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Radiation Pressure Acceleration in the “Light Sail” regime

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1. Light Sail Revisited

- The basic LS concept
- LS improved: Self-Induced Transparency effects
- The “dark mass” puzzle
- Electron and ion dynamics and self-organization

2. Radiation Reaction effects

- RR inclusion in PIC via Landau-Lifshitz equation

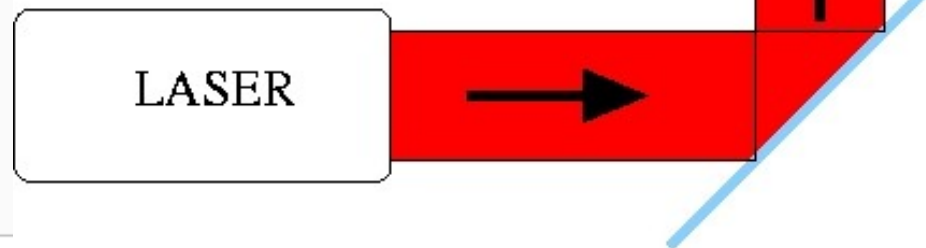


THE “LIGHT SAIL” CONCEPT

Originally proposed as a way to **accelerate a massive mirror** by the Radiation Pressure of an Earth-based laser

R.L.Forward, “Roundtrip interstellar travel using laser-pushed lightsails”,
J. Spacecraft and Rockets **21** (1964) 187

G.Marx, “Interstellar vehicle propelled by terrestrial laser beam”, Nature **211** (1966) 22

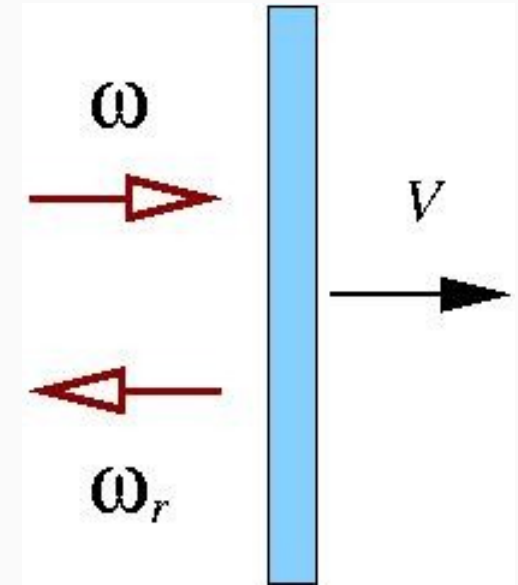




ACCELERATING MIRROR MODEL

perfectly reflecting, **rigid mirror** of mass $M = \rho \ell S$ boosted by a plane light wave

Mirror velocity as a function of the light pulse intensity I and duration τ and of the surface density $n_e \ell$ of the target:



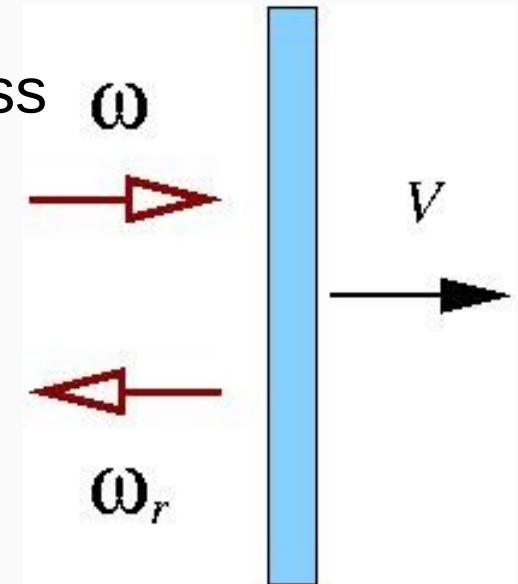
$$\beta(t) = \frac{(1 + \mathcal{E})^2 - 1}{(1 + \mathcal{E})^2 + 1}, \quad \mathcal{E} = \frac{2F(t)}{\rho \ell c^2} = 2\pi \frac{Z m_e a_0^2 \tau}{A m_p \zeta}$$

$$F(t) = \int_0^t I(t') dt' \propto a_0^2 \tau, \quad \zeta = \pi \frac{n_e \ell}{n_c \lambda}$$



MECHANICAL EFFICIENCY

The efficiency η of the acceleration process can be obtained by a simple argument of conservation of “number of photons” plus the **Doppler shift** of the reflected light:



$$N = \frac{IS\tau}{\hbar\omega}, \quad \omega_r = \omega \frac{1 - \beta}{1 + \beta}$$
$$\eta = \frac{\mathcal{E}_{\text{abs}}}{\mathcal{E}_{\text{laser}}} = \frac{N\hbar(\omega - \omega_r)}{IS\tau} = \frac{2\beta}{1 + \beta}$$
$$\beta \rightarrow 1 \Rightarrow \eta \rightarrow 1$$

100% efficiency in the relativistic limit!



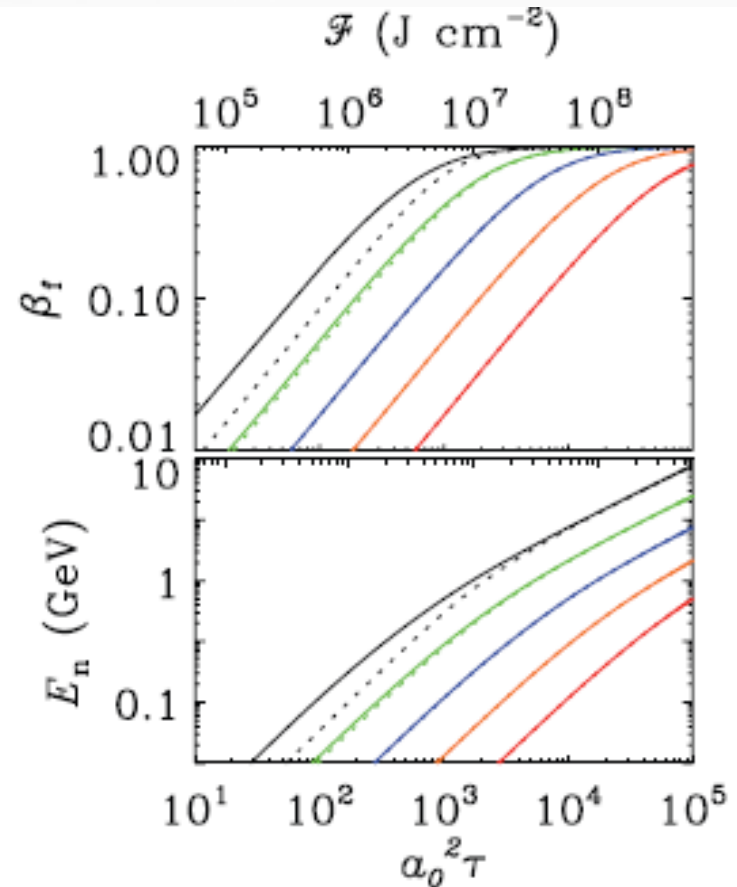
SCALING TO LASER-SOLID INTERACTIONS

Velocity and energy/nucleon for LS-RPA of a ultrathin solid target vs. laser pulse fluence \mathcal{F} for (dimensionless) surface target densities

$$\zeta = 1, 3.16, 10, 31.6, 100$$

Experimental requirements:

- nm foil targets (e.g. DLC)
- ultrahigh contrast (plasma mirrors)
- circular polarization



a_0 : dimensionless amplitude,

τ : duration in cycles



WHY CIRCULAR POLARIZATION?

Using **CP** and **normal incidence** fast electron generation is strongly **suppressed**, maximizing radiation pressure and preventing foil expansion of the foil target

Early study in “thick” targets:

Macchi et al, PRL **94** (2005) 165003

Proposal of CP-RPA of **ultrathin foils** for efficient and monoenergetic acceleration:

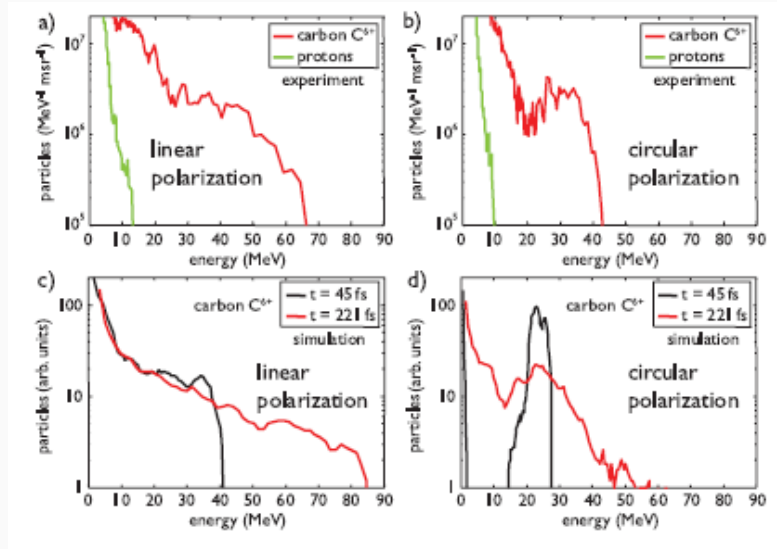
Zhang et al, PoP **14** (2007) 073101

Robinson et al, NJP **10** (2008) 013201;

Klimo et al, PRST-AB **11** (2008) 031301.

First **experimental study** reported:

Henig et al, PRL **103** (2009) 245003





THEORETICAL INTEREST IN CP-RPA...

Thick (semi-infinite) targets

("Hole Boring"):

Liseikina & Macchi, APL **94** (2007) 165003;
Naumova et al, PRL **102** (2009) 025002;
Schlegel et al, PoP **16** (2009) 083103;
Robinson et al, PPCF **51** (2009) 024004 & 095006;
Macchi & Benedetti, NIM A (2010), in press

Ultrathin (sub-wavelength) targets

("Light sail"):

Yan et al, PRL **100**, (2008) 135003 ;
Qiao et al, PRL **102** (2009) 145002;
Tripathi et al, PPCF **51** (2009) 024014;
Eliasson et al. NJP **11** (2009) 073006;
Yan et al, PRL **103** (2009) 135001;
Macchi et al, PRL **103** (2009) 085003;
Macchi et al, NJP **12** (2010) 045013.

Variations on the theme

(side effects, structured targets, ...):

Liseikina et al, PPCF **50** (2008) 124033;
Rykovanov et al., NJP **10**, (2008) 113005;
Ji et al, PRL **101** (2008) 164802;
Yin et al, PoP **15** (2008) 093106;
Holkundkara and Gupta, PoP **15** (2008)
123104;
Chen et al, PoP **15** (2008) 113103;
Zhang et al, PRST-AB **12** (2009) 021301;
Gonoskov et al, PRL **102** (2009) 145002;
Chen et al, PRL **103** (2009) 024801
Grech et al, NJP **11** (2009) 093035



LS MODEL VS 1D PIC SIMULATIONS - I

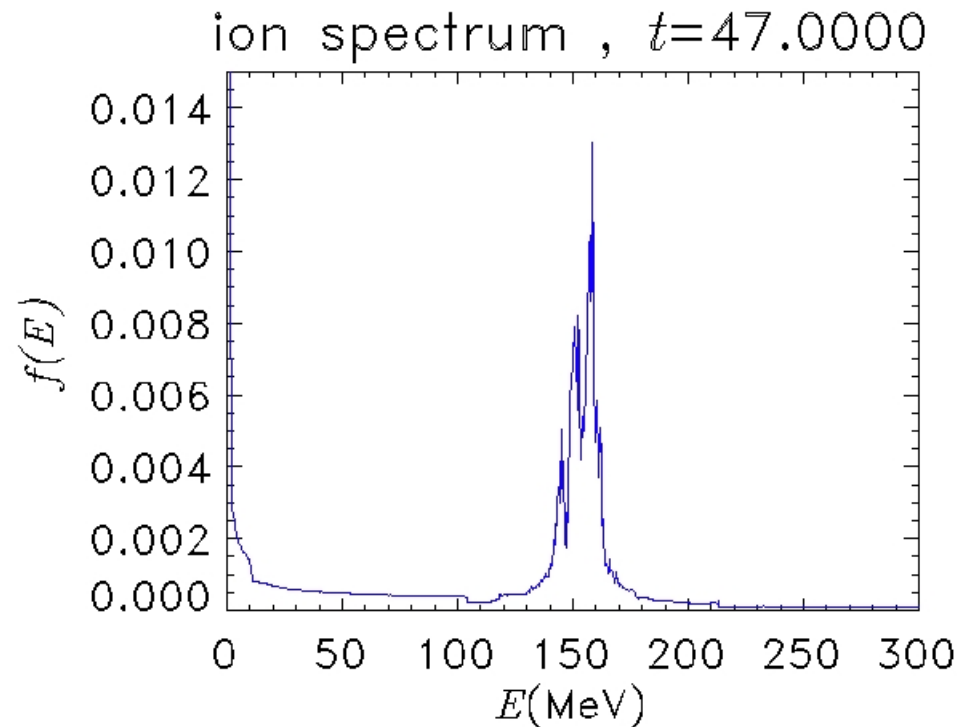
Laser pulse: $a_0 = \mathbf{5-50}$, $\tau = 8$ cycles (“flat-top” envelope)

Thin foil target: $n_e = 250n_c$, $\ell = \mathbf{0.01-0.1\lambda}$ ($\zeta = \mathbf{7.8-78.5}$)

A narrow spectral peak is observed for $a_0 < \zeta$.

The energy of the peak is in **good agreement with the LS formula**

For $a_0 > \zeta$, the dynamics is dominated by a **Coulomb explosion** of the foil

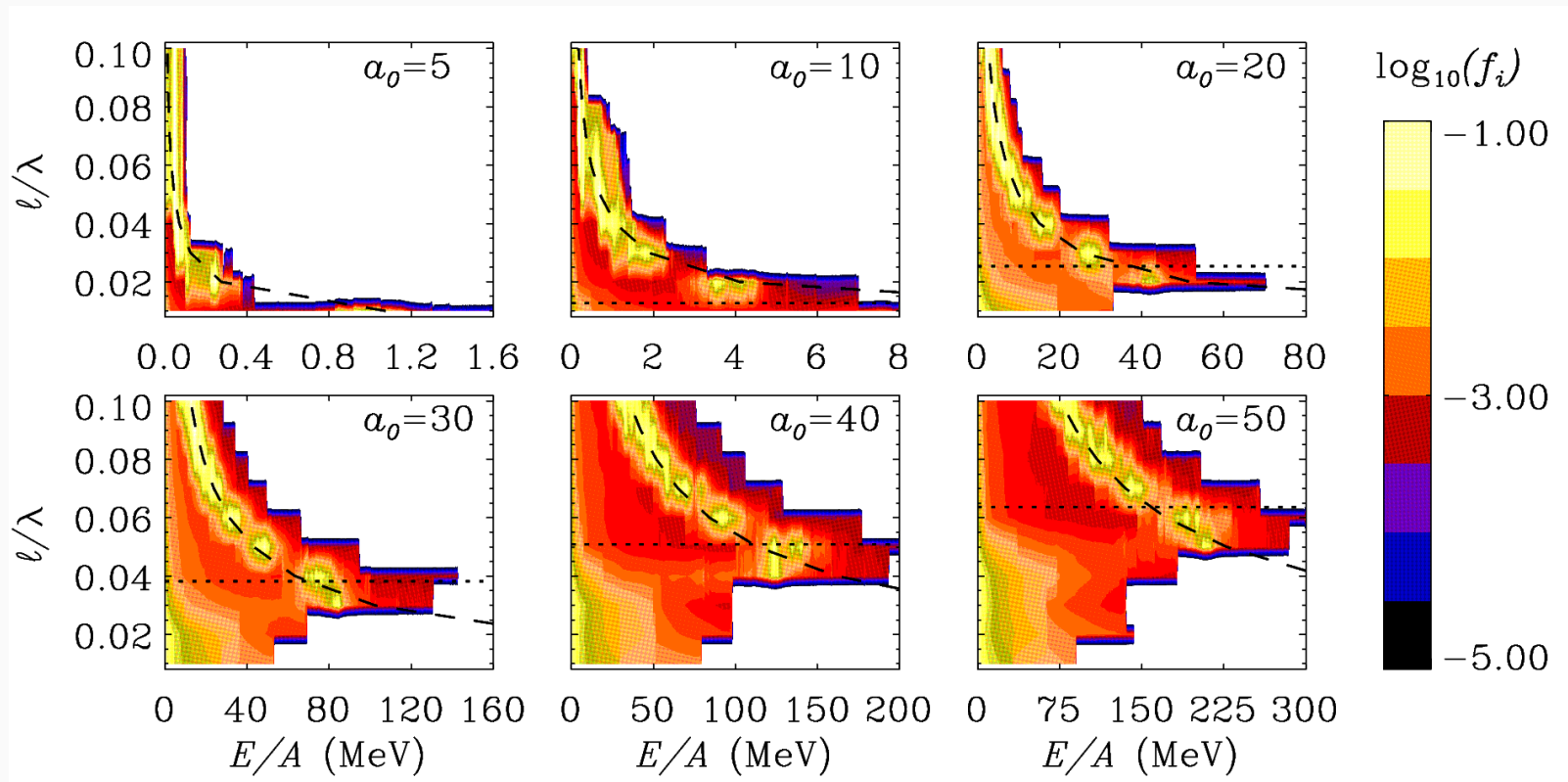




LS MODEL VS 1D PIC SIMULATIONS - II

Energy spectra vs. a_0 and ℓ :

(Dashed line: LS model prediction, dotted line: $a_0 = \zeta$)



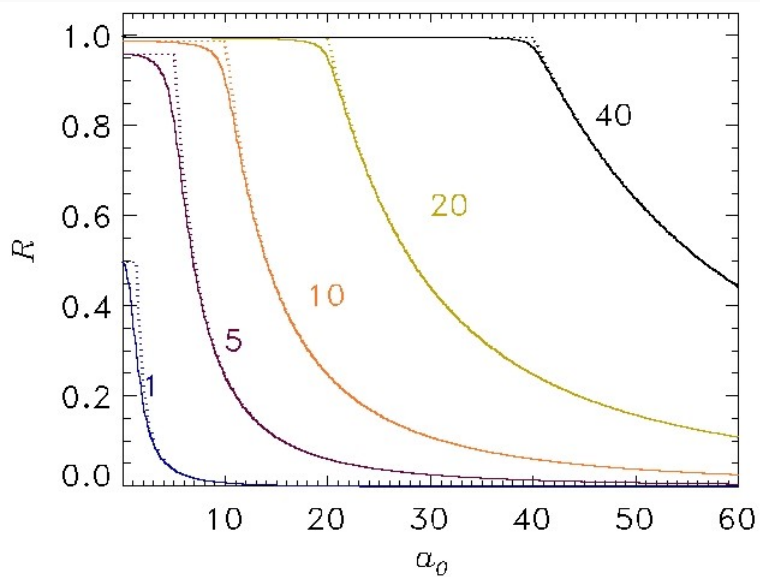


TRANSPARENCY AND “OPTIMAL” THICKNESS

Ultrathin slab model: $n_e(x) = n_0 \ell \delta(x)$, foil thickness $\ell \ll \lambda$

Total radiation pressure in rest frame $P_{\text{rad}} = (2I/c)R$

Nonlinear reflectivity $R = R(\zeta, a_0)$ includes Self-Induced Transparency



$$R \approx \zeta^2 / (\zeta^2 + 1) \quad (a_0 < \zeta)$$

$$R \approx \zeta^2 / a_0^2 \quad (a_0 > \zeta)$$

P_{rad} does not depend on

a_0 for $a_0 > \zeta$! (since $I \sim a_0^2$)

The maximum boost of the foil is at $a_0 \approx \zeta$



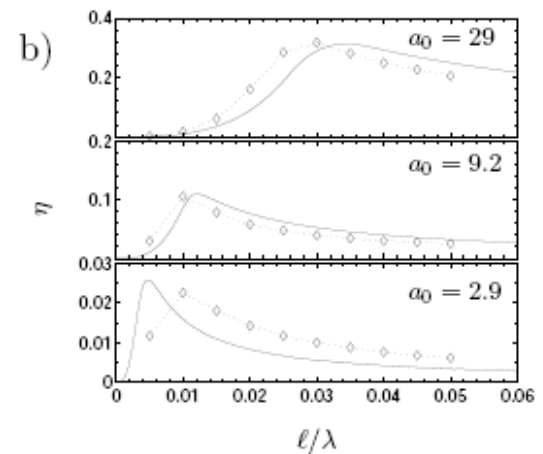
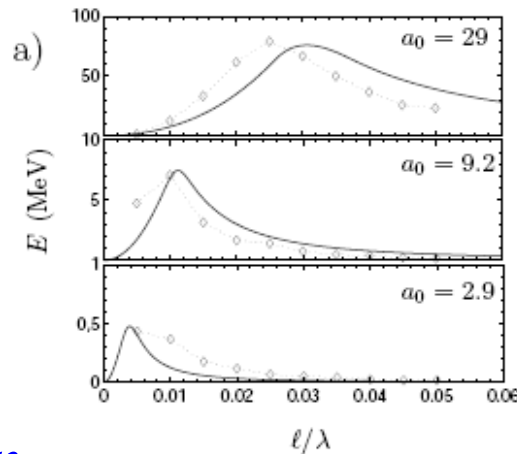
LS MODEL WITH SIT INCLUDED

Modified foil velocity formula for $R < 1$, $a_0 < \zeta$

$$\beta(t) = \frac{(1 + \mathcal{E} - \zeta^{-2})^2 + (1 + \mathcal{E} - \zeta^{-2})\sqrt{(1 + \mathcal{E} - \zeta^{-2})^2 + 4\zeta^{-2}} + 2\zeta^{-2} - 2}{(1 + \mathcal{E} - \zeta^{-2})^2 + (1 + \mathcal{E} - \zeta^{-2})\sqrt{(1 + \mathcal{E} - \zeta^{-2})^2 + 4\zeta^{-2}} + 2\zeta^{-2} + 2}$$

$$\mathcal{E} = \frac{2F(t)}{\rho l c^2} = 2\pi \frac{Z m_e a_0^2 \tau}{A m_p \zeta}$$

Ion energy and
conversion efficiency
vs. intensity and
thickness
(solid: theory,
points: PIC sims.)



9 cycles pulse, $n_e = 250 n_c$

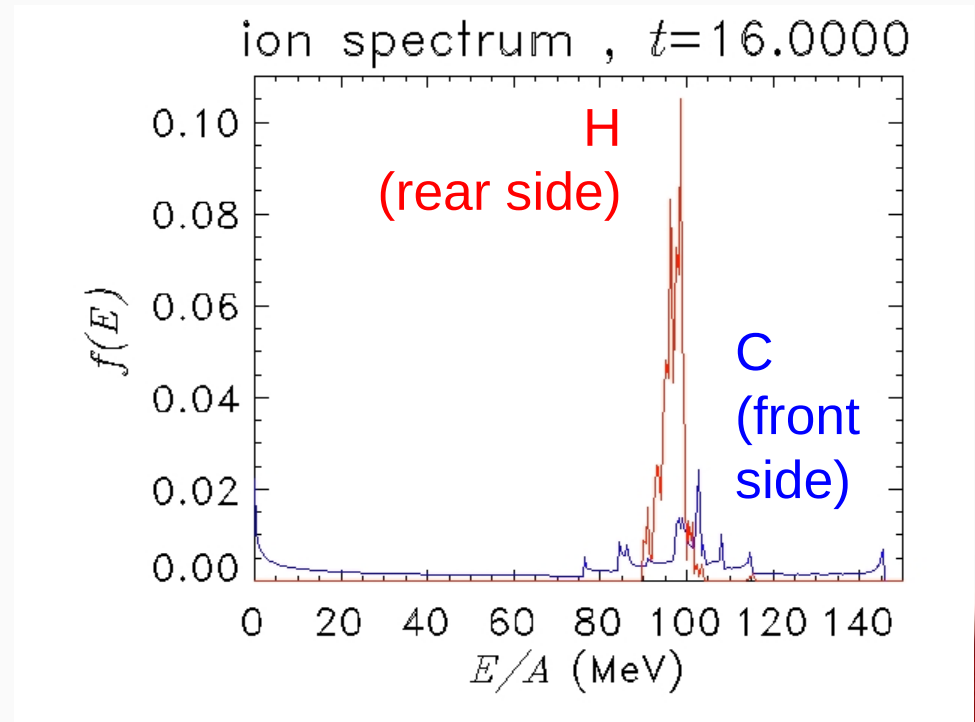


A PUZZLING ISSUE: “DARK” MASS

The RPA peak contains only
~30% of all the ions
(and ~64% of their energy)

Only the **rear side** of the foil
is accelerated (thus LS RPA
may work for double-layer
targets!)

→ *Why there is very good
Agreement of the energy with
the LS formula when inserting
there the whole mass of the target
(and not ~30% of it)?*





RADIATION VS ELECTROSTATIC PRESSURE

Radiation pressure drives electron depletion and generates back-holding electrostatic pressure

$$P_{\text{rad}} = (2I/c)R \leq P_{\text{es}} = 2\pi(en_0 \ell)^2 \quad \text{for } a_0 \leq \zeta$$

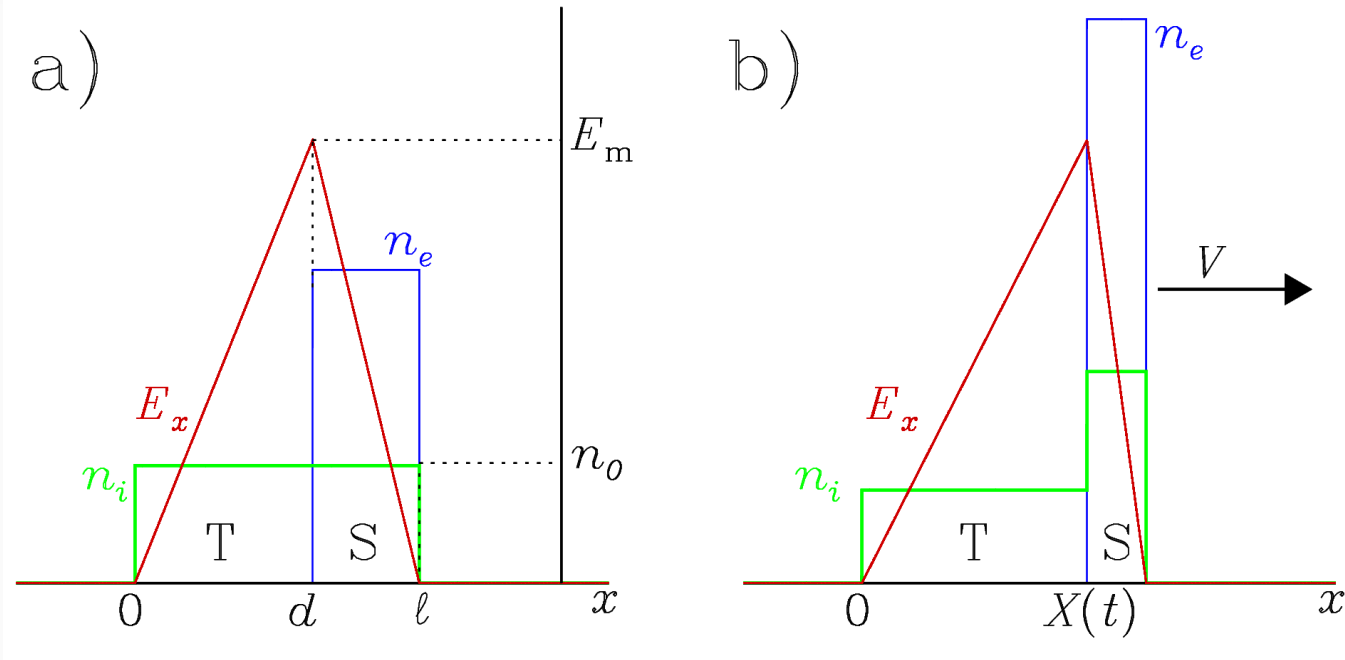
If $a_0 < \zeta$ and $\zeta \gg 1$, $R \approx 1$ and **no electrons are pushed away**

For $a_0 \rightarrow \zeta$ all electrons must pile up near the rear surface in order that $P_{\text{rad}} \simeq P_{\text{es}}$.

- the electron pile-up layer is **much thinner than the foil**
- only **a fraction of the foil is accelerated**



TWO ION POPULATIONS: TAIL & SAIL



Sail (S): ions are bunched accelerated by $E_x = f_p / e$ and move coherently as a “foil” : **monoenergetic component**

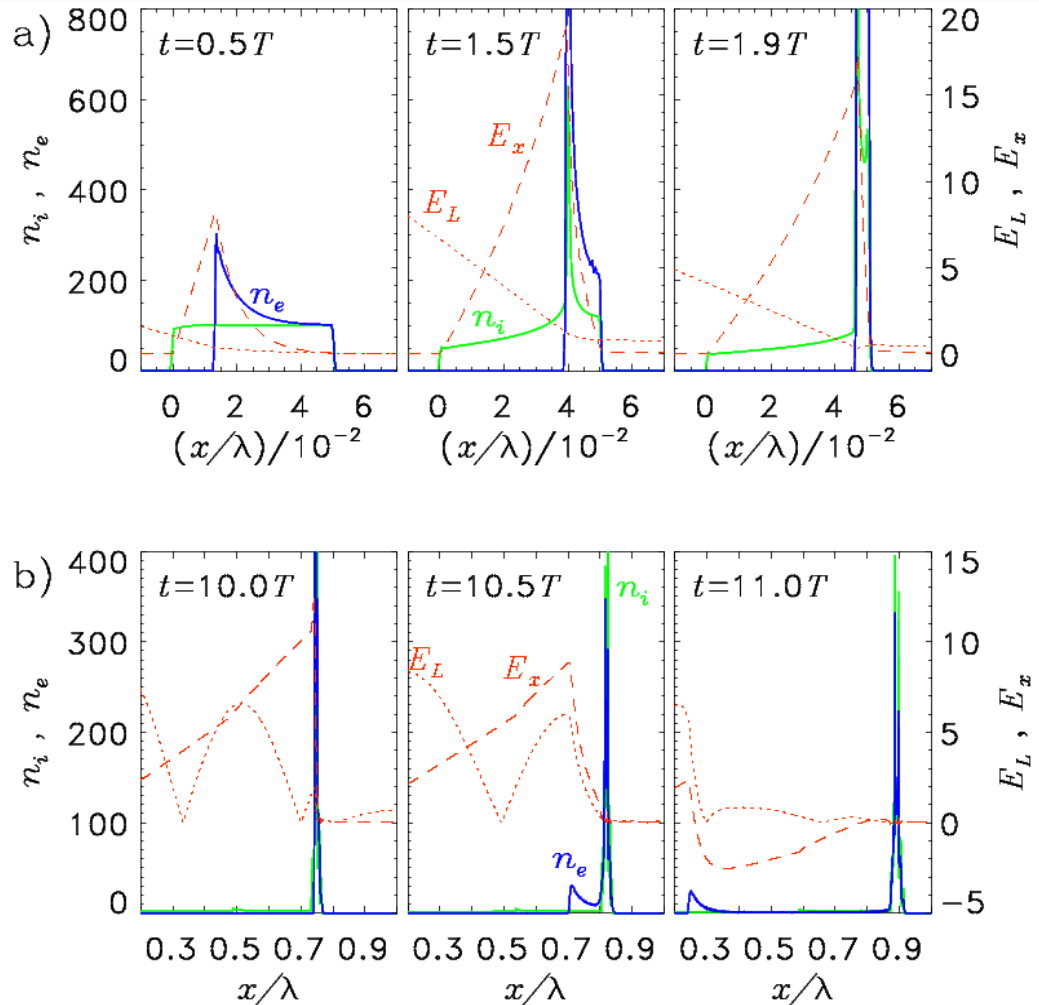
Tail (T): ions are accelerated by their own space-charge field and “Coulomb explode”: **broad spectrum component**



SAIL CHARGING/DISCHARGING

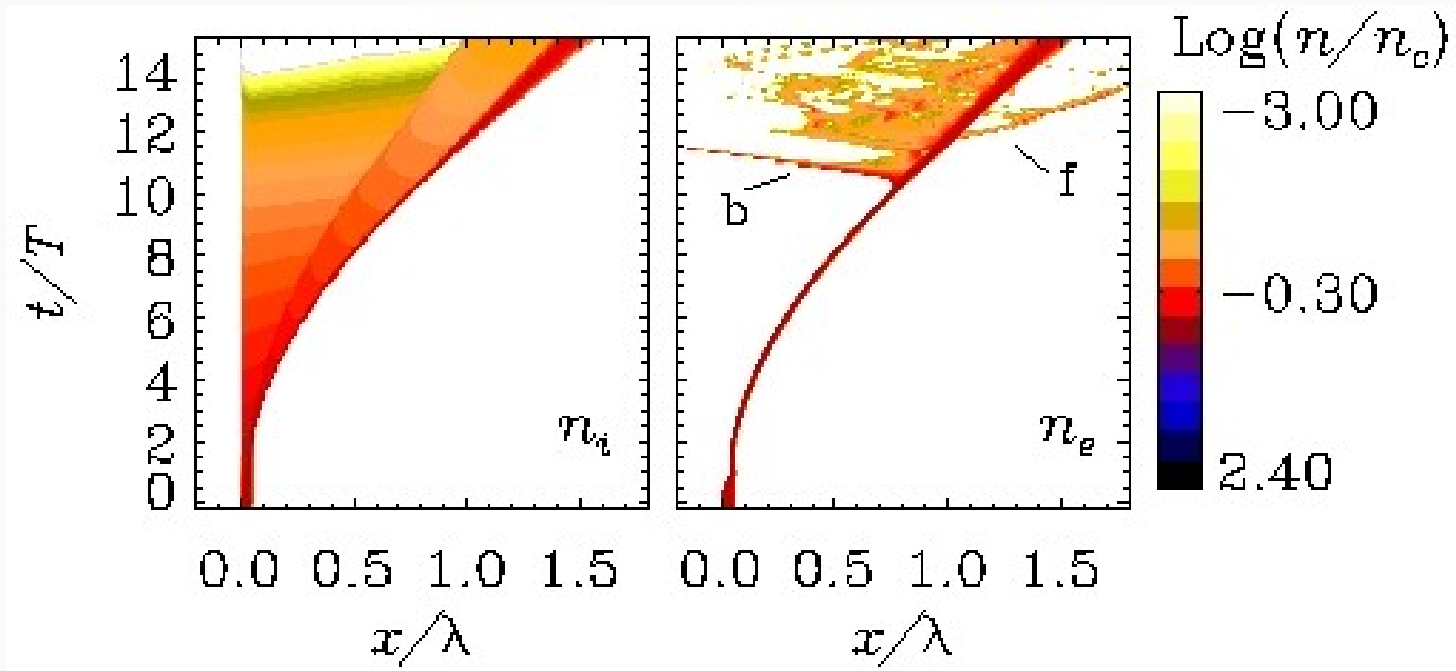
PIC simulations show ions in the compression layer to form a “Sail” thinner than the original foil and negatively charged (excess of electrons)

The excess electrons “detach” from the Sail near the end of the laser pulse, moving backwards and leaving a neutral plasma bunch





COLLAPSE OF ELECTRON EQUILIBRIUM



Near the end of the acceleration stage the pressure balance for electrons breaks down: formation of bunches in forward & backward directions and heating

Ion spectrum broadens in the post-acceleration stage



MISSING INERTIA IS LOWER PRESSURE

Equilibrium of radiation
and electrostatic
pressure on *electrons*:

$$P_{\text{rad}} \doteq \int (-e)n_e E_x dx = \int n_e f_p dx$$

Electrostatic pressure
on *ions*:

$$P_{\text{es}} = \int Z e n_i E_x dx < P_{\text{rad}} \quad (Z n_i < n_e)$$

Calculation on equilibrium
profiles yields:

$$P_{\text{es}} = \frac{M_{\text{Sail}}}{M_{\text{Foil}}} P_{\text{rad}}$$

Equation of motion:

$$P_{\text{es}} = \frac{d}{dt} (M_{\text{Sail}} \mathbf{V}) \iff P_{\text{rad}} = \frac{d}{dt} (M_{\text{Foil}} \mathbf{V})$$

→ **The Sail moves as if it had the total mass of the foil**

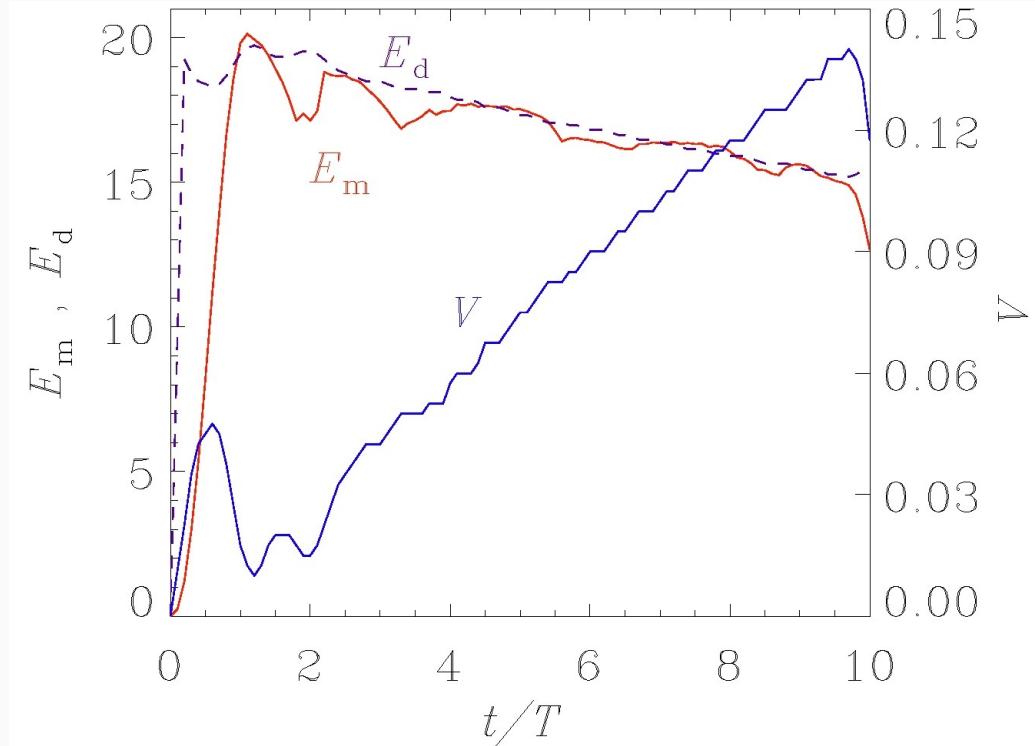


DYNAMIC PRESSURE BALANCE

P_{rad} decreases with
velocity in the Lab frame
 $(P_{\text{rad}})^L = (1-\beta)/(1+\beta) P_{\text{rad}}$

To keep pressure
equilibrium there is
a **mass flow**
(ion current) from

M_{tail} to M_{sail}





ENERGY BALANCE

Efficiency depends only on β (the Sail velocity)
BUT the kinetic energy of the Sail is less than the total!

Energy stored in the
electrostatic field E_x :

“Conversion efficiency”
into electrostatic
energy η_{es} :

$$U_{es} = U_{es}(t) = \int_0^{X(t)} \frac{E_x^2(x, t)}{8\pi} dx$$
$$\frac{dU_{es}}{dt} = \frac{1}{8\pi} E_x^2[X(t), t] \frac{dX}{dt} = \frac{1}{8\pi} E_0^2 \beta c$$
$$\eta_{es} = \frac{1}{I} \frac{dU_{es}}{dt} = 2\beta \left(\frac{d}{\ell}\right)^2 \left(\frac{\zeta}{a_0}\right)^2$$

For $a_0 = \zeta$, the depletion width $d \approx \ell$ thus $\eta_{es} \approx 2\beta$:

most of the stored energy is converted into electrostatic energy
and eventually goes to Tail ions

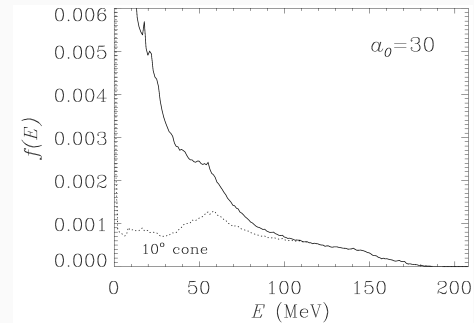
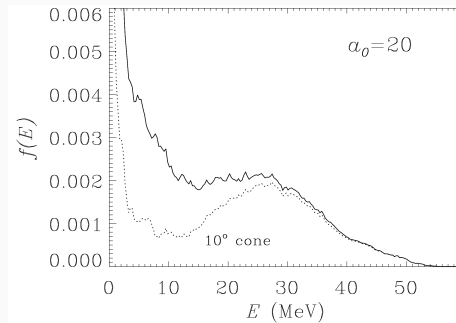
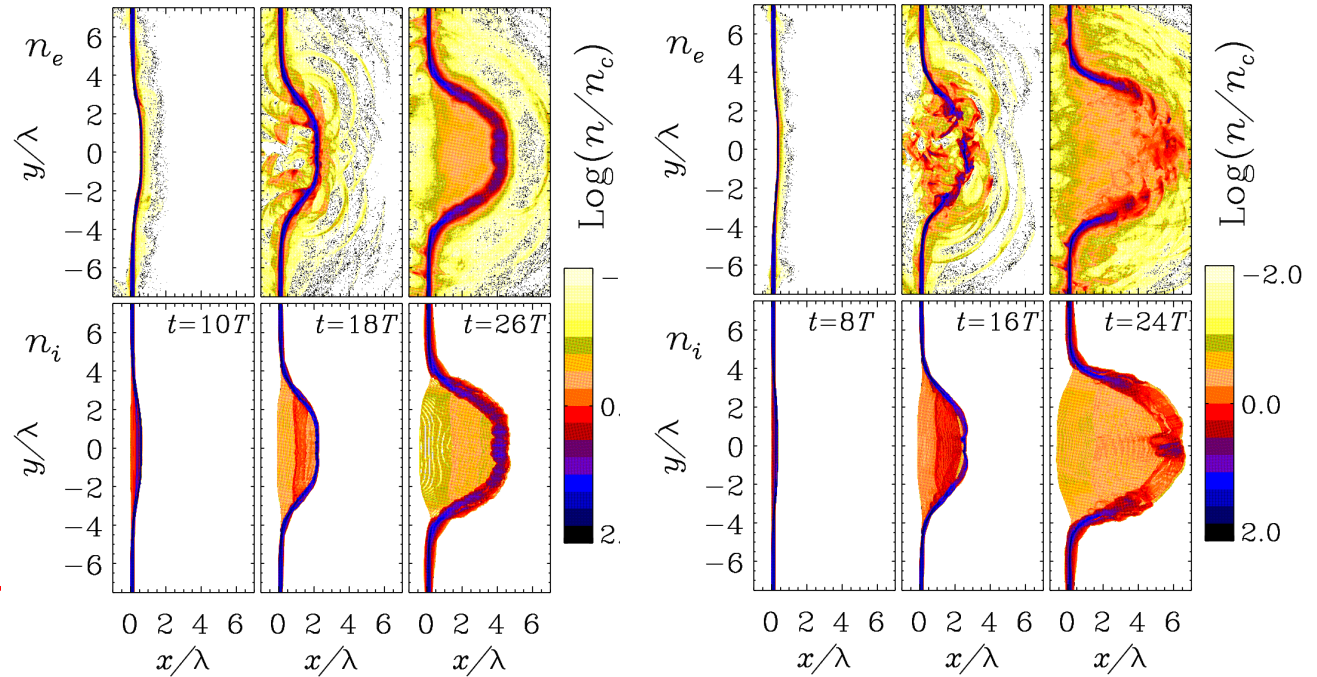


TWO-DIMENSIONAL SIMULATIONS

2D sims for
 $\zeta=31.4$ and
 $a_0=20$ (left)

$a_0=30$ (right)

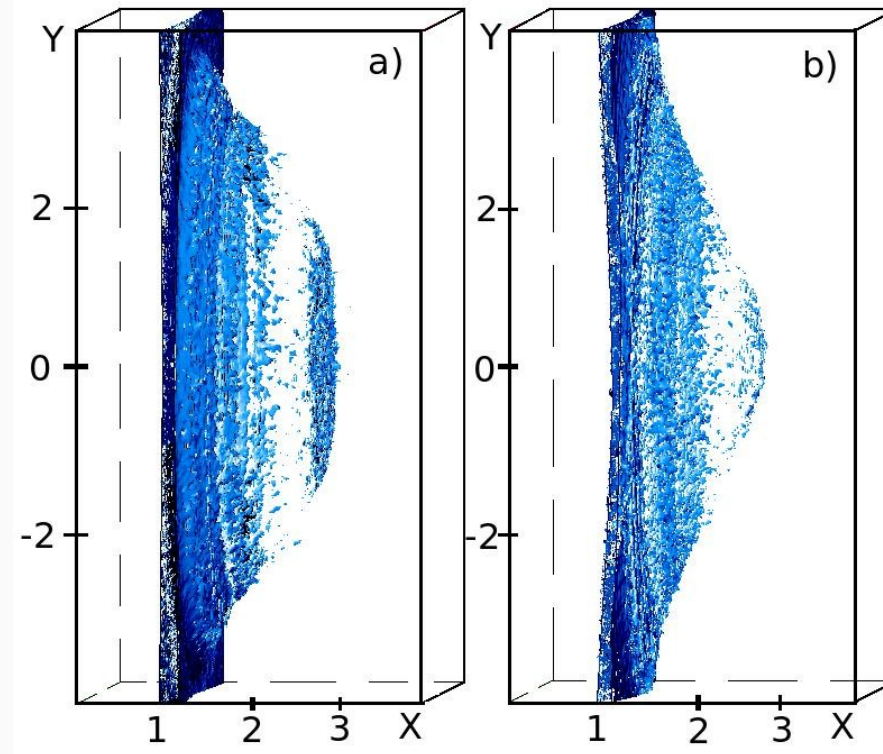
stronger electron
heating and lower
“penetration”
threshold with
respect to 1D:
ion spectrum is
broad





THREE-DIMENSIONAL SIMULATIONS

3D sims for $\zeta=15$,
 $a_0 = 5$, $\tau=18$ cycles
left: Supergaussian
spot profile
right: Gaussian



Supergaussian “flat-top” profiles keep a “quasi-1D” geometry and prevent early breakthrough of laser pulse due to lateral expansion



RADIATION REACTION EFFECTS

Motivation: RR is important for ultra-relativistic particles in EM fields and is thus expected to play a strong role in next generation experiments at ultra-high intensities.

The typical intensity for relevant RR effects is estimated to be $\sim 10^{23}$ W/cm². This corresponds, e.g., to the foreseen regime of RPA dominance (even for Linear Polarization)
[Esirkepov et al, PRL **92** (2004) 175003]

Two RPA simulation studies (for CP at lower intensity) suggest a “beneficial” effect of “electron cooling” by RR
[Schlegel et al, PoP **16** (2009) 083103;
Chen et al, arXiv:0909.5144]



RADIATION REACTION FORCES

$$mc \frac{du^\mu}{d\tau} = f_L^\mu + f_{RR1}^\mu + f_{RR2}^\mu + f_S^\mu$$

EoM of classical particle
with spin in EM field

$$f_L^\mu = eF^{\mu\nu} u_\nu$$

$$f_{RR1}^\mu = e\tau_0 (\partial_\alpha F^{\mu\nu} u_\nu u^\alpha)$$

$$f_{RR2}^\mu = \frac{e^2}{mc} \tau_0 \left(F^{\mu\nu} F_{\nu\alpha} u^\alpha + (F^{\nu\beta} u_\beta F_{\nu\alpha} u^\alpha) u^\mu \right)$$

$$f_S^\mu = -\frac{1}{2c} S^{\gamma\delta} \partial^\mu F_{\gamma\delta} + \frac{1}{2c} \left(S^{\gamma\delta} \partial_\alpha F_{\gamma\delta} u^\alpha \right) u^\mu$$

$$u^\alpha = \left(\gamma, \gamma \frac{\mathbf{v}}{c} \right) \quad \partial^\mu \equiv \left(\frac{\partial}{\partial t}, -c\nabla \right)$$



RADIATION REACTION MODELING

EoM with Landau-Lifshitz force in non-covariant notation

$$\begin{aligned} \frac{d\mathbf{p}}{dt} = & -\left(\mathbf{E} + \mathbf{v} \times \mathbf{B}\right) \\ & - \left(\frac{4}{3}\pi\frac{r_e}{\lambda}\right) \gamma \left[\left(\frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla\right) \mathbf{E} + \mathbf{v} \times \left(\frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla\right) \mathbf{B} \right] \\ & + \left(\frac{4}{3}\pi\frac{r_e}{\lambda}\right) \left[\left(\mathbf{E} + \mathbf{v} \times \mathbf{B}\right) \times \mathbf{B} + \left(\mathbf{v} \cdot \mathbf{E}\right) \mathbf{E} \right] \\ & - \left(\frac{4}{3}\pi\frac{r_e}{\lambda}\right) \gamma^2 \left[\left(\mathbf{E} + \mathbf{v} \times \mathbf{B}\right)^2 - \left(\mathbf{v} \cdot \mathbf{E}\right)^2 \right] \mathbf{v} \end{aligned}$$

The last “friction” term is the dominant one
(first two terms are smaller than spin contribution)

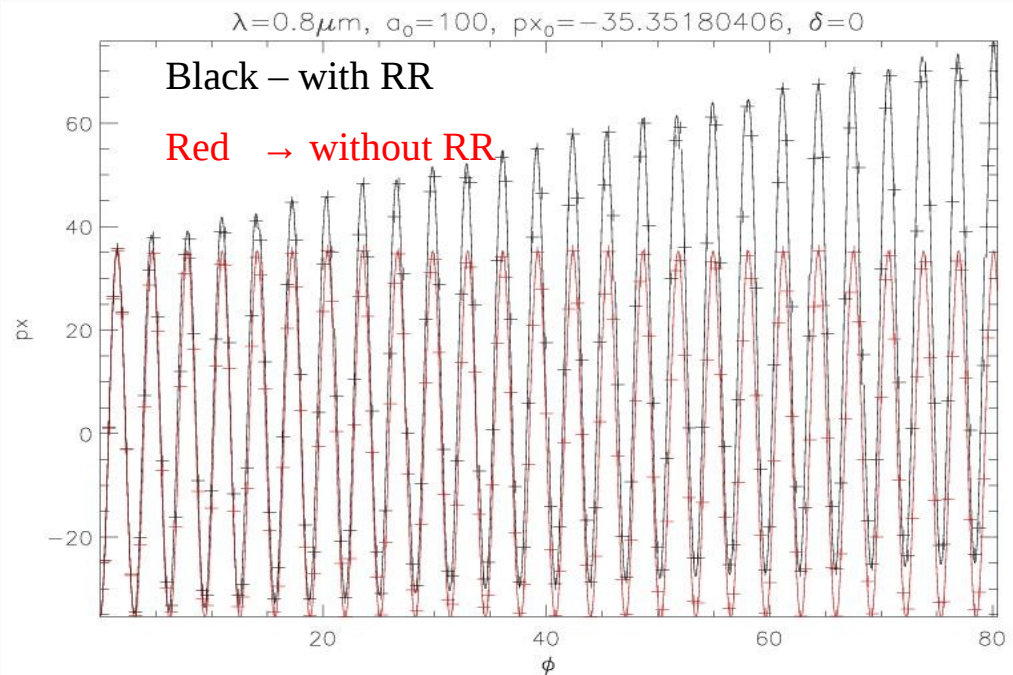


BENCHMARK WITH EXACT SOLUTIONS

Exact solution of the Landau-Lifshitz equation in a plane wave
A.Di Piazza, Lett.Math.Phys. **83** (2008) 305

Test of particle pusher
algorithm with “reduced” LL
included:

- excellent agreement for intensities up to 10^{24} W/cm²
- straightforward to include in a “standard” PIC code (based on Boris pusher)
- only ~10% increase in CPU time





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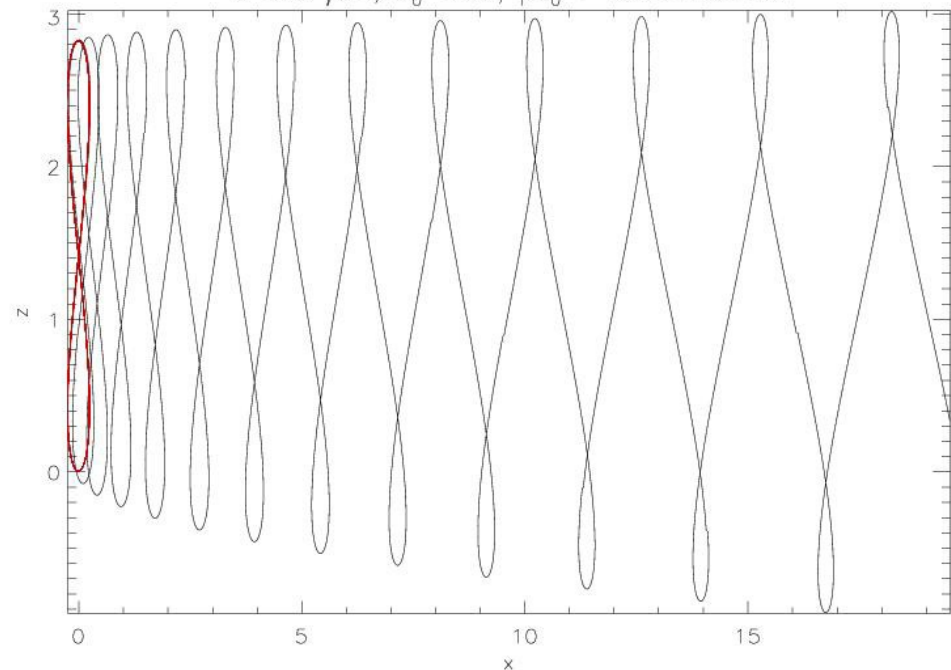
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Black – with RR

Red → without RR

$\lambda=0.8 \mu\text{m}$, $\alpha_0=100$, $p_{x0}=-35.35180406$



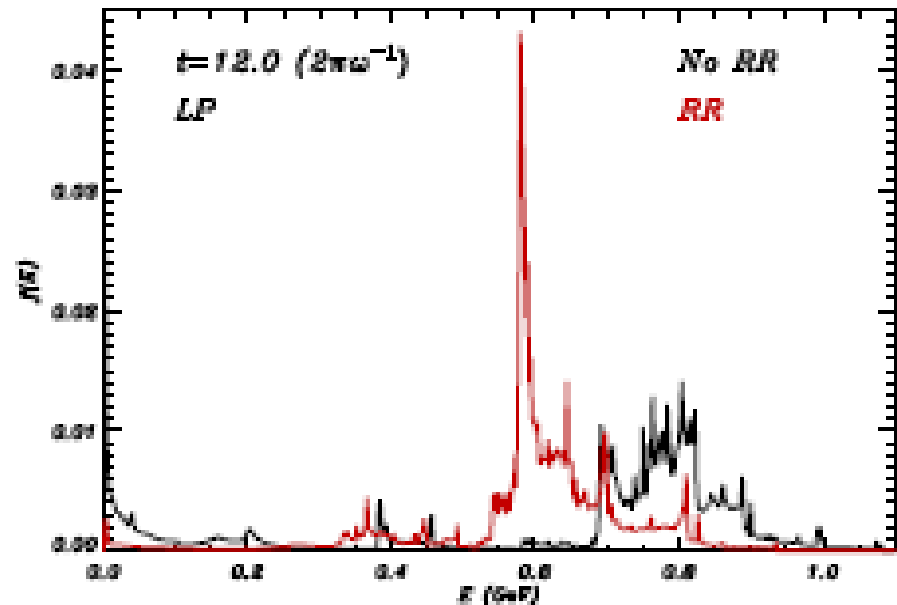


RR EFFECTS ON ION SPECTRA – I (LP)

RP-dominated regime: 10^{24} W/cm² , 11 cycles pulse
1 μ m foil, $100n_c$, linear polarization

Lower energy,
narrower spectrum
with RR included

~20% energy
“dissipated” by RR





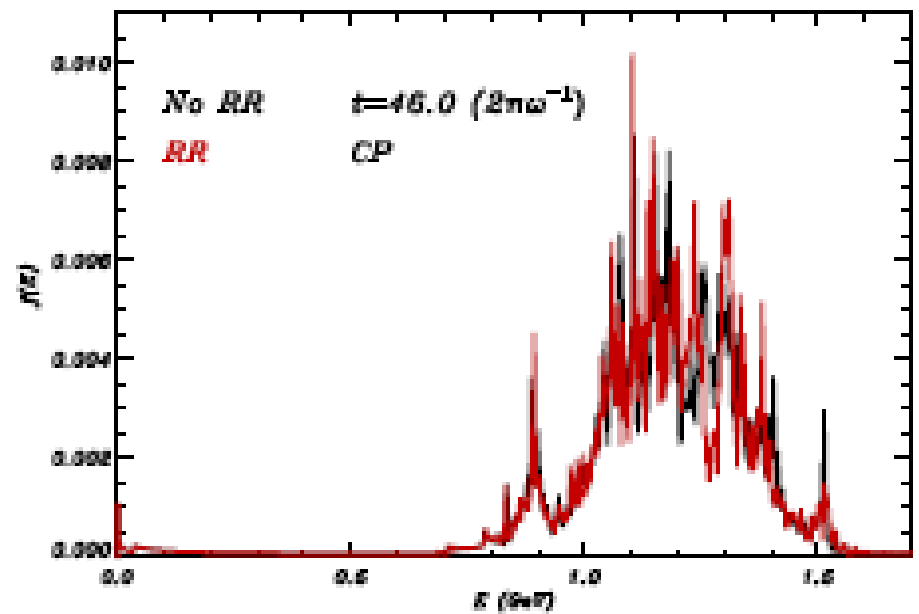
RR EFFECTS ON ION SPECTRA – II (CP)

RP-dominated regime: 10^{24} W/cm² , 11 cycles pulse

1 μ m foil, $100n_c$, circular polarization

Negligible RR effects
on ion spectrum!

Higher energy than in
LP case





CONCLUSIONS

- The simple Light Sail model correctly predicts the ion energy and conversion efficiency observed in simulations
- Not all the ions are accelerated as a monoenergetic bunch
- Dynamics and self-organization underlying LS has been unfolded
- Keeping a monoenergetic spectrum is difficult
- Radiation Reaction effects at ultrahigh intensities are important only for Linear Polarization

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