REAZIONI NUCLEARI DI INTERESSE ASTROFISICO

Laura Elisa Marcucci (Università di Pisa e INFN-Pisa) Riassunto del corso:

 teoria del momento angolare → costruzione delle funzioni d'onda

- <u>teoria perturbativa dipendente dal tempo</u>: calcolo della sezione d'urto (e fattore astrofisico)
- <u>funzioni d'onda</u>: stati legati (*d*, ³He, ³H) e stati di scattering (*pp*, *np*, *Nd*) anche con potenziali realistici
- <u>corrente nucleare</u>: termine one-body e meson-exchange currents
- reazioni considerate in dettaglio:

$$p + p \rightarrow d + e^+ + v_e e \quad n + p \rightarrow d + \gamma$$

Campi di applicazione

- Big Bang nucleosynthesis (BBN)
- Produzione solare di energia attraverso una catena di reazioni nucleari

Big Bang nucleosynthesis (BBN) (naive picture)

- Step 1 (kT < few MeV): thermal eq.; *n/p* ~ 1
- Step 2 (kT ~ 0.7 MeV): weak int. freeze-out, $n/p \sim 1/6 \rightarrow 1/7$
- Step 3 (kT ~ 0.5 MeV): nuclear reactions, which form *d*, ³He, ⁴He, ⁷Li

11 key nuclear reactions, among which:

- $n + p \rightarrow d + \gamma$
- $p + d \rightarrow {}^{3}\text{He} + \gamma$
- ${}^{3}\text{H} + \alpha \rightarrow {}^{7}\text{Li} + \gamma$
- ${}^{3}\text{He} + \alpha \rightarrow {}^{7}\text{Be} + \gamma$
- $^{7}\text{Be} + n \rightarrow ^{7}\text{Li} + p$







Solar neutrinos: the pp chain



www.sns.ias.edu/~jnb



The SuperKamiokande (SK) experiment

Elastic scattering $v + e \rightarrow v + e$

Sensitive to all the v, but in fact particularly to v_{e}







The Sudbury Neutrino Observatory (SNO)





Neutrino interactions in D_2O

Charged Current Reaction (CC)



Neutral Current Reaction (NC)



Elastic Scattering (ES)



Total Rates: Standard Model vs. Experiment Bahcall-Pinsonneault 2000



Theoretical calculations: state-of-the-art

- 1) <u>Realistic description of initial and final nuclear bound- and scattering</u> <u>states</u> (*realistic Hamiltonians: two- and three-nucleon interactions*)
- 2) <u>Realistic description of the electro-weak nuclear current</u> (*one-body and meson-exchange currents*)

For both 1) and 2):

- high accuracy
- widely tested especially in the electromagnetic sector (electron scattering)

Marcucci *et al.*, Nucl. Phys. A **777**, 111 (2006) Adelberger *et al.*, Rev. Mod. Phys. **83**, 195 (2011)

 $n + p \rightarrow d + \gamma$

Total cross section [mb] with the AV18

| One-body | 304.6 |
|----------|----------|
| Full | 332.7 |
| Expt. | 332.6(7) |

 $p + p \rightarrow d + e^+ + v_e$

$S(E) = E \sigma(E) \exp(2 \pi Z_1 Z_2 \alpha / v)$ $S(E) \approx \Lambda(E)^2$

| NN model | AV18 | CD-Bonn |
|----------|-------|---------|
| One-body | 6.965 | 6.985 |
| Full | 7.076 | 7.060 |

H.A. Bethe and C.L. Critchfield, Phys. Rev. 54, 248 (1938)

| | $r_0 = e^2/mc^2$ | $	au_0 = e^2/2 m c^2$ | |
|-----------------------|------------------|------------------------|--|
| x_0 . | 0.645 | 0.322 | |
| V_0 (Mev) | 20.9 | 66.5 | |
| D (Mev) | 10.3 | 47.0 | |
| μ., | 2.94 | 5.45 | |
| y | 2.18 | 4.65 | |
| $(rd \log w/dr)$ | 0.236 | 0.110 | |
| $\hat{\Phi}(r_0)$ | 1.050 | 1.025 | |
| $\Theta(r_0)$ | 0.769 | 0.854 | |
| $(rd \log \Phi/dr)$ | 0.050 | 0.025 | |
| Č, Š | 0.814 | 0.915 | |
| λ | 2.63 | 4.80 | |
| Δ1 | 0.689 | 0.277 | |
| Λ_2 | 1.949 | 1.547 | |
| Aa | 1.205 | 1.030 | |
| $(1+x_0)(1+\mu^{-2})$ | 1.835 | 1.367 | |
| Λ | 2.84 | 2.44 | |
| Λ ² | 8.08 | 5.93 | |

TABLE I. Numerical results for two values of the radius.

¹³ Breit, Condon and Present, Phys. Rev. 50, 825 (1936).

 $n + d \rightarrow {}^{3}\mathrm{H} + \gamma$

Total cross section [mb] (AV18/UIX)

| One-Body | 0.227 |
|----------|-----------|
| Full | 0.556 |
| Expt. | 0.508(15) |

$$p + d \rightarrow {}^{3}\text{He} + \gamma$$



 $p+^{3}\text{He} \longrightarrow ^{4}\text{He}+e^{+}+v_{e}$: the hep reaction

• $S_{\rm SSM98} = 2.3 \times 10^{-20} \text{ keV b}$

 New calculation (Marcucci et al., PRL 84, 5959 (2000); PRC 63, 015801 (2001)):

1. S (10 keV) = 10.1×10^{-20} keV b

- 2. no c.m. energy dependence
- no Hamiltonian model dependence (if H reproduces accurately the initial and final state w.f.'s)
- 4. importance of the P-waves (40%)
- 5. importance of MEC (Δ)

• SSM revisited (Bahcall *et al.*, Astr. J. 555, 990 (2001)): $S_{\text{SSM00}} = 10.1 \times 10^{-20} \text{ keV b}$





The SK collaboration, PRL 86, 5651 (2001)



Limits

- No clear connection with $QCD \rightarrow model-dependence$
- No clear determination of theoretical uncertainty



Chiral Effective Field Theory (*x***EFT)**

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- $QCD \rightarrow quark and gluons$ ("heavy" degrees of freedom)
- Nuclear physics → nucleons and pions ("light" degrees of freedom)
- EFT \rightarrow processes with $E \sim p \sim m_{\pi} \ll \Lambda_{\text{OCD}} \sim 1 \text{ GeV}$

► "heavy" d.o.f. integrated out \rightarrow contact interactions with "light" d.o.f. and low-energy constants (LECs) obtained from experiment

► perturbative theory: matrix elements $\propto O(p/\Lambda_{OCD})^{\vee}$

 $\chi EFT \rightarrow$ implement EFT and spontaneous breaking of QCD's chiral symmetry

Advantages: a) "right" treatment of πN interaction b) nuclear force "hierarchy" \rightarrow accurate $V_{NN} + V_{NNN}$

Disadvantage: limited to processes occurring at low-energy $E \sim 1-2 \text{ m}_{\pi}$

First steps (~2001): "hybrid" χ EFT

- Nuclear wave functions from realistic phenomenological potentials (AV18, etc)
- Nuclear electroweak current from χEFT

Study of the *pp* and *hep* reactions in agreement with "old" phenomenological ones T.S. Park *et al.*, Phys. Rev. C **67**, 055206 (2003)

On-going work toward a fully consistent χ EFT calculation





Girlanda et al., Phys. Rev. Lett. 105, 232502 (2010)

Fully consistent χEFT calculation

- Nuclear interactions and currents in χEFT at N3LO (three-nucleon interaction at N2LO)
- First χ EFT calculation for $\mu^- + d \rightarrow n + n + \nu_{\mu}$ and $\mu^- + {}^{3}\text{He} \rightarrow {}^{3}\text{H} + \nu_{\mu}$
- On-going work for $p + p \rightarrow d + e^+ + v_e$
- Future work for

$$-p + {}^{3}\text{He} \rightarrow {}^{4}\text{He} + e^{+} + v_{e}$$
$$-p + d \rightarrow {}^{3}\text{He} + \gamma$$

Results (with RC)

| | ${}^{1}S_{0}$ | ${}^{3}P_{0}$ | ${}^{3}P_{1}$ | ${}^{3}P_{2}$ | Γ^D | Γ_0 |
|--|---------------|---------------|---------------|---------------|------------|------------|
| $IA-\Lambda=500~{\rm MeV}$ | 238.8 | 21.1 | 44.0 | 72.4 | 381.7 | 1362 |
| $IA-\Lambda=600~{\rm MeV}$ | 238.7 | 20.9 | 43.8 | 72.0 | 380.8 | 1360 |
| $FULL-\Lambda=500~{\rm MeV}$ | 254.4(9) | 20.5 | 46.8 | 72.1 | 399.2(9) | 1488(9) |
| $FULL-\Lambda=600~{\rm MeV}$ | 255(1) | 20.3 | 46.6 | 71.6 | 399(1) | 1499(9) |
| $\chi EFT^* - \Lambda = 500 \text{ MeV}$ | 251.5(7) | 20.0 | 46.6 | 71.8 | 395.1(7) | 1489(9) |

$\Gamma^{D} = 399(3) \text{ s}^{-1} \& \Gamma_{0} = 1494(21) \text{ s}^{-1}$

Marcucci et al., PRL 108, 052502 (2012)

 $\Gamma_0(\exp)=1496(4) \text{ s}^{-1}$



