Advanced schemes for laser-plasma ion acceleration

Andrea Macchi

National Institute of Optics, National Research Council (CNR/INO), Adriano Gozzini research unit, Pisa, Italy

Enrico Fermi Department of Physics, University of Pisa, Italy



Chalmers University of Technology, Göteborg, Sweden, November 20, 2014

< E > < E >

Andrea Macchi

CNR/INO

Main coworkers for this talk

- L. Fedeli¹, A. Grassi^{1,2}, A. Sgattoni, A. Singh Nindrayog^{1,†},
- S. Sinigardi^{*}, F. Pegoraro¹

CNR/INO, Pisa, Italy ¹Dipartimento di Fisica "Enrico Fermi", Università di Pisa, Italy ²Université Pierre et Marie Curie, Paris and LULI, Ecole Polytechnique, France [†]presently at Department of Computer Science, Lovely Professional University, Jalandhar, India ^{*}presently at Dipartimento di Fisica, Università di Bologna, Bologna, Italy

... and many colleagues from several institutes abroad (Queen's University – Belfast, LULI Ecole Polytechnique – Palaiseau, CEA/LyDyL – Saclay, Technical University –Prague, ...)

Outline

- Some history: from the coherent (collective) acceleration paradigm (1957) to the (re–)discovery of laser-driven proton beams (2000)
- Acceleration mechanisms: theory & experiment
- Target Normal Sheath Acceleration (TNSA): use of structured targets and high-field plasmonics
- Radiation Pressure Acceleration (RPA): "Light Sail" acceleration, high gai regimes and instability
- Collisionless Shock Acceleration (CSA): conditions for mononergetic spectra, Vlasov simulations

・ロト ・回ト ・ヨト ・ヨト

Recent ion acceleration reviews (parochial selection)

A. Macchi, M. Borghesi, M. Passoni, Ion Acceleration by Superintense Laser-Plasma Interaction, Rev. Mod. Phys. **85** (2013) 571

A. Macchi, A. Sgattoni, S. Sinigardi, M. Borghesi, M. Passoni, Advanced Strategies for Ion Acceleration using High Power Lasers,

Plasma Phys. Contr. Fus. 55 (2013) 124020

A. Macchi, A Superintense Laser-Plasma Interaction Theory Primer (Springer, 2013) Chap.5 "Ion Acceleration" (for absolute beginners)



CNR/INO

イロン イヨン イヨン イヨン

Andrea Macchi

Other ion acceleration reviews

H. Daido, M. Nishiuchi, A. S. Pirozhkov, *Review of Laser-Driven Ion Sources and Their applications*, Rep. Prog. Phys. **75** (2012) 056401

J.C. Fernández, B.J. Albright, F.N. Beg, M.E. Foord, B.M. Hegelich, J.J. Honrubia, M. Roth, R.B. Stephens, and L. Yin, *Fast ignition with laser-driven proton and ion beams*, Nucl. Fusion **54** (2014) 054006

< ロ > < 同 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

The vision of "coherent" acceleration: Veksler (1957)

V. I. Veksler, At. Energ. 2 (1957) 525



< ロ > < 同 > < 回 > < 回 >

- accelerating field on each particle proportional to the number of accelerated particles
- automatic synchrony between the particles and the accelerating field
- field localization in the region where the particles are
- acceleration of quasi-neutral bunches with large numbers of particles

The dawn of laser-plasma physics (1964)

"The laser is a solution looking for a problem" (D'Haenens to Maiman, 1960) Q-switched lasers (1962): 10 GW on ~ 10^{-2} cm spot $\rightarrow I \simeq 10^{13}$ W cm⁻²



イロン イヨン イヨン イヨン

THE PHYSICS OF FLUIDS

VOLUME 7, NUMBER 7

JULY 1964

CNR/INO

On the Production of Plasma by Giant Pulse Lasers

John M. Dawson

Plasma Physics Laboratory, Princeton University, Princeton, New Jersey (Received 10 October 1963; final manuscript received 10 March 1964)

Calculations are presented which show that a laser pulse delivering powers of the order of 10¹⁰ W to a liquid or solid particle with dimensions of the order of 10^{-2} cm will produce a hot plasma with temperatures in the range of several hundred eV. To a large extent the plasma temperature is held down by its rapid expansion and cooling. This converts much of the energy supplied into ordered energy of expansion. This ordered expansion energy can amount to several keV per ion. If the expanding plasma can be caught in a magnetic field and its ordered motion converted to random motion this might be utilized as a means for filling controlled thermonuclear fusion devices with hot plasma. Further, it should also be possible to do many interesting plasma experiments on such plasmas.

Andrea Macchi

Modern ultraintense laser-matter interactions

Short (~ 10 fs = 10^{-14} s) pulses of Petawatt (10^{15} W) power focused near diffraction limit ($w \sim 1 \mu$ m): $I \simeq 10^{22}$ W cm⁻² (Hercules Laser, CUOS, University of Michigan Ann Arbor, US)

Proposed ELI laser: 100 PW, 15 fs, $I > 10^{23}$ W cm⁻² A future vision: multi-fibre laser [Mourou et al, Nature Photonics **7** (2013) 258]



Figure 1) Principle of a coherent amplifier network. An initial puble from a seed laser (1) is stretched (2), and split into many fibre channels (3). Each channel is amplified in several stages, with the final stages producing publes of -1 m1 at high repetition rate (4). All the channels are combined coherently, compressed (5) and focused (6) to produce a puble with an energy of >10 J at a repetition rate of -10 kHz (7).



CNR/INO

Andrea Macchi

Multi-MeV protons from solid targets (2000)



Andrea Macchi

CNR/INO

Multi-MeV protons from solid targets (2000)

Up to 58 MeV protons observed at LLNL (Livermore) Petawatt Snavely et al, PRL **85** (2000) 2945 Other observations:

Clark et al, PRL **84** (2000) 670 Maksimchuk et al, PRL **84** (2000) 4108

Protons have been observed and characterized in a large number of laboratories and for different laser pulse regimes Figure from Borghesi et al, Plasma Phys. Contr. Fus. **50** (2008) 124040



Andrea Macchi

Target Normal Sheath Acceleration (TNSA)

Physics: sheath field generation by "fast" relativistic electrons at the rear surface of a solid target Field lifetime:

$$\sim (1-10) \times 10^{-12} \text{ s}$$

→ ultrashort ion bunches

Protons originate from a surface impurity layer at the target rear: favorable initial position and Z/A ratio (target cleaning \rightarrow heavier ions acceleration)





イロト イヨト イヨト イヨト

Fast electron generation: a simplified picture

Needs laser-driven push-pull of electrons across the density gradient





Electrons perform "half-oscillations" in vacuum and re-enter in the plasma with approximately the "quiver" energy

Oscillations driven by:

- E for P-polarization
- $\mathbf{v} \times \mathbf{B}$ for *S*-polarization or normal incidence

Image: A matrix

Andrea Macchi

Fast electron generation: 1D simulations



Linear Polarization: fast electron bunches at rate ω (for $\theta = 30^{\circ}$, *P*-pol.) or 2ω (for $\theta = 0^{\circ}$)

Circular Polarization at $\theta = 0^{\circ}$: *no fast electrons* (($\mathbf{v} \times \mathbf{B}$)_{2 ω} = 0)

CNR/INO

Andrea Macchi

Fast electron generation: typical parameters

Typical energy ("ponderomotive scaling")

$$\mathcal{E}_f = m_e c^2 (\gamma - 1) \sim m_e c^2 \left(\sqrt{1 + a_0^2 / 2} - 1 \right)$$

*a*₀: "relativistic" amplitude parameter

$$a_0 = \left(\frac{I\lambda^2}{10^{18} \text{ W/cm}^2}\right)^{1/2} = \frac{eE_L}{m_e\omega c} = \frac{p_{\text{osc}}}{m_e c}$$

- conversion efficiency $\eta_f \simeq 10^{-2} 10^{-1}$
- density $n_f \simeq 10^{20} 10^{21} \text{ cm}^{-3}$
- current density $\sim 10^{12} \text{ A/cm}^2 \rightarrow 10 \text{ MA}$ over the laser spot

Andrea Macchi

CNR/INO

Static modeling of TNSA

Assume fast electrons in Boltzmann equilibrium with density n_e and temperature T_e as the only parameters to evaluate sheath extension L_s and potential drop $\Delta \Phi$



Looking for resonant coupling at the surface

Idea: enhancement of the surface field and of absorption by exciting a normal mode of the target plasma Resonant coupling requires matching of the phase $\varphi = \mathbf{k}_{\parallel} \cdot \mathbf{r} - \omega t$ between the laser -----and the resonant mode characterized by θ k

$$\omega \doteq \omega_m \qquad k_{\parallel} = k \cos \theta \doteq (k_m)_{\parallel}$$

Normal modes of step boundary metal/plasma: surface waves

```
SW: building blocks of plasmonics (at low fields)
```



ヘロト ヘヨト ヘヨト ヘヨト

CNR/INO

Andrea Macchi

Surface wave coupling: the matching problem

Plasma-vacuum interface

$$\varepsilon_1 = 1$$
 $\varepsilon_2 = 1 - \frac{\omega_p^2}{\omega^2} = 1 - \frac{n_e}{n_c}$

Dispersion relation $\omega = \omega(k)$

$$k = \frac{\omega}{c} \left(\frac{\omega_p^2 - \omega^2}{\omega_p^2 - 2\omega^2} \right)^{1/2}$$

$$\omega < \omega_p / \sqrt{2}$$
 $v_{\rm ph} < c$

No matching with $\omega = c k_{\parallel}$



CNR/INO

Andrea Macchi

Surface wave matching in periodic structures

In a spatially periodic medium with period *d*, folding of $\omega_{SW}(k)$ in the Brillouin zone $|k| < \pi/d$ (Floquet-Bloch theorem) allows phase matching

Figure: M.Lupetti, M.Sc. Thesis, 2011

Resonant coupling with EM _____ wave is possible in a grating at an angle of incidence

$$\sin\theta_{\rm res} + \lambda/d = \left(\frac{1 - \omega_p^2/\omega^2}{2 - \omega_p^2/\omega^2}\right)^{1/2}$$

(provided $\omega_{SW}(k)$ does not change much)





CNR/INO

Andrea Macchi

Plasmonics at high fields?

Aim: exploit plasmonic field enhancement and surface wave excitation in sub-wavelength (periodic) structures at high intensity (> 10^{18} W cm⁻², relativistic regime)

- hydrodynamics may wash target structuring out
- use ultrashort pulses (10s of fs)
- prepulse effects may destroy structures at target surface
- use high-contrast techniques (e.g. plasma mirrors)
- features of plasmonics and surface waves unknown in the high-field, nonlinear, relativistic regimes

・ロト ・回ト ・ヨト ・ヨト

CNR/INO

hope for the best ...

Experimental evidence: grating-enhanced TNSA

PRL 111, 185001 (2013)

PHYSICAL REVIEW LETTERS

week ending 1 NOVEMBER 2013

CNR/INO

イロト イヨト イヨト イヨト

Evidence of Resonant Surface-Wave Excitation in the Relativistic Regime through Measurements of Proton Acceleration from Grating Targets

T. Ceccotti,^{1,a} V. Floquet,¹ A. Sgattoni,^{2,3} A. Bigongiari,⁴ O. Klimo,^{5,6} M. Raynaud,⁷ C. Riconda,⁴ A. Heron,⁸ F. Baffigi,² L. Labate,² L. A. Gizzi,² L. Vassura,^{9,10} J. Fuchs,⁹ M. Passoni,³ M. Květon,⁵ F. Novotny,⁵ M. Possolt,⁵ J. Prokůpek,^{5,6} J. Proška,⁵ J. Pikial,^{5,6} L. Štolcová,^{5,6} A. Velyhan,⁶ M. Bougeard,¹ P. D'Oliveira,¹ O. Tcherbakoff,¹ F. Réau,¹ P. Martin,¹ and A. Macchi^{2,11,†} ¹CEA/IRAMISSPAM, F-91191 Gif-sur-Yvette, France
 ²Istituto Nazionale di Ottica, Consiglio Nazionale delle Ricerche, research unit "Adriano Gozzini," 56124 Pisa, Italy ³Dipartimento di Energia, Politecnico di Milano, 20133 Milano, Italy
 ⁴UUL, Université Pierre et Marie Curie, Ecole Polytechnique, CNRS, CEA, 75252 Paris, France
 ⁵ENSPE, Czech Technical University in Prague, CR: R1519 Prague, Czech Republic
 ⁶Institute of Physics of the ASCR, ELI-Beamlines project, Na Slovance 2, 18221 Prague, Czech Republic
 ⁷CEATDSMLS, CNRS, Ecole Polytechnique, 91128 Palaiseau Cedex, France
 ⁹ULUI, UMR7605, CNRS-CEA-Ecole Polytechnique-Paris 6, 91128 Palaiseau Cedex, 17ance
 ¹⁰Dipartimento di Energia di Roma "La Sapienza," Via A. Scarpa 14, 00161 Roma, Italy
 ¹¹Dipartimento fisica "Enrico Fermi," Università di Liva, Largo Bruno Pontecorvo 3, 1-56127 Pisa, Italy

Andrea Macchi

Experimental evidence: grating-enhanced TNSA LaserLAB experiment at SLIC, CEA Saclay laser UHI, 28 fs, 5 × 10¹⁹ W cm⁻²



Andrea Macchi

CNR/INO

Surface wave in PIC simulations

Snapshots of EM fields show localized wave propagating along the surface at resonant angle of incidence (plus reflection at various diffraction orders)



Foam targets for low density-enhanced absorption

Deposit layer at $n_e \simeq n_c$ on thin foil \rightarrow proton energy doubles up to 15 MeV (3D simulation, 25 fs pulse, 1 J energy)

Sgattoni, Londrillo, Macchi, Passoni, PRE **85** (2012) 036405



Fig. 8 - SEM plane view images of foam samples deposited, respectively, with 0.03 mg/s (1 sccm) (a) and 3 mg/s (100 sccm) (b) longitudinal gas flow.

Low-density foam production and characterization:

Zani et al, Carbon 56 (2013) 358





イロン イヨン イヨン イヨン

Foam-enhanced fast electron generation



Coupling of *P*-component of **E** with channel walls accelerates electrons: similarity with cone targets (\rightarrow)

Experiment recently performed on Petawatt laser at GIST, Gwangju, South Korea (data analysis in progress)

High absorption and fast electrons temperature $T_f \gtrsim 3T_{\text{pond}}$ for optimal foam areal density $n_e \ell$



CNR/INO

Andrea Macchi

Enhanced TNSA in microcone targets

[Gaillard et al, PoP 18 (2011) 056710]

Up to \mathcal{E}_{co} =67.5 MeV protons with 80 J pulse energy in cone targets

Efficient coupling to side walls as in the channel case: similar mechanism in action

[Kluge et al, New J. Phys. 14 (2012) 023038]





イロト イヨト イヨト イヨト

Beyond TNSA: searching for other mechanisms

Unresolved issues and limitations of TNSA:

- broad (~ exponential) energy spectrum (slow progress from engineered targets)
- ► slow scaling with laser intensity ($\mathscr{E}_{max} \sim I^{1/2}$)
- high repetition rate not easy with thin solid targets
- structured targets may be complex and/or expensive

< ロ > < 同 > < 回 > < 回 >

Early vision of radiation pressure acceleration (1966)

22

NATURE

JULY 2, 1966 VOL. 213 α -Centauri

INTERSTELLAR VEHICLE PROPELLED BY TERRESTRIAL LASER BEAM

By PROF. G. MARX Institute of Theoretical Physics, Roland Eötvös University, Budapest

A solution to "Fermi's paradox": "Laser propulsion from Earth ...would solve the problem of acceleration but not of deceleration at arrival ...no planet could be invaded by unexpected visitors from outer space"



イロト イヨト イヨト イヨト

The accelerating mirror model of RPA

Perfect mirror boosted by a plane wave: mechanical efficiency η and momentum transfer to mirror derived by Doppler shift and photon number conservation



イロト イヨト イヨト イヨト

CNR/INO

$$\frac{dp}{dt} = \frac{2I}{c}\frac{1-\beta}{1+\beta} \qquad \eta = \frac{2\beta}{1+\beta}$$

High efficiency $(\eta \rightarrow 1)$ but slow gain $(dp/dt \rightarrow 0)$ as $\beta \rightarrow 1$

Andrea Macchi

Analogy with Acceleration by Thomson Scattering

Light Sail equations of motion have the same form as those of a particle undergoing Thomson Scattering Landau & Lifshitz, *The Classical Theory of Fields*, ch.78 p.250 (1962).



Veksler's idea: coherent scattering by a cluster of radius $a \ll \lambda$ with $N \gg 1$ particles, $P_{sc} \rightarrow N^2 P_{sc} \Rightarrow \sigma_T \rightarrow N^2 \sigma_T$

イロト イヨト イヨト イヨト

Light Sail formulas, scaling, and needs

Favorable scaling with normalized fluence *F* "Perfect" monoenergeticity for "rigid", coherent sail motion Need of ultrathin (nm) foils and ultrahigh contrast pulses Circular polarization to reduce undesired heating Issues: slow energy gain, heating, transparency, deformation ...

Andrea Macchi

Relativistic transparency and optimal thickness

Thin plasma foil becomes transparent because of relativistic effects when

$$a_0 > \zeta \equiv \pi \frac{n_e}{n_c} \frac{\ell}{\lambda}$$
 $n_c = \frac{m_e \omega^2}{4\pi e^2}$ (cut-off density)

- → optimal thickness as trade-off between reduced mass and transparency: $a_0 = \zeta$
 - Diamond-Like Carbon ultrathin (nm) targets
- → avoid "prepulses" to cause early target disruption
 - ultrahigh-contrast systems
 - wide spots (and large energy), supergaussian intensity profiles to avoid strong deformation ?

イロト イヨト イヨト イヨト

CNR/INO

Andrea Macchi

*F*² scaling experimentally observed

PRL 109, 185006 (2012)

PHYSICAL REVIEW LETTERS

week ending 2 NOVEMBER 2012

Ion Acceleration in Multispecies Targets Driven by Intense Laser Radiation Pressure

S. Kar,^{1,*} K. F. Kakolee,¹ B. Qiao,^{1,†} A. Macchi,^{2,3} M. Cerchez,⁴ D. Doria,¹ M. Geissler,¹ P. McKenna,⁵ D. Neely,⁶
 J. Osterholz,⁴ R. Prasad, ¹ K. Quinn,¹ B. Ramakrishna,^{1,4} G. Sarri,¹ O. Willi,⁴ X. Y. Yuan,^{5,3} M. Zepf,^{1,7} and M. Borghesi^{1,8}
 ¹Centre for Plasma Physics, School of Mathematics and Physics, Queen's University Belfast, Belfast BT7 1NN, United Kingdom
 ²Istituto Nazionale di Ottica, CNR, Pisa, Italy
 ³Department of Physics, "Enrico Fermi," Largo B. Pontecorvo 3, 56127 Pisa, Italy
 ⁴Instituti für Laser-und Plasmaphysik, Heinrich-Heine-Universitä, Bülastedorf, Germany
 ⁵Department of Physics, SUPA, University of Strathclyde, Glasgow G4 00G, United Kingdom
 ⁶Central Laser Facility, Rutherford Appleton Laboratory, Didcot, Oxfordshire OX11 0QX, United Kingdom
 ⁸Institute of Physics of the ASCR, ELI-Beamlines Project, Na Slovance 2, 18221 Prague, Czech Republic (Received 23 February 2012; published 2 November 2012)

Andrea Macchi

・ロト ・回ト ・ヨト ・ヨト



*F*² scaling experimentally observed

VULCAN laser, RAL/CLF: Laser pulse: $t_p \approx 800 \ fs$ $3 \times 10^{20} \ W \ cm^{-2}$ $\sim 10^9 \ contrast$ Target: $\sim 0.1 \ \mu m$ metal foil



Multispecies (Z/A = 1, 1/2) peaks observed with $\Delta \mathscr{E}/\mathscr{E} \simeq 20\%$ Up to $\simeq 10$ MeV/amu observed at high flux Simulations suggest > 100 MeV/nucleon are within reach

S.Kar et al PRL 109 (2012) 185006

Other recent expts: Steinke et al, PRST-AB **16** (2013) 11303; Aurand et al, NJP **15** (2013) 33031

イロン イヨン イヨン イヨン

Pushing LS forward: "unlimited" acceleration?

Transverse expansion of the target reduces surface density $\rho\ell$

 \Rightarrow "unlimited" acceleration possible at the expense of the number of ions [Bulanov et al, PRL **104** (2010) 135003]

"Faster" gain $E_{ion}(t) \simeq (2It/\rho \ell c^2)^{3/5}$ predicted Mechanism is effective for *relativistic* ions ($\mathscr{F} \gg 1$)

Limitation: relativistic transparency when $a_0 > \zeta \equiv \pi \frac{n_e}{n_c} \frac{\ell}{\lambda}$ Relativistic increase of λ in "sail" frame delays breakthrough



< ロ > < 同 > < 回 > < 回 >

3D simulation campaign: LSAIL project

(Large Simulations of ion Acceleration by Intense Lasers)

- PRACE award for access to FERMI BlueGene/Q at CINECA, Italy
- Typical set-up: 4096 × 1792² grid points, 2 × 10¹⁰ particles, 16384 cores used
- Particle-In-Cell (PIC) codes:
- ALADYN: C. Benedetti, A. Sgattoni, G. Turchetti, P. Londrillo, IEEE Trans. Plasma Science **36** (2008) 1790
- PICCANTE: Open Source code (L.Fedeli, A.Sgattoni, S.Sinigardi, et al) github.com/ALaDyn/piccante



CNR/INO

Andrea Macchi

High energy gain in 3D RPA-LS simulations

Laser: 24 fs, 4.8 μ m spot, $I = 0.85 \times 10^{23}$ W cm⁻² \implies **1.5 kJ** Target: 1 μ m foil, $n_e = 1.1 \times 10^{23}$ cm⁻³, $\zeta \simeq a_0 \simeq 200$



<ロ> (四) (四) (日) (日) (日)

CNR/INO

 $\mathscr{E}_{max} \simeq 2.6 \text{ GeV} > 4 \text{ times 1D model prediction}$ Macchi et al, Plasma Phys. Contr. Fus. **55** (2013) 124020 Sgattoni et al, Appl. Phys. Lett. **105** (2014) 084105

Andrea Macchi

Sructures in ion density: Rayleigh-Taylor instability?



Formation of net-like structures with size ~ λ (laser wavelength) and ~ hexagonal

shape



CNR/INO

Interpretation: Rayleigh-Taylor instability driven by radiation pressure Symmetry analysis of nonlinear RTI in 3D: S. I. Abarzhi, Phys. Rev. E **59** (1999) 1729

Sgattoni, Sinigardi, Fedeli, Pegoraro, Macchi, arXiv:1404.1260

RTI in Radiation Pressure Acceleration

Thin foil target of areal density σ accelerated by a laser of intensity *I* is unstable with growth rate $\gamma = (P_0 q/\sigma)^{1/2}$ with $P_0 = 2I/c$ and *q* the wavevector [Ott, PRL **29** (1972) 1429]



(a) 5nm 4.4MeV

2D simulation [F.Pegoraro & S.V.Bulanov, PRL **99** (2007) 065002] Experimental indication from accelerated ion beam profile structures [C.Palmer et al, PRL **108** (2012) 225002]

Question: what sets the dominant wavevector q (and spatial scale of structures ~ q^{-1} ?)

Plasmonic modulation of radiation pressure

The EM field at a rippled surface (e.g. 2D reflecting, sinusoidal grating of period *d*) is modulated with plasmonic enhancement of the *P*-component when $d \sim \lambda$



Andrea Macchi

CNR/INO

(4) (3) (4) (4) (4)

Thin foil RTI with self-consistent pressure modulation

Model: reflection from shallow 2D grating of depth δ (first order in δ/λ) + modified Ott's theory with modulated pressure:

$$-(q^2 - k^2)^{1/2}$$
 (*S*)

CNR/INO

$$P \simeq P_0(1+K(q)\delta\cos qy), \qquad K(q) = \begin{cases} k^2 q(q^2-k^2)^{-1/2} & (P) \\ (k^2-q^2/2)(q^2-k^2)^{-1/2} & (C) \end{cases}$$



Andrea Macchi

Open Issues for Light Sail RPA

- Experimental results show important "non-pure LS" effects:
- broad, non-monoenergetic peaks in energy spectrum
- species separation
- weak dependence on polarization (tight focusing and target deformation effects?)
- Rayleigh-Taylor instabilities
- Use of ultrathin foil targets may be not optimal for high repetition rate operation

・ロト ・ 日 ・ ・ ヨ ・ ・ ヨ ・

Collisionless Shock Acceleration

• Concept: shock wave of velocity $v_s = Mc_s$ (M > 1, $c_s = \sqrt{ZT_e/Am_p}$) driven by the laser pulse into an ideal (collisionless) plasma (low density, high temperature)



- Shock front is a moving potential barrier → reflection of some ions from the shock front: v_i ≃ 2v_s
- → acceleration of *monoenergetic*, multi–MeV ions if v_s is constant and $T_e \simeq T_{pond}$ at $a_0 > 1$

Collisionless Shocks: Existence and Generation

Shocks do *not* exist in an *ideal* gas or plasma: some "dissipation" is necessary

→ ion reflection itself can provide dissipation in a collisionless plasma!

[Tidman & Krall, Shock Waves in Collisionless Plasmas (Wiley, 1971)]

Shock generation requires some "strong and sudden driver" (e.g. explosion): for laser-plasma interaction it may be driven by (a combination of)

<ロ> <同> <同> < 同> < 同>

- rapid heating of electrons at the interface
- "piston" effect of radiation pressure
- plasma instabilities (connection with astrophysics)

Monoenergetic CSA in CO₂ laser-H gas interaction

Proton spectra:

 $\mathcal{E}_{max} = 22 \text{ MeV} \qquad \Delta \mathcal{E} \lesssim 10^{-2} \mathcal{E}_{peak}$ Laser: 100 ps train of 3 ps pulses $I = 6.5 \times 10^{16} \text{ W cm}^{-2} (a_0 = 2.5)$ linear polarization Target: H₂ gas jet, $n_0 \leq 4n_c$ Claim: shock driven by fast electron pressure

- monoenergetic spectrum
- suitability for high repetition rate (flowing target)
- low efficiency



Figure 21 Proton energy spectra. a, Proton spectra obtained with a 100-ps-long laser pulse (red) and a 100 ps macropulse consisting of a number of 3 ps micropulses (blue) both containing 60.1. The typical noise level on a single CR39 detector was 100 pits. The total number of protons contained within the moneenergetic peak was 2.5×10^5 . b, The details of the energy spectra on four different laser shots with different macropulse structures (number of pulses and a_0 values ranging from 15 to 2.5).

イロン イロン イヨン イヨ

Haberberger et al Nature Phys. 8 (2012) 95

Andrea Macchi

Efficiency vs. monoenergetic spectrum

Reflected ions are on the tail of the ion distribution

 $v_i > v_s - (2e\Phi_m/m_i)^{1/2}$ (Φ_m potential jump)

Shock must lose part of its energy to accelerate ions

- \rightarrow shock's kinetic energy and velocity v_s decrease
- → velocity $2v_s$ of reflected ions decreases and spectrum broadens ("chirps") towards low energy
- → small loss is necessary for monoenergetic spectrum
- → number of accelerated ions must be small Demonstration in 1D simulation: vary the load of the shock (number of accelerated ions) by varying the background ion temperature T_i



A few ions in the tail of the warm distribution are reflected as a monoenergetic beam (v_s is constant)

イロト イヨト イヨト イヨト

CNR/INO

Andrea Macchi



CNR/INO

Andrea Macchi

Shock Loading and Energy Chirping in 2D

laser : $\tau_p = 45T$, $a_0 = 1$, $w = 5\lambda$; target: $n_e = 2n_c$, $T_i = 100 \text{ eV}$, Z/A = 1



Sgattoni et al, Proc. Spie 8779 (2013)

Resolution of low-density tail in the ion distribution is crucial \rightarrow challenge for PIC codes (high number of particles needed)

.

CNR/INO

"Vlasov" simulations of collisionless shocks

Alternative approaches to Vlasov equation for $f(\mathbf{r}, \mathbf{p}, t)$

$$\partial_t f + \mathbf{v} \cdot \partial_\mathbf{r} f + q \left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right) \cdot \partial_\mathbf{p} f = 0$$
 (+Maxwell eqs.)

- Particle-In-Cell: Lagrangian, discrete momentum space
- "Vlasov": Eulerian , continuum phase space
 Pro et contra for Vlasov:
- fine resolution of low-density tails
- no numerical noise
- computational effort much greater than PIC

CSA in 1D Vlasov simulations



For cold ions no reflection from the shock should occur (unless the shock becomes a "snowplow" and cannot propagate) BUT quite stable ion reflection is observed: what causes the necessary energy spread?

(Anna Grassi, M.Sc. Thesis, University of Pisa, 2013)

→ 프 → < 프 →</p>

Ion turbulence



Quasi-static ($\nu_f \simeq 0$), very low-frequency ion density modulation with $\simeq \lambda/4$ wavelength (λ : laser) originating from the target rear boundary (opposite to the laser interaction side) \rightarrow velocity spread for ion reflection

CNR/INO

Andrea Macchi

Origin of ion turbulence from "fast electron beats"



Fast electron bunches accelerated by $\mathbf{v} \times \mathbf{B}$ force have 2ω frequency, velocity ~ *c* and are *coherent* \rightarrow nonlinear plasma wave with $\lambda_p \simeq \lambda/2$ Reflection from the rear side causes "beating" of incident and reflected waves generating $\lambda_p/2 \simeq \lambda/4$ modulation

Model for ion turbulence - I

Pulsed electron bunches as "pump waves" ($\omega_{\pm} \simeq 2\omega, k_{\pm} \simeq \omega_{\pm}/c$):

$$(n_e^{(p)}, v_e^{(p)}) = \frac{(n_+, v_+)}{2} e^{ik_+ x - i\omega_+ t} + \frac{(n_-, v_-)}{2} e^{-ik_- x - i\omega_- t} + \text{C.C.}$$

Nonlinear coupling to background electrons and ions:

$$\partial_t n_e^{(NL)} + n_0 \partial_x v_e^{(NL)} \simeq -\partial_x (n_e^{(p)} v_e^{(p)}) \qquad \partial_t n_i \simeq -n_0 \partial_x v_i$$
$$\partial_t v_e^{(NL)} \simeq -\frac{e}{m_e} E^{(NL)} - v_e^{(p)} \partial_x v_e^{(p)} \qquad \partial_t v_i \simeq +\frac{e}{m_i} E^{(NL)} - \frac{T_e}{m_i} \frac{\partial_x n_i}{n_0}$$

CNR/INO

Low frequency component $E^{(NL)} = \tilde{E}e^{iKx-i\Omega t}$ etc.:

$$\Omega = \omega_+ - \omega_- \ll \omega \qquad K = k_+ + k_- \simeq 2 \frac{2\pi}{\lambda/2} = \frac{2\pi}{\lambda/4}$$

Andrea Macchi

Model for ion turbulence - II

Reflection of electron bunches from expanding sheath at rear side $\mathbf{v}' \simeq \mathbf{v} - 2\mathbf{u}$ $\lambda' \simeq \lambda(1 - u/v)$ $\Omega \simeq 2\omega(u/v) \simeq Ku$



Ion density oscillations

$$\tilde{n}_i = n_0 \frac{i e \tilde{E}}{m_i} \frac{K}{\Omega^2 - K^2 c_s^2} \qquad \tilde{v}_i = \frac{i e \tilde{E}}{m_i} \frac{\Omega}{\Omega^2 - K^2 c_s^2}$$

CNR/INO

→ 프 → < 프 →</p>

Andrea Macchi

Some conclusions and perspectives ...

- Progress in ion acceleration has been obtained on various sides (cut-off energy, efficiency, energy spread ...) in separate experiments with different mechanisms (more or less suitable for foreseen applications)
- Laser-driven ion beams already used for ultrafast plasma diagnostic and warm dense matter production
- Reaching required performance for other applications is still challenging:
- exploit new generation lasers
- improve target engineering
- develop large-scale simulations for experiment design

イロト イヨト イヨト イヨト

Funding acknowledgments

- ► LASERLAB-EUROPE, grant No. 284464, EU's 7th Framework Programme, proposal n.SLIC001693.
- ► PRACE supercomputing award, project LSAIL, for access to resource FERMI BlueGene/QTM at CINECA (Italy)
- MIUR (Italy):
- FIR project "Superintense Ultrashort Laser-Driven Ion Sources"

イロト イヨト イヨト イヨト

CNR/INO

- PRIN project "Laser-Driven Shock Waves"

EXTRA SLIDES

Andrea Macchi

CNR/INO

æ

<ロ> <同> <同> < 同> < 同>

Charging and "truncation" by electron escape

- An isolated, warm plasma in "real" 3D space gets charged due to the escape of N_{esc} electrons with energy > U_{esc} (since the binding potential is limited)
- For a simple spherical emitter of radius R having N_0 electrons at T_e :

$$N_{\rm esc} = N_0 \exp(-U_{\rm esc}/T_e)$$
 $U_{esc} = e^2 N_{\rm esc}/R$

- Message: cut-off energy U_{esc} (hence &_{max}) depends on target density, size, ...
- A: the system is neither steady nor in Boltzmann equilibrium, the target is neither isolated nor grounded, ...

CNR/INO

Andrea Macchi

Dynamic modeling of TNSA

Plasma expansion model: isothermal rarefaction wave solution "patched" at the ion front where quasi-neutrality breaks down

$$c_s = \left(\frac{ZT_e}{m_i}\right)^{1/2}, \quad u_f \equiv u_i(x_f) = c_s[2\ln(\omega_{pi}t) + 1], \quad \mathscr{E}_{\max} = \frac{m_i}{2}u_f^2 \propto ZT_e$$

∴ ion energy diverges due to infinite energy reservoir! assume finite model (e.g thin foil expansion) with $T_e(t)$ assume finite acceleration time (extra patch)



CNR/INO

Andrea Macchi

Some models fit better than others

Comparison of several models with experimental energies [Perego et al, Nucl.Inst.Meth.Phys.Res.A **653** (2011) 89]



Fitting parameters: laser pulse energy, power, intensity, duration; fast electron energy, density, divergence; ion density and distribution; target thickness; ... and various "phenomenological" quantities

Proton beam focusing and manipulation

TNSA-based "lenses" for spatial and spectral control of protons



Toncian et al, Science 312 (2006) 410



Kar et al, PRL **100** (2008) 105004

Andrea Macchi

Engineering the initial distribution of ions

Hydrogen-rich microdot for monoenergetic acceleration

Use of *bacteria* as hydrogencontaining layer



Schwoerer et al, Nature **439** (2006) 445

Dalui et al, Scient. Rep. 4 (2014) 1

CNR/INO

Andrea Macchi

Structured targets for enhanced TNSA

Fast electron production can be increased in targets with front side shaping or structuring (foams, microfunnels, ..., gratings)

Periodic structures allow resonant excitation of surface waves \rightarrow field enhancement and high absorption ("light caught by a grating") Matching possible

at a plasmavacuum interface if the grating is preserved (ultrahigh pulse contrast necessary!)

Andrea Macchi

 $\omega = ck$

 π/d

 $\omega_n/\sqrt{2}$

 $-q - \pi/d$

< □ > < 同

-2a

Need for ultraclean pulses: plasma mirrors

Plasma mirrors

yielding ~ 10^{12} pulseto-prepulse contrast allow to preserve target structuring until the short pulse interaction



Image: A matrix

figure from P. Gibbon, ibid., 369.

Andrea Macchi

How to make radiation pressure dominant?



The "Optical Mill" rotates in the sense *opposite* to that suggested by the imbalance of radiation pressure: *thermal* pressure due to *heating* dominates

< ロ > < 同 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < 回 > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ > < □ >

CNR/INO

Enforcing radiation pressure dominance requires to suppress heating of the surface Possible solution for ultraintense lasers: circular polarization

Transparency regime: Break-Out Afterburner

Transition to transparency: strong instability and volumetric heating of electrons Proton and C broad spectra at high energies and large number of particles (6% efficiency) Highest energies observed off-axis



[Jung et al NJP 15 (2013) 023007]

・ロト ・回ト ・ヨト ・ヨト

CNR/INO

Indication of > 150 MeV cut-off for protons! [Hegelich et al, APS Conf. 2011; arXiv:physics/1310.8650]

RPA in Thick Targets: "Hole Boring"

"Piston" push at a reflecting plasma surface bores a hole accelerating ions at velocity v_i Momentum flow balance:

 $P_{\rm EM} \doteq P_{\rm kin}$

(steady assumption)



・ロン ・四 ・ ・ 回 ・ ・ 日 ・

CNR/INO

 $I/c \sim (m_i n_i v_i) v_i \Rightarrow v_i \sim (I/m_i n_i c)^{1/2}$

Energy scaling $\mathscr{E}_i \sim v_i^2 \sim n_i^{-1}$ suggests to use low densities $n_i \gtrsim n_c$ (cut-off density): possible with gas targets

Andrea Macchi

Hole Boring RPA with gas H target and CO₂ laser

Narrow proton spectra at $\mathscr{E}_{peak} = 0.8 - 1.2$ MeV ($\Delta \mathscr{E} / \mathscr{E}_{peak} \simeq 20\%$ spread) observed from H gas jet at $n_e = 4 - 8n_c$ CO₂ ($\lambda = 10 \ \mu$ m) laser $I = 6.5 \times 10^{15}$ W cm⁻² circular polarization

Scaling with I/n_e and number of protons consistent with HB acceleration



FIG. 1 (color online). Raw and processed proton spectra for varying peak density n and vacuum intensity I showing scaling of peak proton energy $E_{max} \ll I/nc$ [MeV]. Parameter I/n shown to the right of the respective raw images. Shots taken with (a) I = 6.4, $n = 6.1n_{cr}$, (b) I = 5.5, $n = 6.1n_{cr}$, (c) I = 5.9, $n = 7.6n_{cr}$, (d) I = 5.7, $n = 8.0n_{cr}$ (I in units of 10¹⁵ W cm⁻²). (e) Background subtracted (solid lines) and also corrected (dashed lines) spectra. Heights of corrected spectra adjusted to match those of raw lineouts. Lineout corresponding to (b) reduced 4× to fit on the same scale.

イロン イヨン イヨン イヨン

CNR/INO

Palmer et al, PRL 106 (2011) 14801

Andrea Macchi