acceleration heating must be suppressed for efficiency at not-so-extreme intensities $(I = 10^{18} - 10^{21} \text{ Wcm}^{-2})$

Circular Polarization Quenches Heating

Normal incidence & linear polarization (LP): electrons perform non-adiabatic oscillations driven by the $\mathbf{v}\times\mathbf{B}$ force across the sharp density gradient

Oscillating (2ω) term vanishes for circular polarization (CP)

CP

→ suppression of electron heating

Light pressure (0ω) term) dominates & ions respond smoothly to steady force Macchi et al. PRL **95** (2005) 185003

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LP

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Energy spectrum improvement for circular polarization

GEMINI laser, $\tau_p = 45$ fs, $I = 6 \times 10^{20}$ W cm⁻², $\frac{10^{10}}{3}$ 15 nm thick CH foils, CP vs LP experiment: CP brings higher energies & spectral peaks for both species 10^{8} simulations: transparency-limited acceleration, proton spectrum not well reproduced (c) 10¹¹ circular polarization linear polarization t=66fst=83fs t=33fs t=50fs 10^{10} **Ω**P/**3**P/NP e 4 0 108 -4 0.1 10 IB_zI₄ 30 15 m TOY -4 ò ò 4 ò 4

x [µm]

(b) Proton Noise 10 1520 30 $\epsilon_{\rm Proton}$ (MeV) Noise 102025 $\epsilon_{\rm Carbon}$ (MeV/u) C.Scullion et al, PRL 119 (2017) 054801 ヘロト ヘヨト ヘヨト ヘヨ

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Light Sail energy from conservation laws

Conservation of 4-momenta in "collision" between laser pulse and moving mirror (mass $M = \rho \ell$)

 $p_i + mc = p_r + \mathcal{E}/c$ $p_i = -p_r + p_s$

Image: A matrix

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Using
$$\mathscr{E}^2 = M^2 c^2 + p_s^2$$
 and $p_i = \int_0^\infty \frac{I(t')}{c} dt' \equiv \frac{Mc}{2} \mathscr{F}$
energy $\frac{\mathscr{E}}{Mc^2} = \frac{\mathscr{F}^2}{2(\mathscr{F}+1)} \left(\simeq \frac{\mathscr{F}^2}{2} \text{ for } \beta = \frac{p_s c}{\mathscr{E}} \ll 1 \right)$
efficiency $\eta = \frac{\mathscr{E}}{p_i c} = \frac{2\beta}{1+\beta} \longrightarrow 100\%$ in the $\beta \to 1$ limit

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Coherent Laser-Plasma Acceleration

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Foreseen laser sailing ...

R.Forward (1964) G.Marx (1966)









Breakthrough Starshot (2016) breakthroughinitiatives.org Critical analysis of (un)feasibility: H.Milchberg, Phys. Today, 26 April 2016

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Light Sail with extreme light on nanofoils

Energy/nucleon & efficiency from the 1D mirror model (τ_p : laser pulse duration)

$$\mathscr{F} = \frac{2I\tau_p}{\rho\ell} = \frac{Z}{A} \frac{m_e}{m_p} \frac{a_0^2}{\zeta} \omega \tau_p \qquad \left(\zeta = \pi \frac{n_e}{n_c} \frac{\ell}{\lambda}\right)$$
Deptimal thickness $a_0 \simeq \zeta$ at threshold of
elativistic transparency
$$\mathscr{E}_{\max} \simeq 2\pi^2 \frac{(m_e c)^2}{m_p} \left(\frac{Z}{A} \frac{c\tau_p}{\lambda} a_0\right)^2$$
 $\mathscr{E} \simeq 10 \text{ nm}, I \simeq 1.6 \times 10^{21} \text{ W cm}^{-2} (a_0 = 22), \tau_p = 40 \text{ fs}$
 $\longrightarrow \mathscr{E}_{\max} \simeq 150 \text{ MeV}, \eta \simeq 50\%$
Coherent motion of the sail \longrightarrow mononergetic ion spectrum
A dream ion beam? ...

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Transparency of ultrathin plasma foil 1D model with relativistic nonlinearity

 $n_e(x) \simeq n_0 \ell \delta(x)$ (ℓ : foil thickness)

[V.A.Vshivkov et al, Phys. Plasmas 5 (1996) 2727] Nonlinear reflectivity:

$$R \simeq \begin{cases} 1 & (a_0 < \zeta) \\ \frac{\zeta^2}{a_0^2} & (a_0 > \zeta) \end{cases} \qquad \zeta \equiv \pi \frac{n_0 \ell}{n_c \lambda}$$

The transparency threshold $a_0 \simeq \zeta$ depends on areal density $n_0 \ell$ (note: it is *not* $n_e < n_c \gamma$ with $\gamma = (1 + a_0^2)^{1/2}$)



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Analytical model for multi-D Light Sail

Self-similar transverse dilatation $r_{\perp}(t) = \Lambda(t)r_{\perp}(0)$

$$\sigma = \sigma(t) = \frac{\sigma(0)}{\Lambda^{D-1}(t)}, \quad \frac{\mathrm{d}}{\mathrm{d}t}(\gamma\beta_{\parallel}) = \frac{2I}{\sigma(0)c^2}\Lambda^{D-1}(t)\frac{1-\beta_{\parallel}}{1+\beta_{\parallel}} \quad (D = 1, 2, 3)$$

Impulsive transverse kick by ponderomotive force

$$\frac{\mathrm{d}p_{\perp}(t)}{\mathrm{d}t} \simeq -m_e c^2 \partial_r (1 + a^2(r, t))^{1/2} \simeq 2m_e c^2 a_0 r / w \qquad (a_0 \gg 1, r \ll w)$$

→ transverse momentum scales linearly with position

$$\frac{\mathrm{d}\Lambda}{\mathrm{d}t} = \frac{\dot{r}_{\perp}(t)}{\dot{r}_{\perp}(0)} = \frac{\alpha}{\gamma(t)} , \qquad \gamma(t) \simeq (p_{\parallel}^2 + m_i^2 c^2)^{1/2} , \qquad \alpha \simeq 2 \frac{m_e a_0 c^2 \Delta t}{m_p w^2}$$

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Solution in the $\gamma \gg 1$ limit $\gamma = \left(\frac{t}{\tau_k}\right)^k$, $k = \frac{D}{D+2}$

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Thin foil RTI with self-consistent pressure modulation

Model: reflection from shallow 2D grating of depth δ + modified Ott's theory [PRL **29** (1972) 1429] with modulated pressure:

$$-(q^2 - k^2)^{1/2}$$
(S)

$$P \simeq P_0(1+K(q)\delta\cos qy), \qquad K(q) = \begin{cases} k^2 q(q^2-k^2)^{-1/2} & (P) \\ (k^2-q^2/2)(q^2-k^2)^{-1/2} & (C) \end{cases}$$



S-polarization P-polarization C-ircular polarization RT: no modulation ($\delta = 0$)

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