

From Intense Fields to Radiation Dominated Regime

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γ -resist kick-off meeting,
November 28, 2012, CNR/INO, Pisa

Outline of the talk

- ▶ A “soft” introduction to Radiation Friction
- ▶ Landau-Lifshitz approach
- ▶ Radiation Friction effects in Laser-Plasma interactions
- ▶ The Radiation Dominated Regime
- ▶ Radiation Friction signatures in Thomson Scattering

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Introducing Radiation Friction

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Pedagogical example:
electron in a magnetic field \mathbf{B}_0

$$\mathbf{f}_L = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B}/c) \text{ Lorentz force}$$

$$m_e \frac{d\mathbf{v}}{dt} = \mathbf{f}_L = -\frac{e}{c} \mathbf{v} \times \mathbf{B}_0$$

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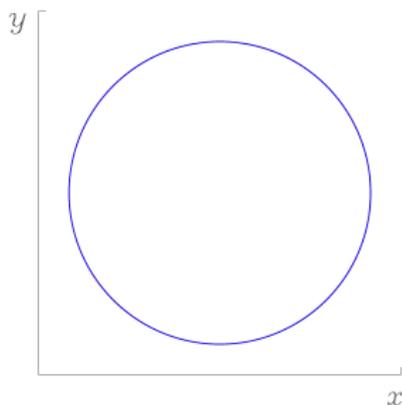
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Solution: uniform circular motion

$$|\mathbf{v}| = v = \text{const.}$$

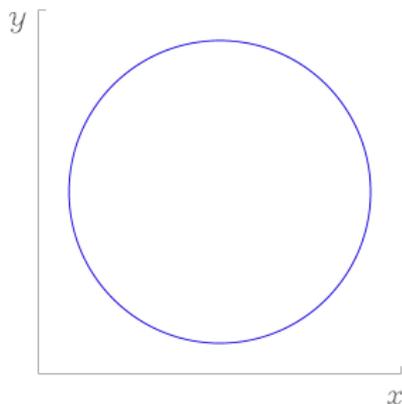
$$K = \frac{1}{2} m_e v^2 = \text{const.} \quad \omega_c = \frac{eB_0}{m_e c} \quad r = \frac{v}{\omega_c}$$



Introducing Radiation Friction

BUT the electron radiates:

$$P_{\text{rad}} = \frac{2e^2}{3c^3} \left| \frac{d\mathbf{v}}{dt} \right|^2 = \frac{2e^2}{3c^3} \omega_c^2 v^2$$



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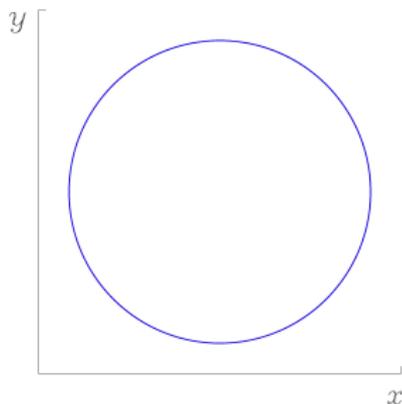
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Energy loss due to radiation:

$$\frac{dK}{dt} = -P_{\text{rad}} \longrightarrow v(t) = v(0)e^{-t/\tau}$$

$$\tau = \frac{3m_e c^3}{2e^2 \omega_c^2} = \frac{3c}{2r_c \omega_c^2}$$



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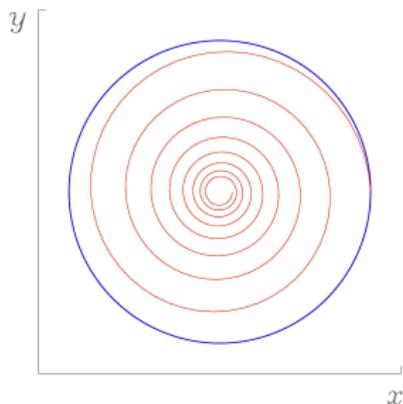
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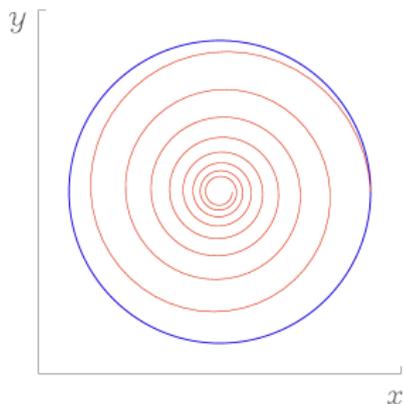
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If $r(t) \simeq v(t)/\omega_c$, electron “falls” along a spiral



Introducing Radiation Friction

The Lorentz force does not describe the electron motion consistently:
need to include an extra force



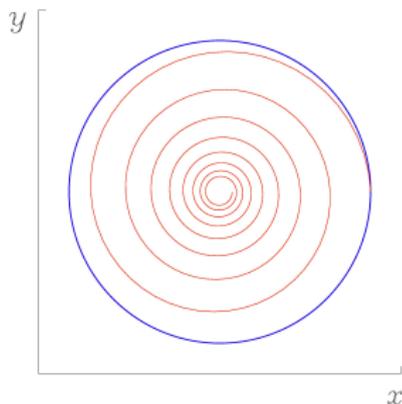
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$$m_e \frac{d\mathbf{v}}{dt} = \mathbf{f}_L + \mathbf{f}_{\text{rad}}$$

Work done by extra force = energy loss

$$\int_0^t \mathbf{f}_{\text{rad}} \cdot \mathbf{v} dt = - \int_0^t P_{\text{rad}} dt \longrightarrow \mathbf{f}_{\text{rad}} = - \frac{2e^2}{3c^3} \frac{d^2\mathbf{v}}{dt^2}$$



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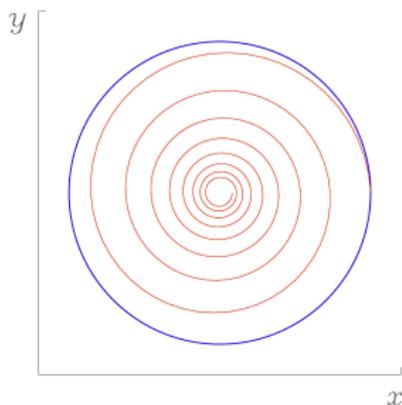
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Physical interpretation: the electron is affected by the self-generated radiation field (radiation *reaction* or *self-force*)



Landau-Lifshitz approach

$\mathbf{f}_{\text{rad}} = -\frac{2e^2}{3c^3} \frac{d^2 \mathbf{v}}{dt^2}$ is unsatisfying:

- unphysical runaway solutions $\dot{\mathbf{v}}(t) = \dot{\mathbf{v}}(0)e^{t/\tau}$
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LL iterative approach brings $\mathbf{f}_{\text{rad}} = \mathbf{f}_{\text{rad}}(\mathbf{E}, \mathbf{B})$:

$$\mathbf{f}_{\text{rad}} \simeq -\frac{2e^2}{3c^3} \left(-\frac{e}{m_e} \frac{d}{dt} \mathbf{f}_L \right) = \frac{2e^3}{3m_e c^3} \left(\dot{\mathbf{E}} - \frac{e}{m_e c} \mathbf{E} \times \mathbf{B} \right)$$

in the “instantaneous” frame where $\mathbf{v} = 0$

L.L.Landau, E.M.Lifshitz, *The Classical Theory of Fields*
(Elsevier, 1975), 2nd Ed., par.76

Validity range of classical LL approach

Iterative approach valid if $|\mathbf{f}_{\text{rad}}| \ll |e\mathbf{E}|$ in the instantaneous frame

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$$\frac{c}{\omega} \gg r_c \equiv \frac{e^2}{m_e c^2} = 2.8 \times 10^{-13} \text{ cm}$$

If $|\dot{\mathbf{E}}| \sim \omega E$:

$$B \ll \frac{m_e c^2}{e r_c} = 6 \times 10^{15} \text{ G} \rightarrow E \ll 2 \times 10^{18} \text{ V cm}^{-1}$$

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$$E < E_s = \frac{m_e c^2}{\lambda_c} = \frac{m_e^2 c^3}{e \hbar} < \frac{m_e c^2}{r_c} \quad \frac{c}{\omega} > \lambda_c = \frac{\hbar}{m_e c}$$

$$E_s = 1.3 \times 10^{16} \text{ V cm}^{-1} \text{ (Schwinger field)}$$

$$\lambda_c = 4 \times 10^{-11} \text{ cm (Compton wavelength)}$$

Why worry about Radiation Friction?

The relevant fields seem out of reach, but . . .

- ▶ Depending on the interaction geometry the field amplitudes and frequencies are much higher in the rest frame of the electron

Example: collision of an electron with $\gamma \gg 1$ and a plane wave

$$F = \frac{2}{3} \left(\frac{e^2}{m_e c^2} \right) |\mathbf{E} \times \mathbf{B}| = \frac{8\pi}{3} r_c^2 I \longrightarrow F' = \frac{8\pi}{3} r_c^2 (4\gamma^2 I) \gg F$$

- ▶ The effect of radiation friction “cumulates” with time
- ▶ From another point of view: so far RF effects has *never* been characterized in experiments despite >100 years of theoretical work!

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Relativistic Landau-Lifshitz RF force

$$f_{\text{rad}}^{\mu} = -\frac{2r_c^2}{3} \left[F^{\mu\nu} F_{\alpha\nu} u^{\alpha} - F^{\alpha\nu} u_{\nu} F_{\alpha\beta} u^{\beta} u^{\mu} + \frac{m_e c}{e} \partial_{\alpha} F^{\mu\nu} u^{\alpha} u_{\nu} \right]$$

Spatial component in the laboratory frame:

$$\mathbf{f}_{\text{rad}} = -\frac{2r_c^2}{3} \left\{ \gamma^2 \left[\left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right)^2 - \left(\frac{\mathbf{v}}{c} \cdot \mathbf{E} \right)^2 \right] \frac{\mathbf{v}}{c} - \left[\left(\mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right) \times \mathbf{B} + \left(\frac{\mathbf{v}}{c} \cdot \mathbf{E} \right) \mathbf{E} \right] + \gamma \frac{m_e c}{e} \left(\dot{\mathbf{E}} + \frac{\mathbf{v}}{c} \times \dot{\mathbf{B}} \right) \right\}$$

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Inclusion of RR in plasma modeling and simulation

Kinetic equation for electrons

$$\partial_t f + \nabla_{\mathbf{r}} \cdot (\mathbf{v}f) + \nabla_{\mathbf{p}} \cdot (\mathbf{F}f) = 0 \quad f = f(\mathbf{r}, \mathbf{p}, t) \quad (*)$$

$$\mathbf{F} = -e(\mathbf{E} + \mathbf{v} \times \mathbf{B}/c) + \mathbf{f}_{\text{rad}}$$

- Eq.(*) is solved by the Particle-In-Cell approach
- Including RF accounts for loss of incoherent radiation of wavelength $\lambda \ll n_e^{-1/3}$: system becomes dissipative

[see e.g. Tamburini, Pegoraro, Di Piazza, Keitel, Macchi, New J. Phys. **12**, 123005 (2010)]

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Radiation energetics with RF included

1D PIC simulations of
laser-plasma interaction
with radiation emission
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[Capdessus et al.,
PRE **86**, 036401 (2012)]

$$n_e = 10n_c, d = 100\lambda, \tau_L = 16T$$

$$1e21 - 1e22 - 8e22 - 3e23 \text{ W cm}^{-2}$$

With RF (left): energy balance is consistent

Without RF (right): radiative loss \simeq input laser energy
for the highest intensity!

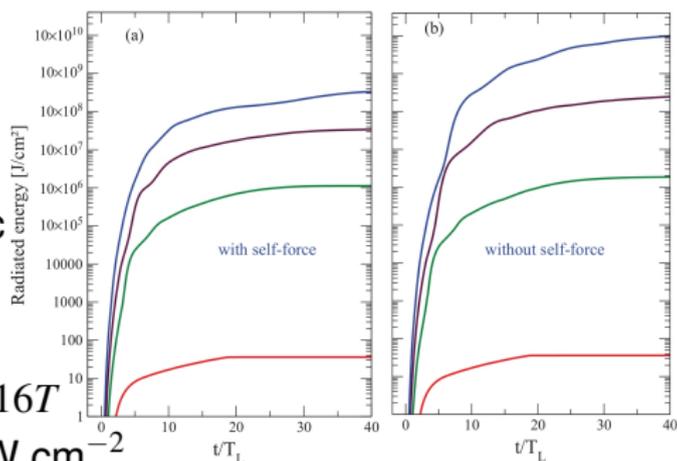
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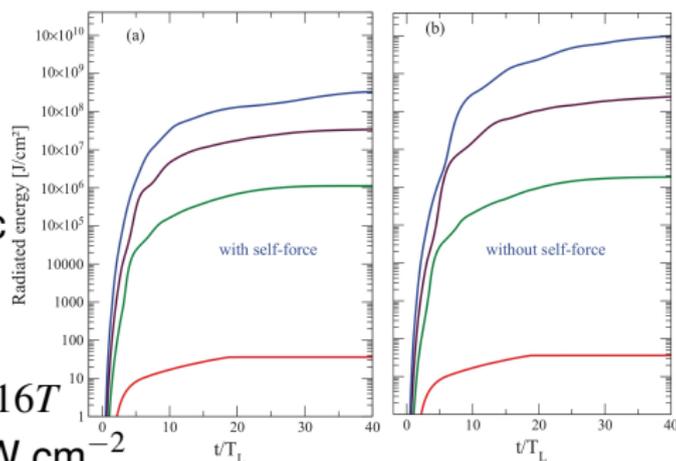
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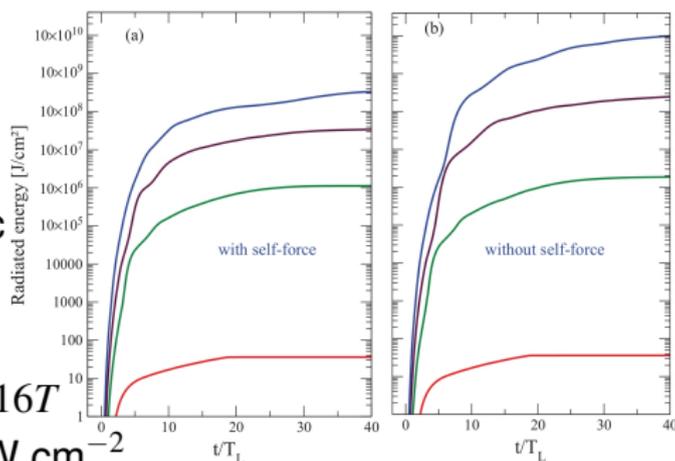
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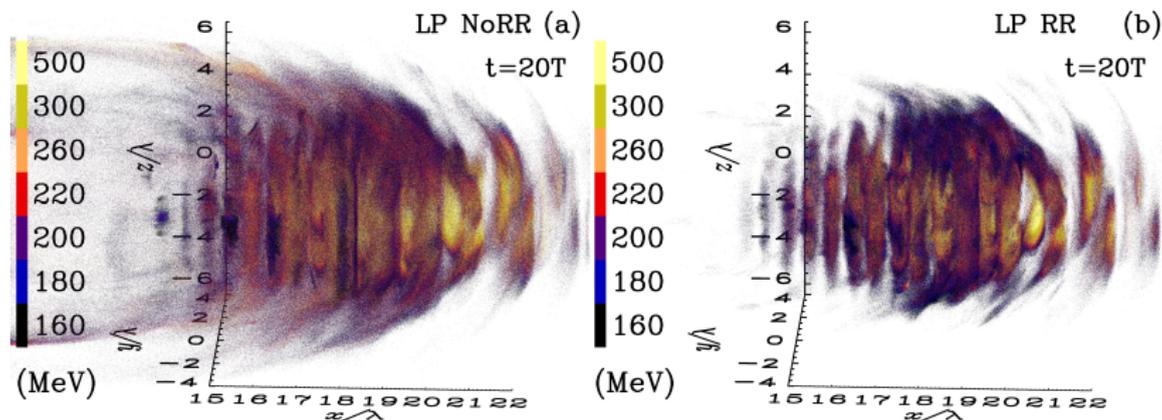
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3D PIC simulations with RF included



Thin foil acceleration ($d = 1\lambda$, $n_e = 64n_c$, $I = 1.7 \times 10^{23} \text{ W cm}^{-2}$)
Space-Energy electron distribution without (left) & with (right) RF
[Tamburini, Liseykina, Pegoraro, Macchi, PRE **85**, 016407 (2012)]

Radiation Dominated Regime

[J.Koga, T.Esirkepov, S.V. Bulanov, PoP **12**, 093106 (2005)]

RDR in a laser field: (radiation loss) \simeq (initial energy)

Electron counterpropagating to laser field $\Rightarrow \mathbf{f}_{\text{rad}}$ is maximized

\Rightarrow Thomson scattering geometry “enhances” RF effects

$$P_{\text{rad}} \frac{2\pi}{\omega} \simeq \mathcal{E}_{\text{osc}} = m_e c^2 \left[\left(1 + \frac{\mathbf{p}^2}{m_e c^2} \right)^{1/2} - 1 \right]$$

$$R \equiv \frac{2r_c \omega}{3c} \gamma_0 (1 + \beta_0) a^2 \simeq 1 \quad a = \frac{eE}{m_e \omega c}$$

($\beta_0 = v_0/c$, $\gamma_0 = (1 - \beta_0^2)^{-1/2}$ initial β - and γ -factor)

$R = 1$ for $\gamma_0 = 300$ (150 MeV) and $a = 336$ (2.4×10^{23} W cm $^{-2}$)

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[J.Koga, T.Esirkepov, S.V. Bulanov, PoP **12**, 093106 (2005)]

RDR in a laser field: (radiation loss) \simeq (initial energy)

Electron counterpropagating to laser field $\Rightarrow \mathbf{f}_{\text{rad}}$ is maximized

\Rightarrow Thomson scattering geometry “enhances” RF effects

$$P_{\text{rad}} \frac{2\pi}{\omega} \simeq \mathcal{E}_{\text{osc}} = m_e c^2 \left[\left(1 + \frac{\mathbf{p}^2}{m_e c^2} \right)^{1/2} - 1 \right]$$

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($\beta_0 = v_0/c$, $\gamma_0 = (1 - \beta_0^2)^{-1/2}$ initial β - and γ -factor)

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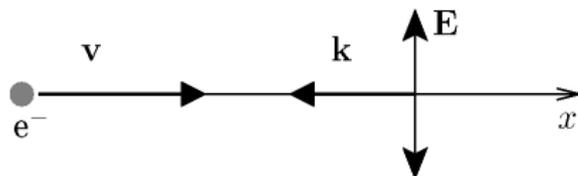
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Thomson Scattering with RF included

TS of a 35 fs pulse by 150 MeV electrons

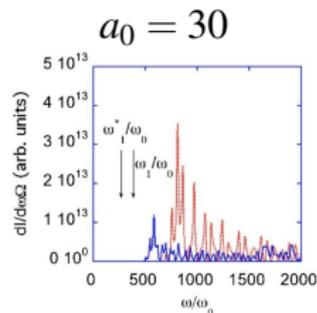


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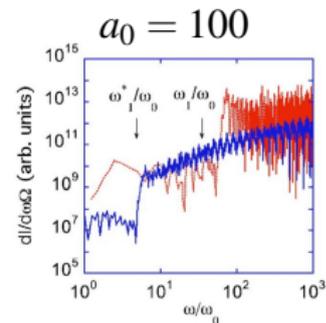
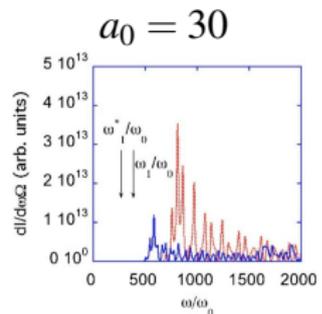
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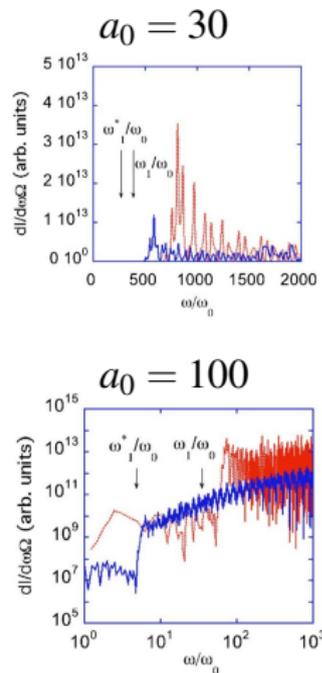
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Total power and low-frequency cut-off
are strongly affected by RF
already when $R \ll 1$

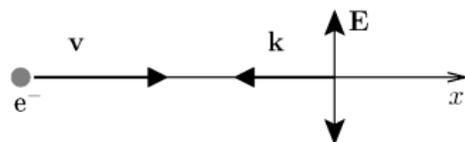


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Angular RF signatures in Thomson Scattering

Key point: the RF decreases v_x
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Change in v_x affects
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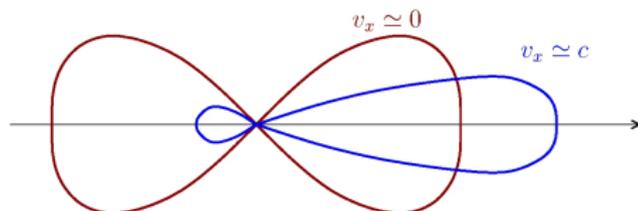
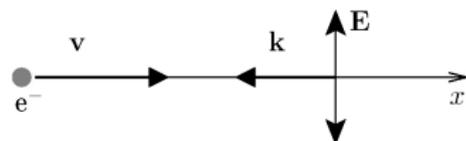
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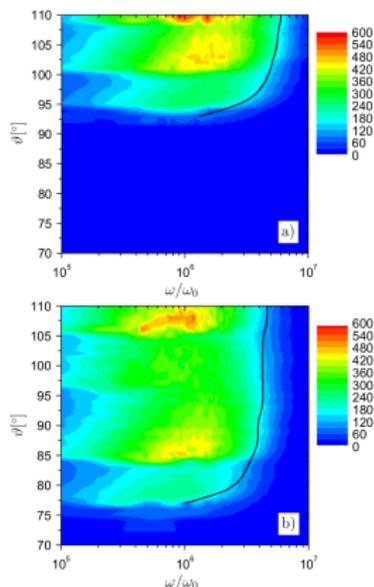
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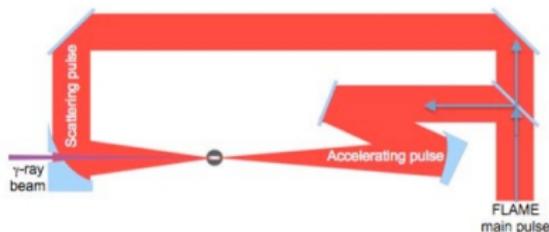
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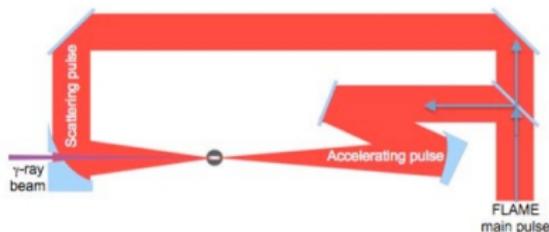
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“All-Optical” TS as a test bed for RF?



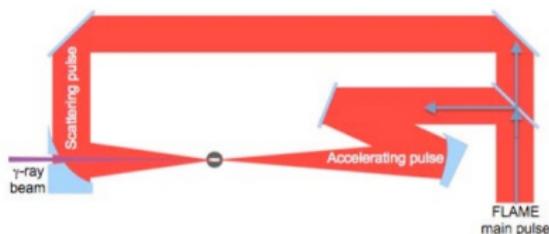
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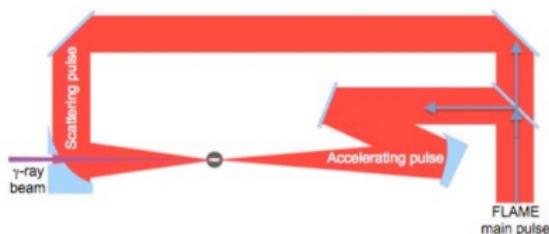
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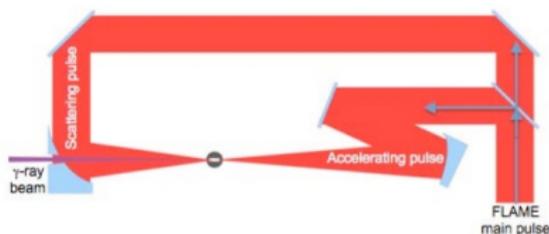
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