Surfin' the Surface Wave: a New Approach to Plasma-based Acceleration

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International Conference on Plasma Science and Applications (ICPSA22) Invited-2-3b, December 29, 2022

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Example of Acceleration by a Strong Surface Wave



From: T. Katsouleas, "On the node of a wave", Nature 444 (2006) 688

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Producing the Perfect Wave for Electrons

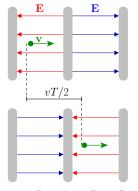
To accelerate either surfers or electrons a wave must have:

- 1. longitudinal force compoment
- 2. phase velocity $v_p \approx$ surfer velocity v
- $\longrightarrow v_p \lesssim c$ for relativistic electrons
- LINAC principle: sequence of cavity with alternating fields

$$L \simeq vT/2 \simeq cT/2 = \pi c/\omega$$

A plasma wave in a cold plasma with phase velocity v_p can be seen as a sequence of "cavities" with $T = \pi/\omega_p$ and $L = v_p T_p$ $\omega_p = (4\pi e^2 n_e/m_e)^{1/2}$ plasma frequency

Tutorial: A. Macchi, Am. J. Phys. 88 (2020) 723



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Wake Waves

An object (e.g. duck) moving at velocity V produces a wake of oscillations with $v_p = V$

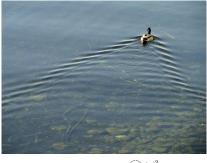
Bodensee at Bad Schachen, Lindau, Germany.

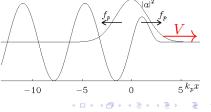
Photo by Daderot, Wikipedia, public domain.

Laser wakefield acceleration: a pulse of duration $\simeq \pi/\omega_p$ drives a plasma wake with $\upsilon_p = c(1 - \omega_p^2/\omega^2)^{1/2} \lesssim c$ T.Tajima & J.Dawson, Phys. Rev. Lett. **43**(1979) 267

[Textbook: A. Macchi, A Superintense Laser-Plasma

Interaction Theory Primer (Springer, 2013)]



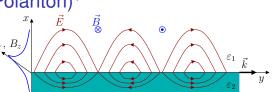


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Surface Plasmon (Polariton)¹

SP: a building block of plasmonics E_y, B_z (mostly studied in the *linear* regime)



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SP excitation \longrightarrow EM field confinement and enhancement Interface between vacuum and "simple metal" (cold plasma):

$$\varepsilon_1 = 1 \qquad \varepsilon_2 = 1 - \frac{\omega_p^2}{\omega^2} < -1$$
$$k = \frac{\omega}{c} \left(\frac{\omega_p^2 - \omega^2}{\omega_p^2 - 2\omega^2} \right)^{1/2} \qquad \omega < \frac{\omega_p}{\sqrt{2}} \qquad v_p = \frac{\omega}{k} < c$$

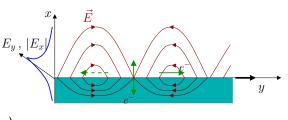
¹aka Surface Plasma Wave

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Surfin' the Surface Wave

SP can accelerate electrons much like "bulk" plasma waves:

- longitudinal
- *E*-component (E_y)
- phase velocity $v_{\rm p} \lesssim c$ (with $v_{\rm p} \rightarrow c$ when $\omega_p \gg \omega$)



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Simple model predicts for maximum energy and emission angle (relative to surface):

$$\mathcal{E} \simeq m_e c^2 a_{\rm SP} \frac{\omega_p^2}{\omega^2} \qquad \tan \phi_e = \frac{p_x}{p_y} \simeq \frac{1}{\gamma_{\rm P}} \qquad \left(a_{\rm SP} \equiv \frac{eE_y}{m_e\omega c}\right)$$

Macchi et al, Phys. Plasmas 25 (2018) 031906; 26 (2019) 042114

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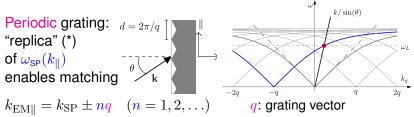
Exciting Surface Plasmons with Laser Light

Phase matching requirement between SP and plane EM wave $(\omega_{\text{FM}} = |\mathbf{k}_{\text{FM}}|c)$: $\varphi_{\mathsf{EM}} = \varphi_{\mathsf{SP}}$ where $\varphi = \mathbf{k}_{\parallel} \cdot \mathbf{r} - \omega t$ i.e. phase velocities $v_p = \omega/k_{\parallel}$ must be equal $v_{\text{EM}} = v_{\text{SP}}$ EM No matching $k / \sin(\theta)$ with EM wave at a plane interface: k_y k.Laser $v_{\text{EM}} = \frac{c}{\sin \theta} > c$ $v_{\text{SP}} = c \left(\frac{2 - \omega_p^2 / \omega^2}{1 - \omega_r^2 / \omega^2} \right)^{1/2} < c$

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Grating Coupling



Matching occurs at "resonant" incidence angles (for $\omega_p/\omega \gg 1$) $\sin \theta \simeq n \frac{\lambda}{d} - 1$ (= diffraction order along the surface) - usually n = 1

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- actually an infinite grating is not strictly required (a local surface modulation may suffice)
- (*) folding in the Brillouin zone Floquet-Bloch theorem

The "Extreme" Interaction Regime

High laser irradiance: electron dynamics is relativistic and strongly nonlinear

Relativistic parameter: $a_0 = \frac{eE}{m_e \omega c} = 0.85 \left(\frac{I\lambda^2}{10^{18} \text{ Wcm}^{-2} \mu \text{m}^2}\right)^{1/2}$ Femtosecond pulse duration: target density profile preserved

first investigation of SP in this regime (with unknown features: no relativistic SP theory available)
SP may be driven to high amplitude close to "breaking" threshold with self-injection of electrons



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Macchi et al, Phys. Plasmas 26 (2019) 042114

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Observation of "Surfing" Acceleration

PRL 116, 015001 (2016) PHYSICAL REVIEW LETTERS

week ending 8 JANUARY 2016

Electron Acceleration by Relativistic Surface Plasmons in Laser-Grating Interaction

L. Fedeli, ^{1,2,*} A. Sgattoni,² G. Cantono, ^{3,4,1,2} D. Garzella,³ F. Réau,³ I. Prencipe,^{5,†} M. Passoni,⁵ M. Raynaud,⁶ M. Květoň,⁷ J. Proska,⁷ A. Macchi,^{2,1} and T. Ceccoti³ ¹Enrico Fermi Department of Physics. University of Pisa, 5lc127 Pisa, Italy ²National Institute of Optics, National Research Council (CNR/INO), u.o.s Adriano Gozzini, 56124 Pisa, Italy ³LIDTL, CEA, CNRS, Université Paris-Saclay, CEA Saclay, 91191 Gif-sur-Yvette, France ⁴University of Paris Sud, Orsay 914005, France ⁵Department of Energy, Politecnico di Milano, Milan 20156, Italy ⁶Laboratoire des Solides irradiés, Ecole Polytechnique, CNRS, CEA/DSM/IRAMIS, Université Paris-Saclay, 91128 Palaiseau Cedex, France ⁷FNSPE, Czech Technical University, Prague 11519, Czech Republic (Received 30 June 2015; published 7 January 2016)

L. Fedeli et al, Phys. Rev. Lett. 116 (2016) 015001

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Observation of "Surfing" Acceleration

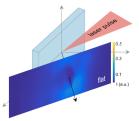


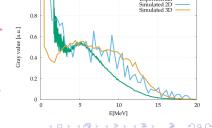
UHI laser: 25 fs pulse $5 \times 10^{19} \text{ W cm}^{-2}$, contrast $\gtrsim 10^{12}$ at 5 ps collimated ($\simeq 20^{\circ}$ cone) electron emission near the surface tangent ($\phi \simeq 2^{\circ}$) multi-MeV energy, Total charge $\simeq 100 \text{ pC}$ (up to $\simeq 650 \text{ pC}$ with blazed gratings: C. Cantana et al. Dava Plaamaa **25** (2018) 021007)

G. Cantono et al, Phys. Plasmas 25 (2018) 031907)



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Observation of SP-enhanced XUV High Harmonics

PHYSICAL REVIEW LETTERS 120, 264803 (2018)

Extreme Ultraviolet Beam Enhancement by Relativistic Surface Plasmons

G. Cantono,^{1,2,3,4,*} L. Fedeli,⁵ A. Sgattoni,^{6,7} A. Denoeud,¹ L. Chopineau,¹ F. Réau,¹ T. Ceccotti,¹ and A. Macchi^{3,4}
 ¹LIDYL, CEA, CNRS, Université Paris-Saclay, CEA Saclay, 91191 Gifsur-Yvette, France
 ²Université Paris Sud, Paris, 91400 Orsay, France
 ³National Institute of Optics, National Research Council (CNRINO) A. Gozzini unit, 56124 Pisa, Italy
 ⁴Enrico Fermi Department of Physics, University of Pisa, 56127 Pisa, Italy
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 ⁷LESIA, Observatoire de Paris, CNRS, UPMC: Sorbonne Universites, 92195 Meudon, France

G. Cantono et al, Phys. Rev. Lett. 120 (2018) 264803

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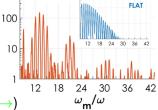
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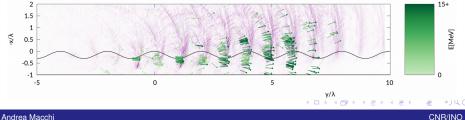
Observation of SP-enhanced XUV High Harmonics

Enhanced High Harmonic emission observed when Surface Plasmons are excited

Simulations show coherent scattering from self-organized electron bunches (\rightarrow) to produce quasi-collinear HH (similar mochanism to collective instability



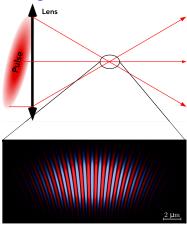
(similar mechanism to collective instability in a FEL)



In-Pulse Rotation of Incidence Angle

Inducing wavefront rotation (WFR) in the laser pulse the effective incidence angle rotates in time \rightarrow "resonant" condition holds only for an interval shorter than the driving pulse





WFR obtained by focusing a tilted wavefront pulse

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Near "Single-Cycle" Surface Plasmon Polariton



Few-Cycle Surface Plasmon Polariton Generation by Rotating Wavefront Pulses

F. Pisani,*^{,†}[®] L. Fedeli,^{*,‡} and A. Macchi*^{,¶,†}[®]

[†]Enrico Fermi Department of Physics, University of Pisa, 56127 Pisa, Italy [‡]Department of Energy, Politecnico di Milano, 20133 Milano, Italy [¶]National Institute of Optics, National Research Council (CNR/INO), A.Gozzini unit, 56124 Pisa, Italv

F. Pisani, L. Fedeli, A. Macchi, ACS Photonics 5 (2018) 1068

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Near "Single-Cycle" Surface Plasmon Polariton

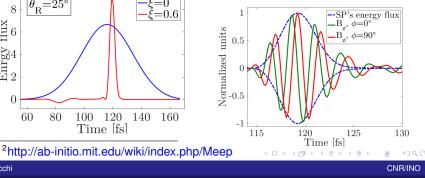
MEEP² simulations of WFR pulse on Ag grating

SPP w/o and with WFR duration: 3.8 fs (~ 1.4 cycles)

 $\times 10^{-3}$ $\theta_{\rm B} = 25^{\circ}$ 8 =0.6Enrgy flux 9 2 0 60 80 140160100120Time [fs]

 $E = E(r, z, t) \exp(-i\omega_L t + ir\xi t + \phi)$ E : WFR parameter laser: 30 fs , $\lambda_L = 0.8 \ \mu m$

> dependence on absolute phase ϕ



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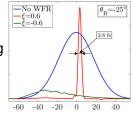
WFR Enhancement of SPP Amplitude

"Sliding focus" effect: WFR makes the laser spot move along the target with velocity

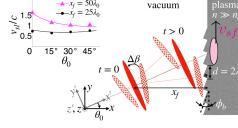
$$v_{sf} \simeq rac{\Delta eta \, x_f}{\lambda \cos^2 heta_0} \propto x_f \boldsymbol{\xi}$$

 $(x_f: waist-to-target distance)$

When $v_{sf} \simeq c$ the SPP is "sustained" along its propagation: increase of peak amplitude (Note the effect of the **sign** of ξ)



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WFR Effect on Electron Acceleration

PHYSICAL REVIEW E 103, L021201 (2021)

Letter

Ultrashort high energy electron bunches from tunable surface plasma waves driven with laser wavefront rotation

S. Marini , ¹² P. S. Kleij , ^{1,2,3} F. Pisani , ³ F. Amiranoff, ² M. Grech , ² A. Macchi , ¹³ M. Raynaud , ¹ and C. Riconda , ^{2,*} ¹LSJ, CEADRF/IRAMIS, CNRS, École Polytechnique, Institut Polytechnique de Paris, F-91128 Palaiseau, France ²LULI, Sorbonne Université, CNRS, CEA, École Polytechnique, Institut Polytechnique de Paris, F-75252 Paris, France ³Enrico Fermi Department of Physics, University of Pisa, Iargo Bruno Pontecorvo 3, 56127 Pisa, Italy ⁴National Institute of Optics, National Research Council (CNR/INO), Adriano Gozzini laboratory, 56124 Pisa, Italy

S. Marini, P. Kleij, F. Pisani, F. Amiranoff, M. Grech, A. Macchi, M. Raynaud, C. Riconda Phys. Rev. E **103** (2021) L021201

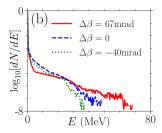
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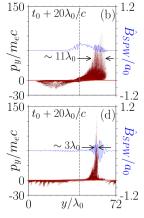
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WFR Effect on Electron Acceleration

SMILEI open source PIC code 27 fs & 4×10^{19} W cm⁻² laser pulse WFR may double the cut-off energy of the electron bunch while shortening its duration down to 8 fs (simulations by S. Marini and P. Kleij)





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waiting for experiments

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Grating Drawbacks

- grating targets are expensive
- need of ultrahigh contrast pulses free from spurious "prepulses" to preserve the shallow modulation from early damage
- strong EM scattering losses of the SP propagating along the grating (inverse to the generation process)

(might be reduced by having the grating only in the laser spot but would require perfect pointing stability)

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— Can we do surfing acceleration without a grating?

Acceleration with No Grating Attached

New Journal of Physics

Surface plasmon-driven electron and proton acceleration without grating coupling

J Sarma¹⁽⁰⁾, A McIlvenny¹⁽⁰⁾, N Das²⁽⁰⁾, M Borghesi^{1,*}⁽⁰⁾ and A Macchi^{3,4,*}⁽⁰⁾

¹ Centre for Plasma Physics, The Queen's University of Belfast, University Road BT71NN, Belfast, United Kingdom

² Tezpur University, Tezpur, India

³ National Institute of Optics, National Research Council (CNR/INO), Adriano Gozzini Laboratory, Pisa, Italy

4 Enrico Fermi Department of Physics, University of Pisa, Pisa, Italy

* Authors to whom any correspondence should be addressed.

J. Sarma et al, New J. Phys. 24 (2022) 073023

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Dephasing vs Acceleration Lengths

Revisiting (un-)matching between SP and incident laser pulse: phase difference after propagating over a distance L

$$\Delta \phi = (k_{\rm EM\parallel} - k_{\rm SP})L = \left(\frac{1}{v_{\rm EM}} - \frac{1}{v_{\rm SP}}\right)\omega L$$
$$\Delta \phi \doteq \pi \quad \longrightarrow \quad L = \frac{\pi}{k_{\rm SP} - k_{\rm EM}} \equiv L_{\rm dep}$$
At grazing incidence $(\alpha = \pi/2 - \theta \ll 1)$ $L_{\rm dep} \simeq \frac{\lambda}{\alpha^2 + \omega^2/\omega_n^2}$

Acceleration is not limited by dephasing when

$$L_{\rm dep} > L_{\rm acc} = \frac{\mathcal{E}_{\rm max}}{eE_{\rm SP}} \simeq \frac{\lambda}{\pi} \frac{\omega_p^2}{\omega^2}$$

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Coupling without Grating

To drive the SP efficiently the incident EM field must have: 1) a non-vanishing component parallel to the surface (E_x) \leftarrow reflectivity $R = |r|^2 < 1$

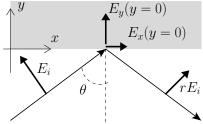


Image: A matrix

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2) a E_x/E_y ratio close as possible to the value for the SP:

$$\frac{E_{\text{EM},x}}{E_{\text{EM},y}}\Big|_{y=0^+} = -\varepsilon \frac{1-r}{1+r} \frac{\cos\theta}{\sin\theta} \qquad \frac{E_{\text{SP},x}}{E_{\text{SP},y}}\Big|_{y=0^+} = -i|\varepsilon|^{1/2}$$

asymptotically equal for $\sin \theta \rightarrow 1$

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Simulation Set-Up

PIC code EPOCH 2 simulations by J. Sarma & A. McIlvenny 1 Length geometry 🗄 Thickness Cartesian 2D 0 SPW Target: fully ionized Au $^{-1}$ with CH contaminant layer electron density -2 $n_e = 1.7 \times 10^{23} \text{ cm}^{-3}$ 20 40 60 thickness $d = 0.8 \ \mu \mathrm{m}$ $x(\mu m)$ length $L_T = (90 - 200) \ \mu m$ Pulse: $\lambda = 0.8 \ \mu m \ (\omega_p = 10\omega)$ Gaussian profiles, width $6.5 \ \mu m$, duration $35 \ fs$ (FWHM) intensity $I = (0.34 - 7.8) \times 10^{20} \text{ W cm}^{-2}$ "relativistic" parameter $a_0 = (5 - 19)$ イロン イヨン イヨン イヨン

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Acceleration

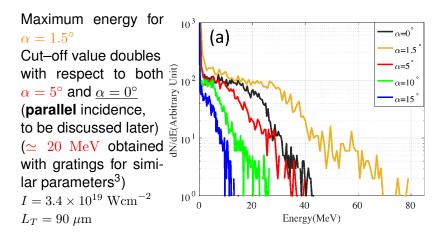
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Electron Energy Increase at Grazing Incidence



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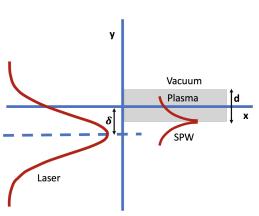
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³Cantono et al, PoP **25** (2018) 031907

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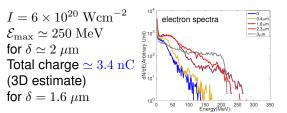
Parallel Incidence

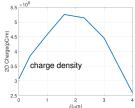
For parallel incidence $(\alpha = 0^{\circ})$ the laser pulse will interact with the (left) short edge of the target Additional parametric dependence on the "focal spot shift" δ \equiv distance between the laser propagation axis and the target midplane (surface at $y = -\delta/2$

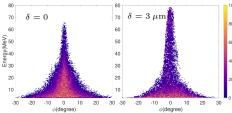


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Enhancement by "Spot Shift" at Parallel Incidence







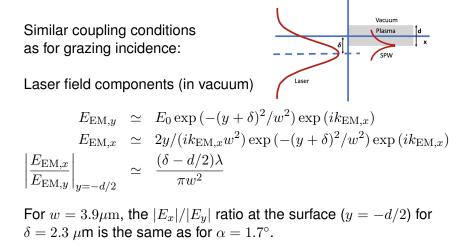
Energy-angle distributions $(I = 3.4 \times 10^{19} \text{ W cm}^{-2})$ Electrons are strongly collimated with almost symmetrical distribution even for "asymmetrical" interaction with $\delta \neq 0$

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On the Effect of "Spot Shift"



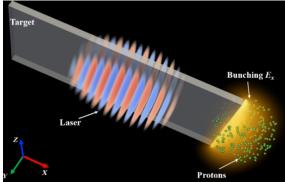
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Proposed "Peeler" Proton Acceleration

"A fs laser (red and blue) is incident on the edge of a micron-thick tape (grey) [...] Abundant electrons are accelerated forward by the intense laser."



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3D simulations with VLPL and EPOCH codes

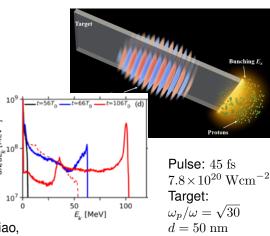
X.Shen, A.Pukhov, B.Qiao, Phys. Rev. X **11** (2021) 041002

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Proposed "Peeler" Proton Acceleration

"[...] at the rear edge a longitudinal bunching field is established (vellow). Protons Jreen us simultaneously accelerated and ting to a highly beam."

X.Shen, A.Pukhov, B.Qiao, Phys. Rev. X **11** (2021) 041002



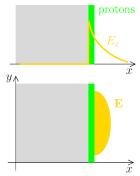
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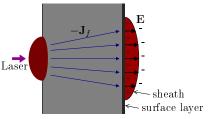
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Origin of Monoenergetic Proton Spectra

Protons from surface impurity layer are accelerated in a sheath generated by "hot" electrons (Target Normal Sheath Acceleration)





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For a standard "wide" target the E-field is inhomogeneous both tranversally and longitudinally (due to self-screening by protons) → broad proton spectra Review on ion acceleration: A. Macchi, M. Borghesi, M. Passoni, Day Mod. Phys. **25** (2012) 751

Rev. Mod. Phys. 85 (2013) 751

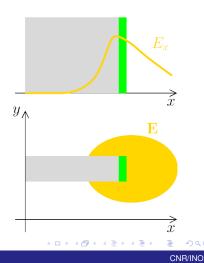
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Origin of Monoenergetic Proton Spectra

"Peeler" scheme reduces proton number with transverse localization and exploits high electron charge produced by SP-driven acceleration

Protons are now less than hot electrons: no E-field screening E-field is spatially "smooth" in both directions

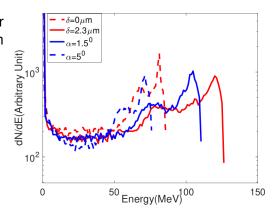
- \longrightarrow all protons see the same field
- \longrightarrow monoenergetic acceleration



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Proton Spectra: 2D Simulation

Highest cut-off energy is reached for parallel incidence with "shifted" pulse $(\delta = 2.3 \ \mu m)$ $(\sim X 2 \text{ increase with})$ respect to $\delta = 0$) Slightly lower energy at grazing incidence $(\alpha = 1.5^{\circ})$ $I = 6 \times 10^{20} \, \mathrm{W cm}^{-2}$



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Conclusions

- Superintense laser-driven Surface Plasmons drive
 - "surfing" acceleration of high charge electron bunches

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- XUV high harmonics by coherent scattering
- Perspectives for:
 - near single-cycle SP generation
 - SP-driven acceleration without grating coupling
 - "peeler" monoenergetic proton acceleration Download this talk:

https://osiris.df.unipi.it/ macchi/talks.html

EXTRA SLIDES

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Surfin' the Surface Wave

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Simple Model of SP Acceleration - I

SP field on the vacuum side is electrostatic in the wave frame S' moving with phase velocity $\beta_p = v_p/c$ with respect to S (lab) Electrostatic potential in S':

$$\Phi' = -\left(\frac{\gamma_{\rm p} E_{\rm SP}}{k}\right) e^{k'x} \sin k'y' \qquad k' = \frac{k}{\gamma_{\rm p}}$$
^(e) in the co-moving gain depends on the "kick angle" from the top of the potential hill

 $(1 \quad 0^2) - 1/2$

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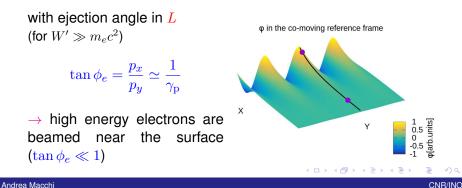
Simple Model of SP Acceleration - II

Assume as the most likely case an electron going downhill along the *x*-direction and acquiring an energy $W' = eE_{SP}/k'$

0

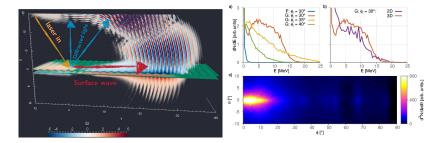
$$W \simeq \gamma_{\rm p} W' \simeq m_e c^2 a_{\rm SP} \frac{\omega_p^2}{\omega^2}$$

$$(a_{\rm SP} = eE_{\rm SP}/m_e\omega c)$$



3D Simulations PICcante code⁴

A.Sgattoni, L.Fedeli, S.Sinigardi et al, arXiv:1503.02464



Simulations match experimental observations quantitatively and in detail

⁴Particle-In-Cell Code for AdvaNced simulations on TiEr-0 systems =

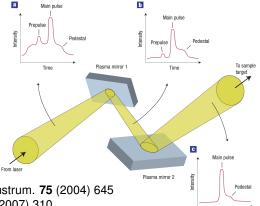
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"Ultraclean" high-contrast pulses

Ionization shutters ("plasma mirrors") yield pulse-toprepulse intensity contrast > 10^{11} \rightarrow sub-wavelength structuring is preserved until the short pulse interaction



Time

CNR/INO

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B. Dromey et al, Rev. Sci. Instrum. **75** (2004) 645
A. Levy et al, Opt. Lett. **32** (2007) 310
C. Thaury et al, Nature Physics **3** (2007) 424
figure from P. Gibbon, *ibid.* 369

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First evidence from proton emission

PRL 111, 185001 (2013)

PHYSICAL REVIEW LETTERS

week ending 1 NOVEMBER 2013

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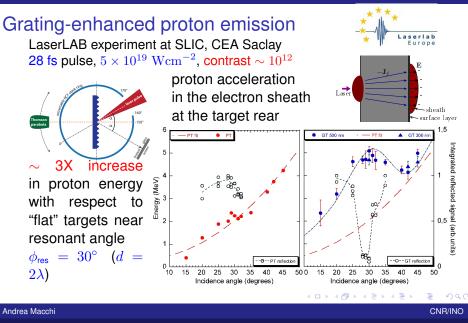
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Evidence of Resonant Surface-Wave Excitation in the Relativistic Regime through Measurements of Proton Acceleration from Grating Targets

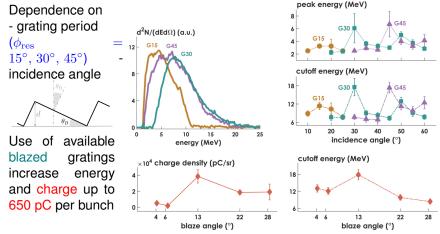
T. Ceccotti,^{1,*} V. Floquet,¹ A. Sgattoni,^{2,3} A. Bigongiari,⁴ O. Klimo,^{5,6} M. Raynaud,⁷ C. Riconda,⁴ A. Heron,⁸ F. Baffigi,² L. Labate,² L. A. Gizzi,² L. Vassura,^{9,10} J. Fuchs,⁵ M. Passoni,³ M. Květon,⁵ F. Novotny,⁵ M. Possolt,⁵ J. Prokůpek,^{5,6} J. Proška,⁵ J. Proška,⁵ J. Proška,^{5,6} L. Velyhan,⁶ M. Bougeard,¹ P. D'Oliveira,¹ O. Tcherbakoff,¹ F. Réau,¹ P. Martin,¹ and A. Macchi^{2,11,7} ¹ CEA/IRAMISSPAM, F-91191 Gif-sur-Yvette, France
²Istituto Nazionale di Onica, Consiglio Nazionale delle Ricerche, research unit "Adriano Gozzini," 56124 Pisa, Italy ³Dipartimento di Energia, Politecnico di Milano, 20133 Milano, Italy ⁴UULI, Université Pierre et Marie Curie, Ecole Polytechnique, CNRS, CEA, 75252 Paris, France ⁵FNSPE, Czech Technical University in Prague, CR-11519 Prague, Czech Republic ⁶Institute of Physics of the ASCR, ELI-Beamlines project, Na Slovance 2, 18221 Prague, Czech Republic ⁷CEA/DSM/LSI, CNRS, Ecole Polytechnique-Paris 6, 91128 Palaiseau Cedex, France ⁸CPHT, CNRS, Ecole Polytechnique-Paris 6, 91128 Palaiseau, France ⁹LULI, UMR7605, CNRS-CEA-Ecole Polytechnique-Paris 6, 91128 Palaiseau, France ¹⁰Dipartimento BAI, Università di Roma ¹¹La Sapara, ¹¹Maria ¹¹Dipartimento di Fisca ¹¹Università di Pisa, Largo Bruno Pontecorvo 3, 1-56127 Pisa, Italy ¹¹Dipartimento di Fisca ¹¹Università di Pisa, Largo Bruno Pontecorvo 3, 1-56127 Pisa, Italy ¹¹Dipartimento di Fisca ¹¹Entrico Fermi, ¹¹Università di Pisa, Largo Bruno Pontecorvo 3, 1-56127 Pisa, Italy ¹¹Dipartimento di Fisca ¹¹CHU, ¹¹CHU,

T. Ceccotti et al, Phys. Rev. Lett. 111 (2013) 185001

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Optimizing SP-enhanced electron emission



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G. Cantono et al, Phys. Plasmas 25 (2018) 031907

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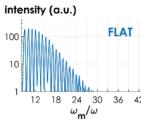
SP-enhancement and optimization of HH

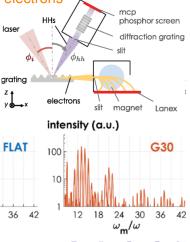
Simultaneous measurements of HH & electrons

HH optimization via density profile tailoring (scalelength $L \simeq 0.1\lambda_L$) by a femtosecond prepulse Kahaly et al, PRL **110** (2013) 175001 **Notice**: $L \sim$ grating depth!

Max HH order:





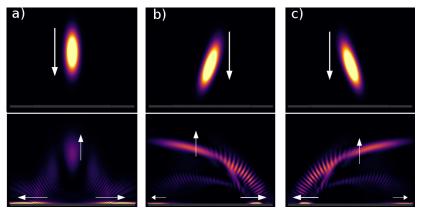


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Effect of WFR Direction

Normal incidence: excitation of two symmetric SPs) a): no rotation b): counterclockwise rotation c): clockwise rotation Note the scattering from the grating \rightarrow radiative loss of energy



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Laser Spot Stretching Effects

At grazing incidence the laser spot is stretched along $x \rightarrow$ the intensity on target decreases

$$I(\theta) = I(0)\cos\theta = I(0)\sin\alpha$$

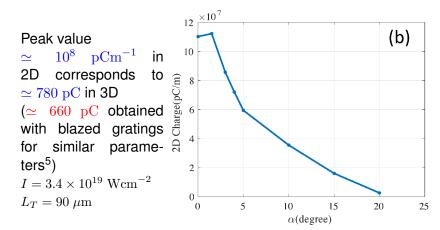
With respect to the case with grating (typical $\theta \leq 45^{\circ}$) the effect on electron acceleration may be compensated by:

- 1. slow scaling of electron energy $\mathcal{E}_e \propto E_{\mathrm{SP}} \propto \sin^{1/2} \alpha$
- 2. no loss from radiative scattering
- 3. no saturation (observed in grating simulations at $a_0 \gtrsim 10$)

Strong (unknown) nonlinear effects on SP are prevented

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Accelerated Charge for Different Angles



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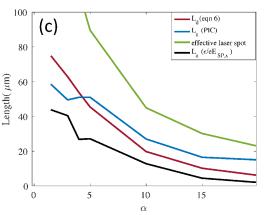
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⁵Cantono et al, PoP **25** (2018) 031907

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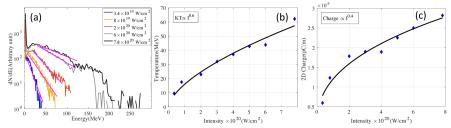
Acceleration Length for Different Angles

Two different estimates of the acceleration length (L_{acc}) are fairly close to the dephasing length L_{dep} as a function of α Spot width on target $L_S > L_{dep}$ (could be further optimized) $I = 3.4 \times 10^{19} \text{ W cm}^{-2}$ $L_T = 200 \ \mu m$



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Scaling with Laser Intensity



a) electron spectra b) electron "Temperature" c) charge density

Temperature exceeds "ponderomotive" values $(T_{\rm p} = m_e c^2 (\sqrt{(1 + a_0^2/2)} - 1))$ by one order of magnitude Peak charge density value $\simeq 3 \times 10^8 \ {\rm pCm^{-1}}$ in 2D corresponds to an estimate $\simeq 1.9 \ {\rm nC}$ in 3D

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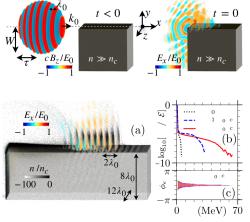
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A Possibly Similar Observation?

3D SMILEI simulations

S.Marini, P.Kleij, M.Grech, M.Raynaud, C.Riconda,

"Electron acceleration by laser plasma wedge interaction" arXiv:2202.08226

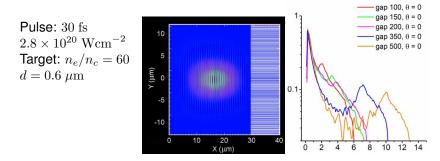


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Earlier Numerical Observation?

2D simulations of an array of parallel foils: electron acceleration attributed to SP peaked proton spectra appear for large gaps



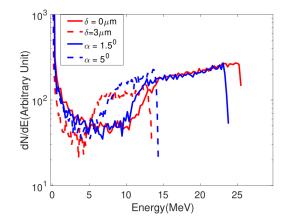
G. Cristoforetti et al, Plasma Phys. Control. Fusion 62 (2020) 114001

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Proton Spectra: 2D simulation (Low I)

Sharply peaked spectra are not apparent for lower intensity $I = 3.4 \times 10^{19} \text{ Wcm}^{-2}$



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Proton Spectra: 3D simulation (preliminary)

- electron spectra similar to 2D
- Shen et al.'s proton spectra reproduced

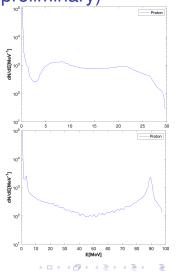
For a lower energy pulse (GEMINI class) the proton spectrum is deeply affected

- Geometrical effect on proton acceleration?
- Insufficient electron production?

$$I = 6 \times 10^{20} \text{ W cm}^{-2}$$

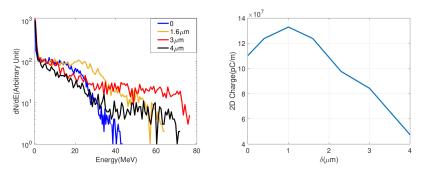
pulse width $3 \ \mu m$

(4 times narrower than Shen's)



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Effect of δ on Spectra and Charge (Low I)



a) electron spectra

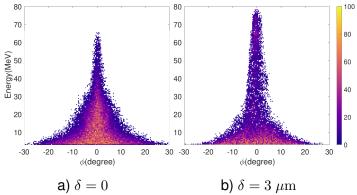
b) charge density

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 $I = 3.4 \times 10^{19} \text{ Wcm}^{-2}$ Maximum values are *not* for $|\delta| \le d/2$

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Electron Collimation



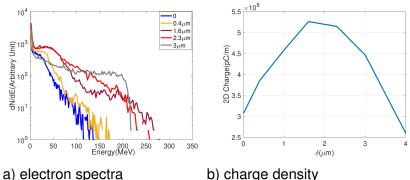
$$I = 3.4 \times 10^{19} \; {\rm W cm^{-2}}$$

Electrons are strongly collimated with almost symmetrical distribution even for "asymmetrical" interaction with $\delta \neq 0$

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Effect of δ on Spectra and Charge (High I)



b) charge density

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 $I = 6 \times 10^{20} \, \mathrm{W cm}^{-2}$ Total 3D charge would be $\simeq 3.4 \text{ nC}$ for $\delta = 1.6 \ \mu \text{m}$

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